

Probing the Three Dimensional Nature of Supernovae; Results from the Supernova Spectropolarimetry Project (SNSPOL)*

*Supernovae Are Not Spherical Cows

NSF AST-1210599

G. Grant Williams
MMT Observatory/U. Arizona
June 3, 2015

\neq

SN 1987A

NASA/ESA/Challis & Kirshner





C. Bilinski



N. Smith



P. Smith



P. Milne



Leah Huk
U. Denver



J. Hoffman
U. Denver



D. Leonard
SDSU



L. Dessart
Lagrange, CNRS



+ many more

Eta Carinae SN Progenitor

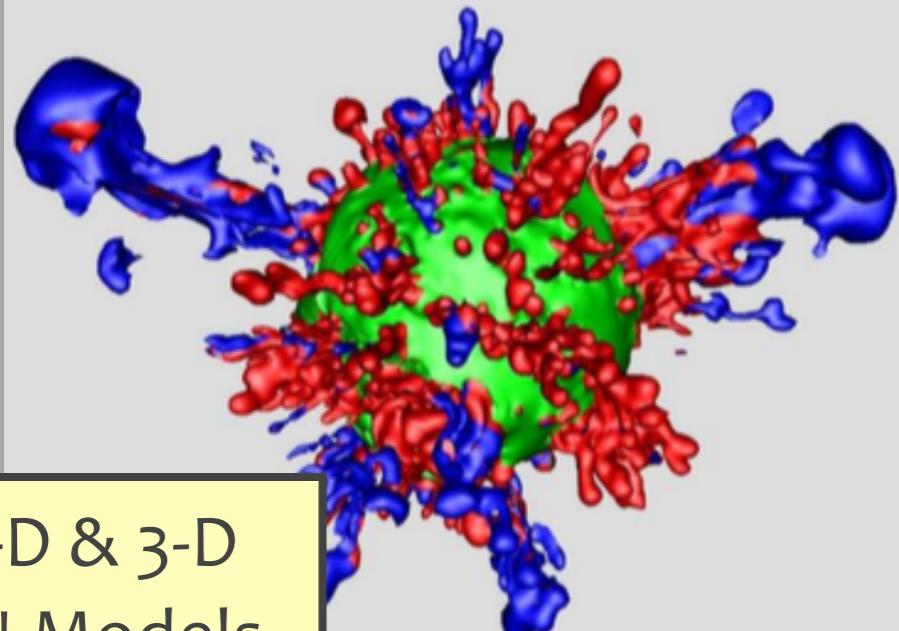


Eta Carinae

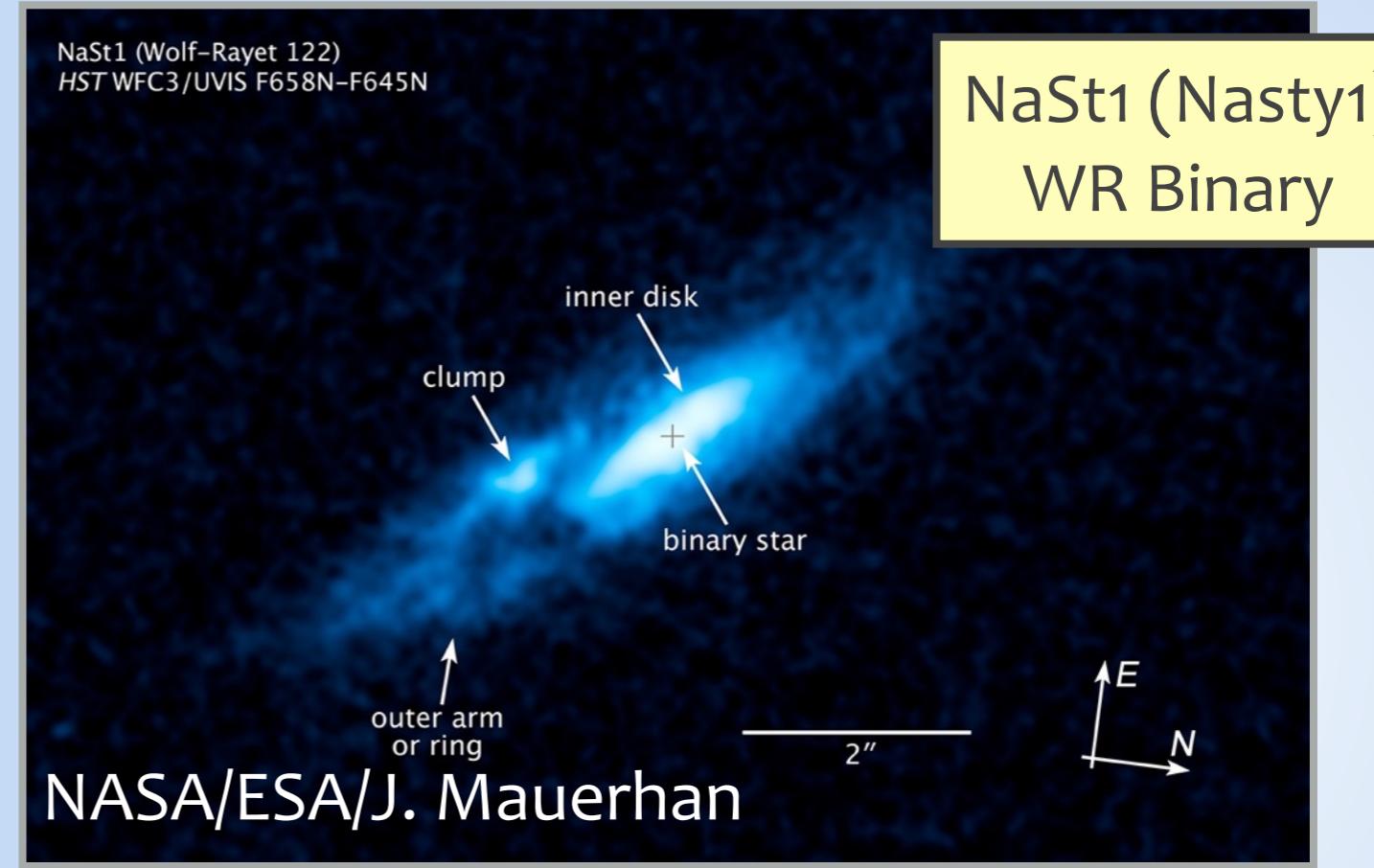
PRC96-23a · ST Scl OPO · June 10, 1996
J. Morse (U. CO), K. Davidson, (U. MN), NASA

HST · WFPC2

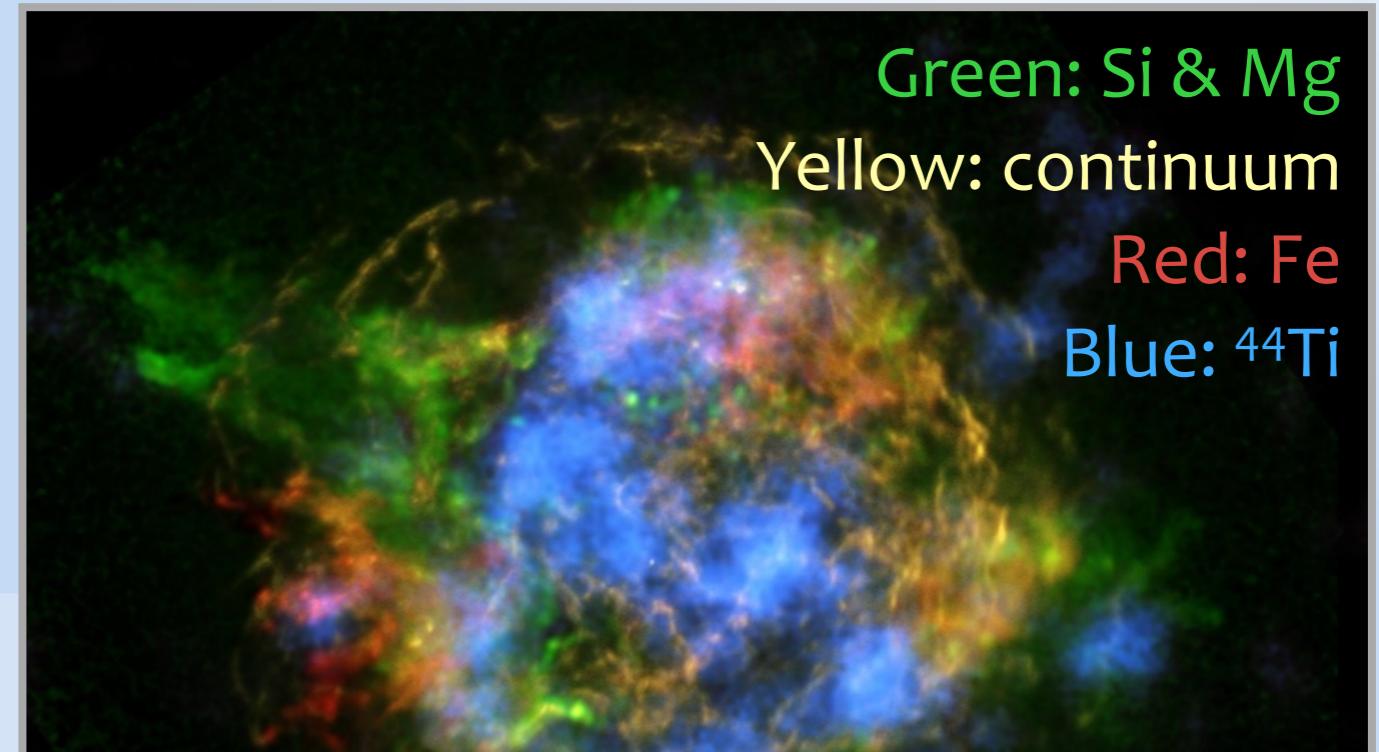
Hammer et al. 2010



2-D & 3-D
SN Models



NaSt1 (Nasty1) WR Binary



NASA/JPL-Caltech/CXC/SAO
Grefenstette et al.

Cas A SN Remnant



Sep, 1994

Mar, 1995

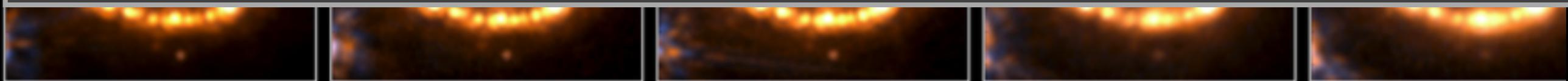
Feb, 1996

Jul, 1997

Feb, 1998

Polarimetry allows us to probe scales in a supernova that cannot be imaged from earth or space.

Spectropolarimetry enhances the power of this technique by allowing us to probe the geometry of specific ionic species.



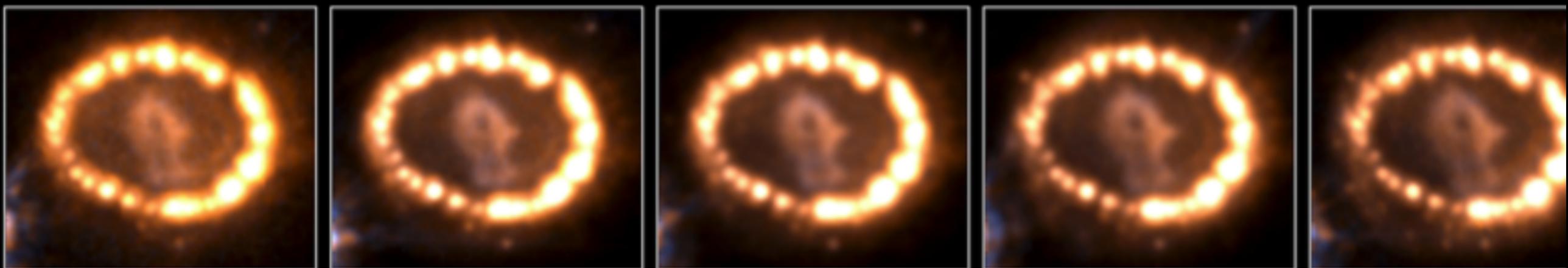
Sep, 2005

Apr, 2006

Dec, 2006

May, 2007

Feb, 2008



Apr, 2009

Dec, 2009

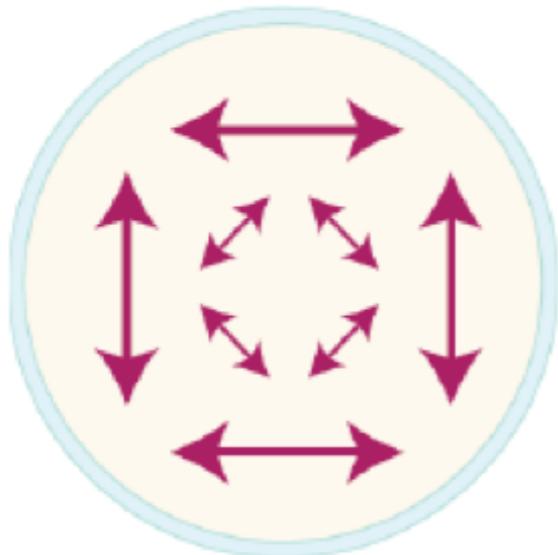
Jan, 2011

Feb, 2013

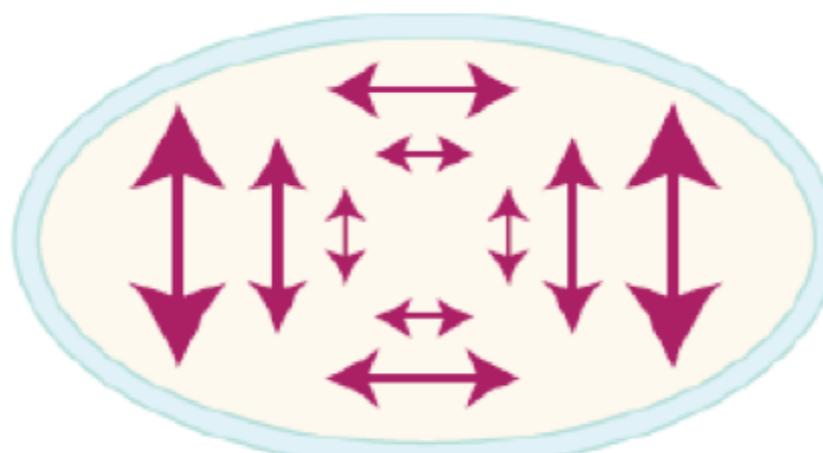
Jun, 2014

Electron Scattering → Polarizing Mechanism

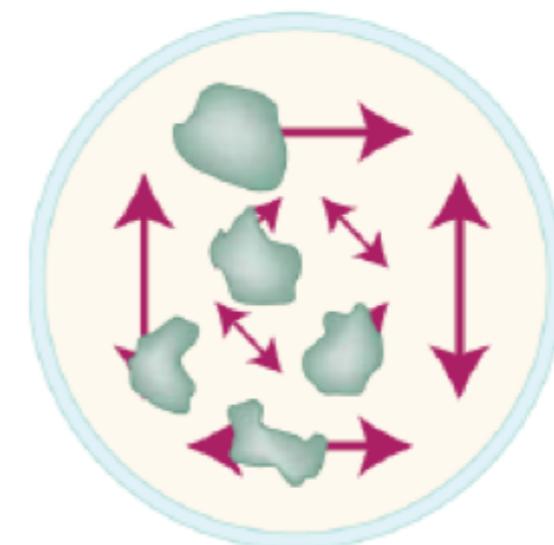
Spherical $P_{\text{net}} = 0\%$



Aspherical $P_{\text{net}} > 0\%$



Clumpy Ejecta



(Image: Leonard 2007, Science, 315,



= Direction of electric vector in plane of sky

Leonard 2007

Wang & Wheeler ARAA 2008

of published SN polarization observations = 94

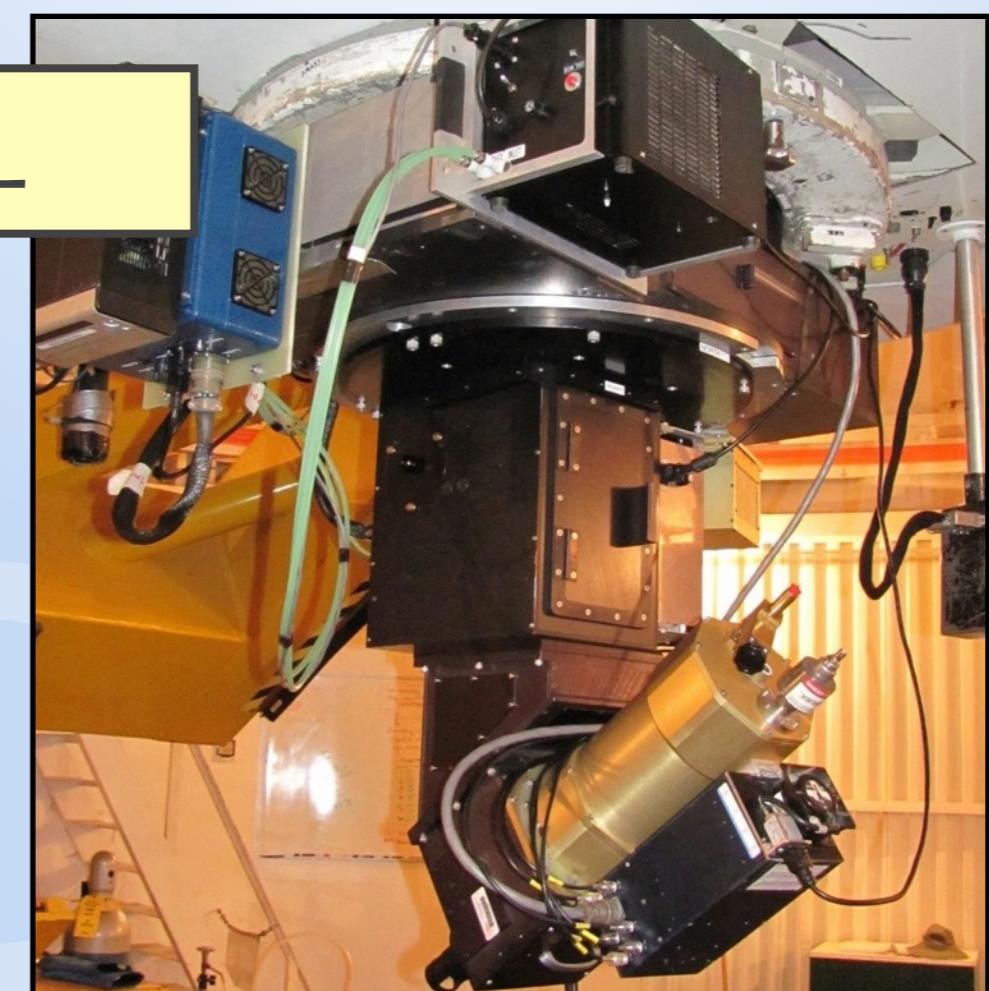
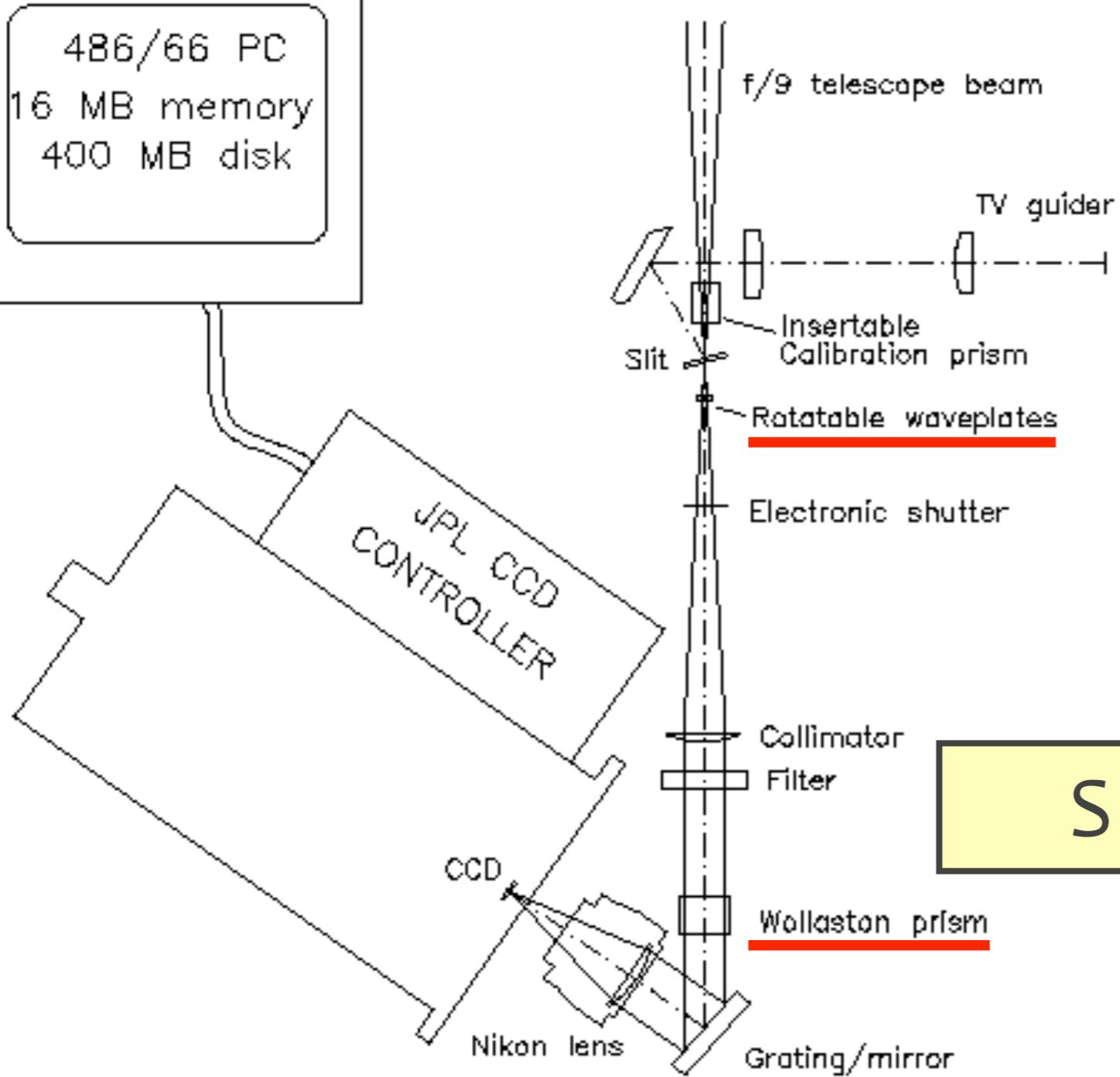
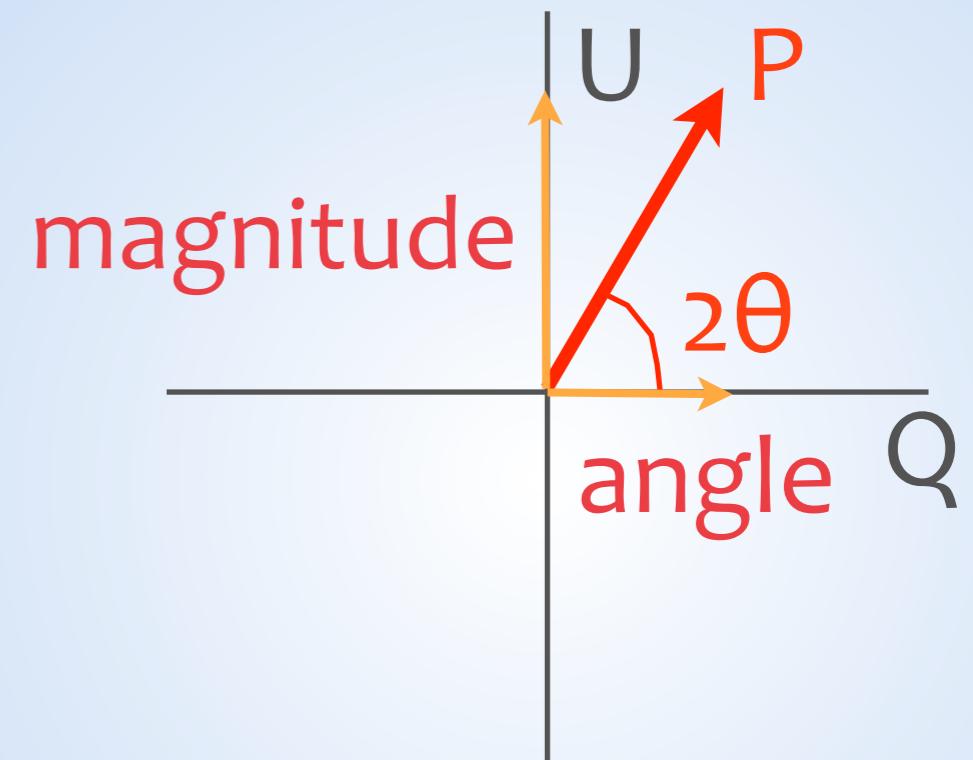
one epoch = 55*

two epochs = 19

three or more = 20

*Interstellar Polarization (ISP) vs. Intrinsic SN Polarization

Linear Stokes Parameters



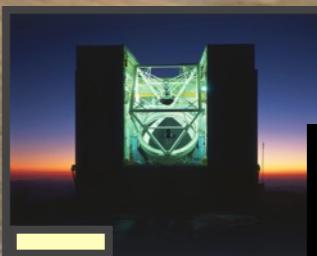
Linear Polarization

$$P = (Q^2 + U^2)^{1/2}$$

$$\theta = 1/2 \arctan (U/Q)$$



Tucson

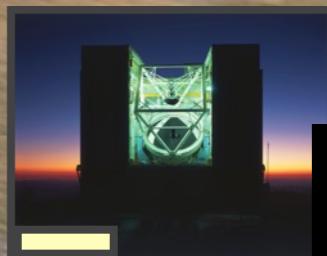


61" Kuiper



Image Credit: T. & J. Polakis

Tucson



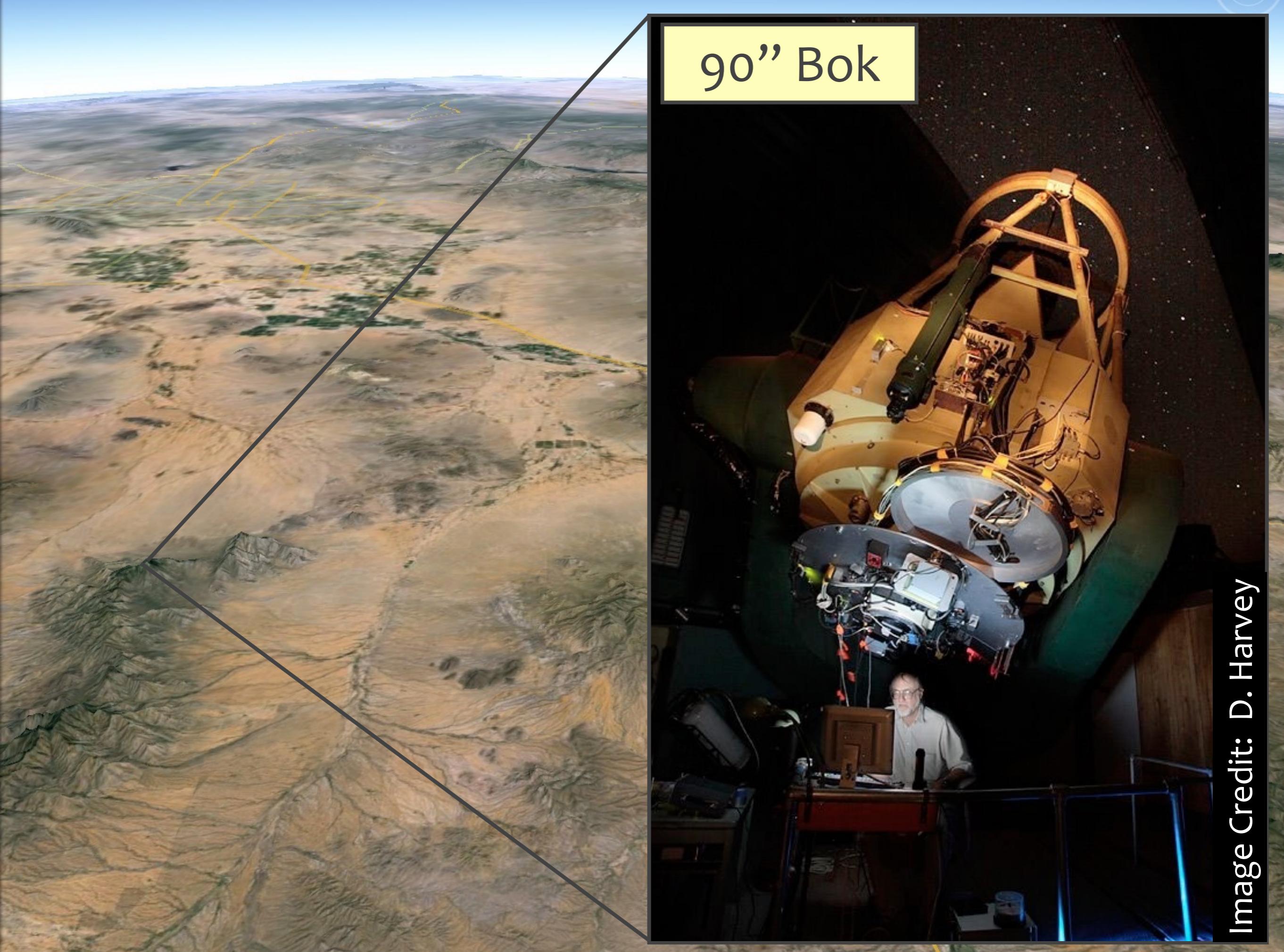


Image Credit: D. Harvey

6.5-m MMT

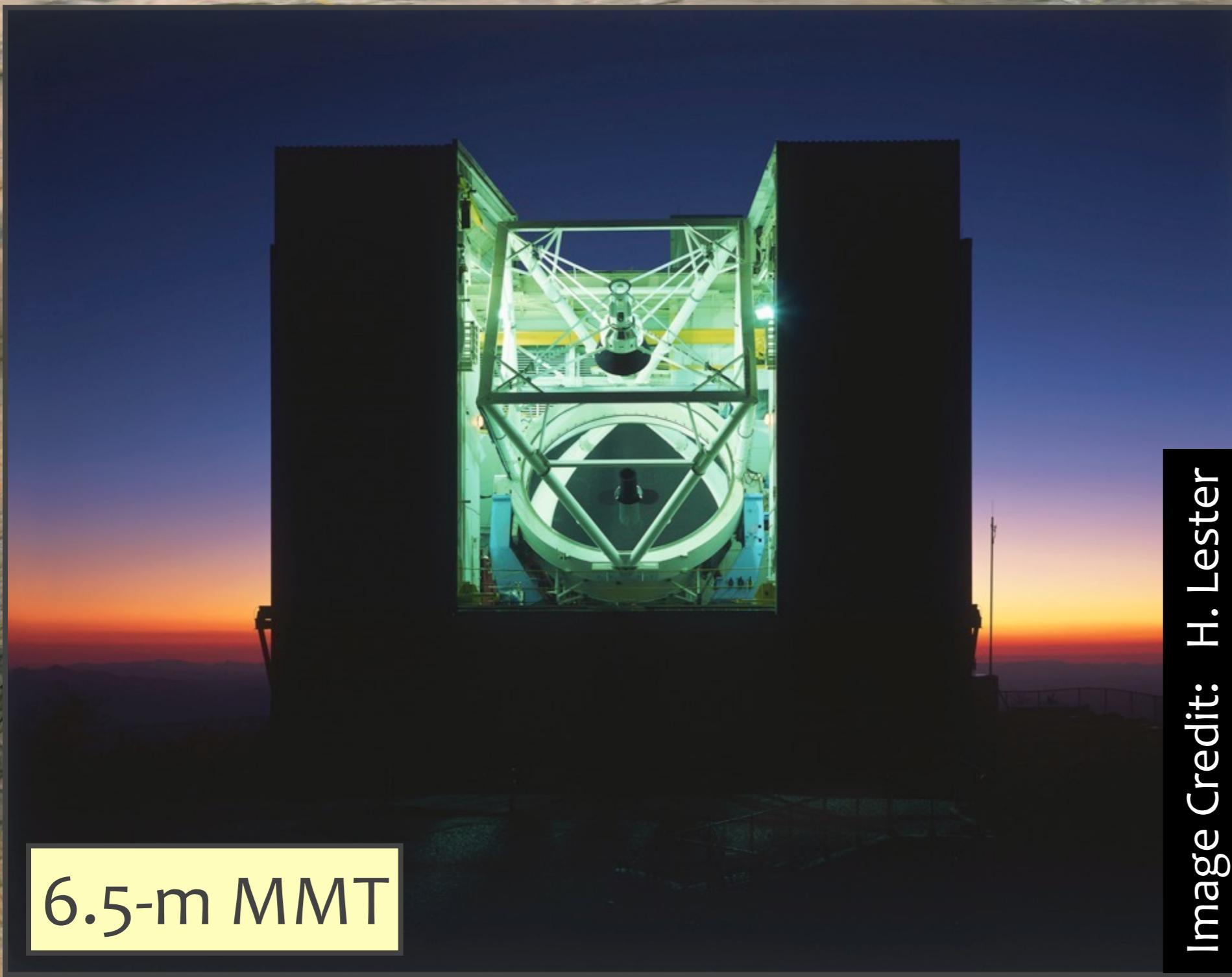
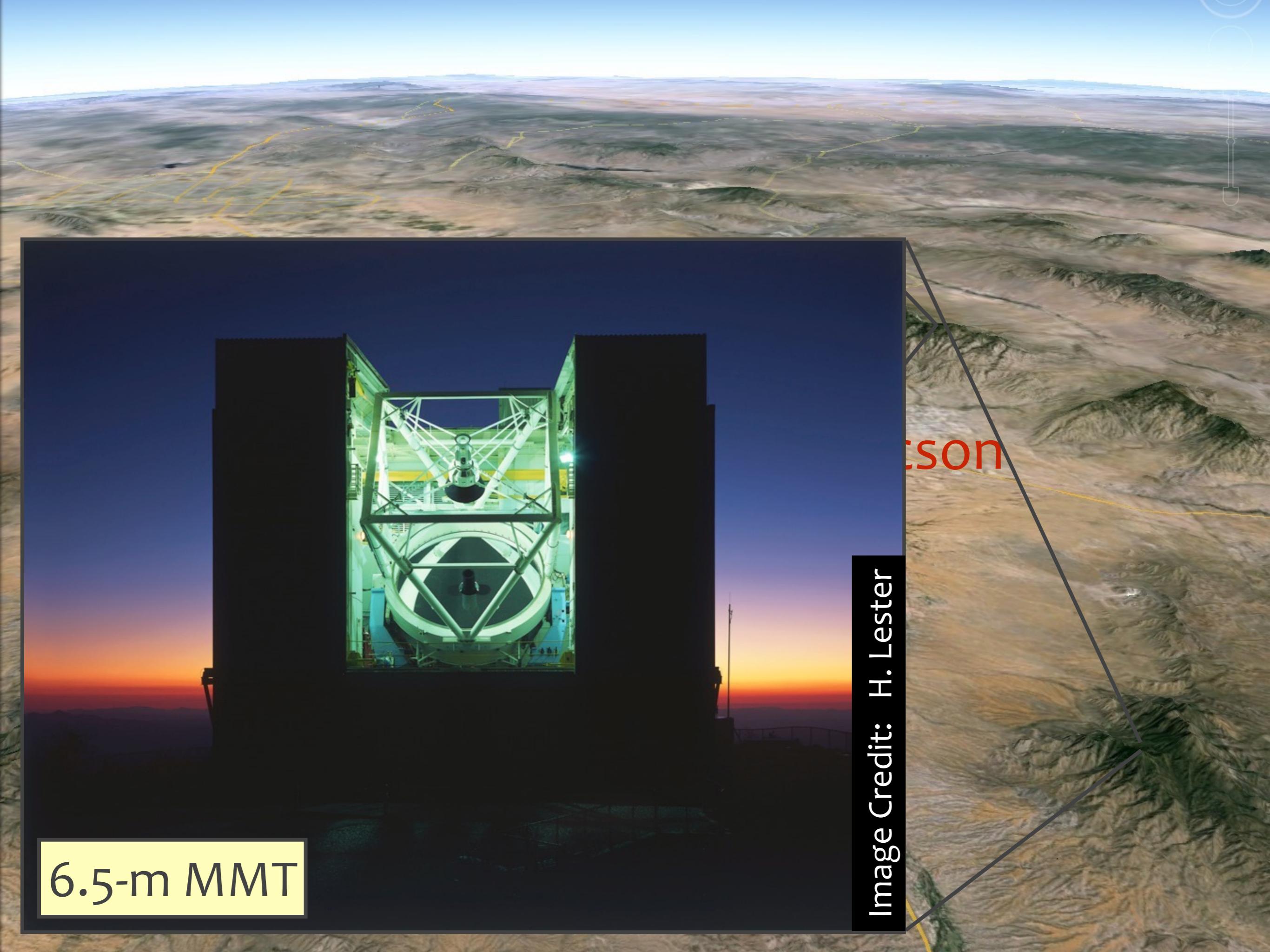


Image Credit: H. Lester

Linson



The Supernova Spectropolarimetry Project (SNSPOL)

GOAL:

To complete a long-term comprehensive **multi-epoch** spectropolarimetry survey of all types of supernovae.

OBJECTIVE:

To improve our understanding of the predominance and characteristics of asymmetries in the different types of supernovae.

Queue observing as part of a larger specpol program guarantees that SN data **every month**.

59 SNe, 44 multiple-epoch

Ia (6/8)

- SN 2011fe (8)
- SN 2012cg (2)
- SN 2012ht (2)
- SN 2013fw (hv, 1)
- SN 2014J (5)
- SN 2014ao (2)
- ASASSN-14lp (6)
- ASASSN-15hy (1)

Ib (2/3)

- SN 2011au (6)
- iPTF13bvn (1)
- M12045 (7)

Ic (7/8)

- PTF12gzk (pec, 2)
- SN 2012ej (4)
- SN 2012fh (6)
- SN 2013ff (3)
- SN 2013ge (2)
- SN 2014L (4)
- SN 2014ad (4)
- SN 2014as (BL,1)

II (5/7)

- SN 2011bv (1)
- SN 2013am (4)
- SN 2013bu (3)
- SN 2013ee (2)
- SN 2013hj (2)
- SN 2014ce (1)
- ASASSN-14kg (2)

IIP (10/15)

- SN 2012A (5)
- SN 2012aw (7)
- SN 2012ch (1)
- SN 2012ec (8)
- SN 2012ho (1)
- SN 2013ab (5)
- SN 2013bi (4)
- SN 2013ej (5)**
- SN 2013fs (4)
- SN 2014A (2)
- SN 2014bc (1)
- SN 2014cx (6)
- SN 2014cy (2)
- J17492 (1)
- J10491 (1)

IIL (1/1)

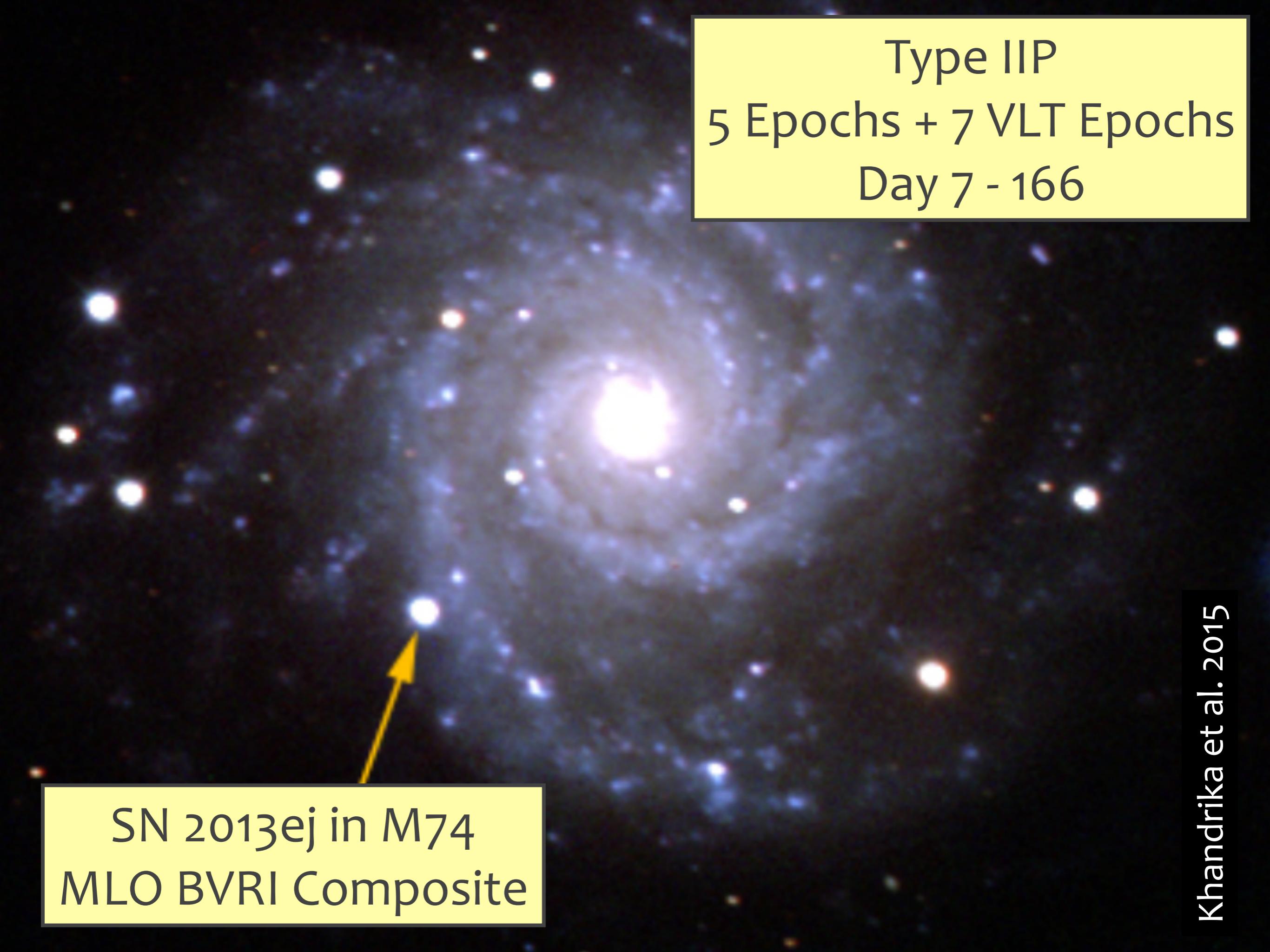
- SN 2014G (4)

IIb (5/5)

- SN 2011dh (2)
- SN 2012fg (3)
- SN 2013ak (2)
- SN 2013 df (2)
- ASASSN-14az (4)

IIn (8/12)

- SN 2010jl (11)**
- SN 2011cc (1)
- SN 2011ht (2)
- SN 2012ab (2)
- PTF11iqb (1)
- SN 2009ip (4)**
- SN 2014ab (5)
- ASASSN-14il (3)
- J07285 (2)
- M04421 (1)
- J13522 (3)
- J090934 (1)

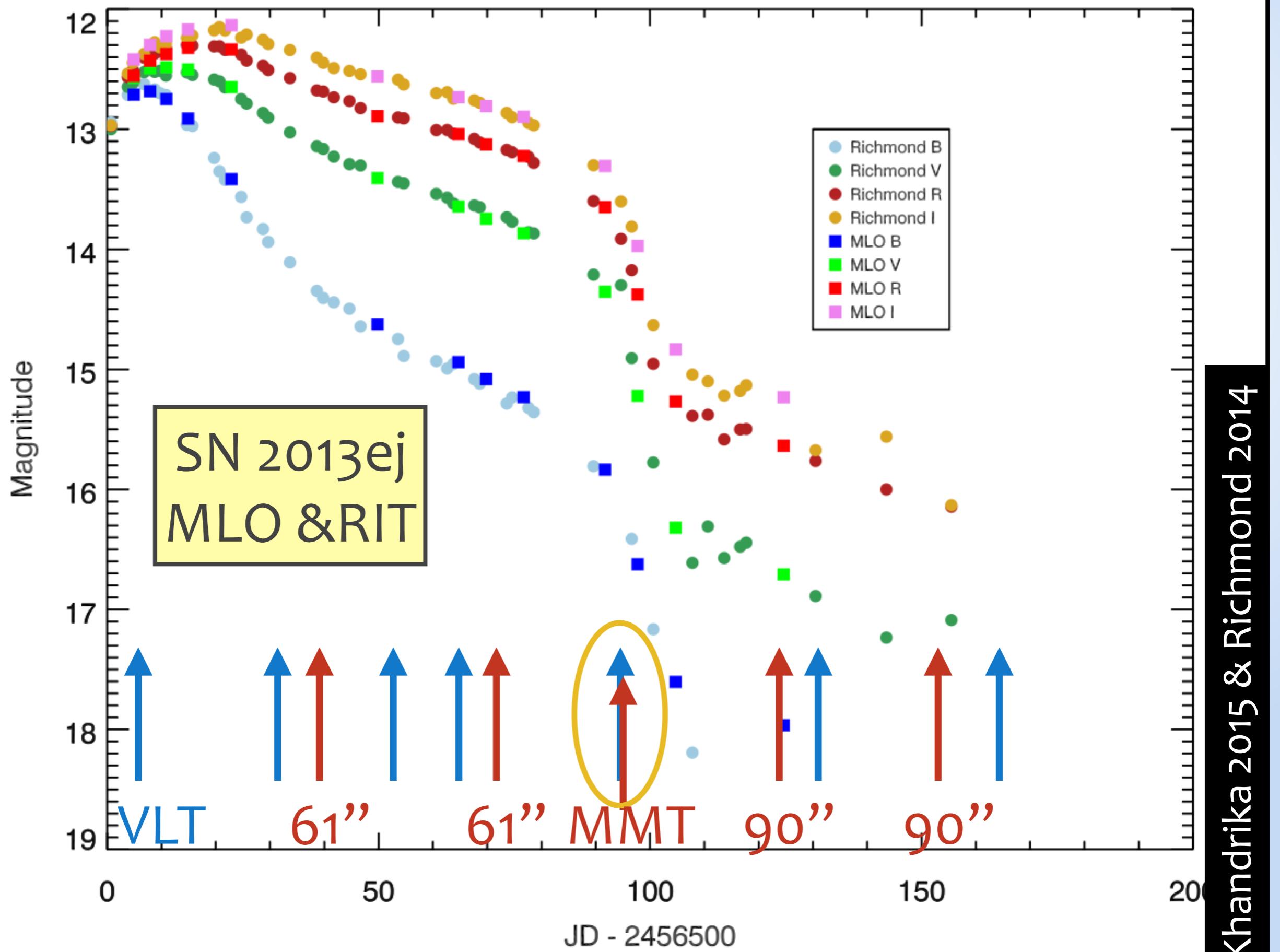


Type IIP
5 Epochs + 7 VLT Epochs
Day 7 - 166

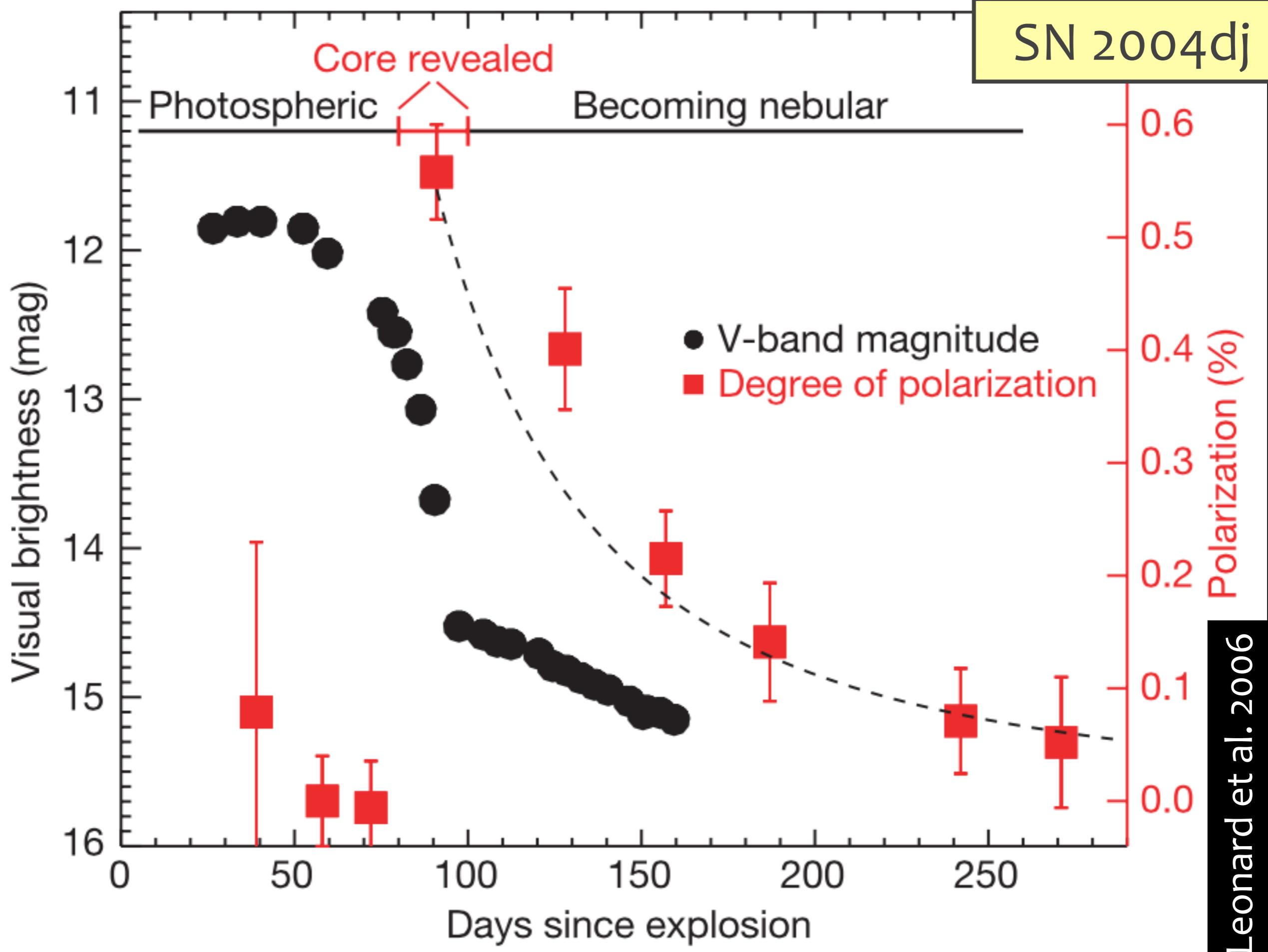
SN 2013ej in M74
MLO BVRI Composite

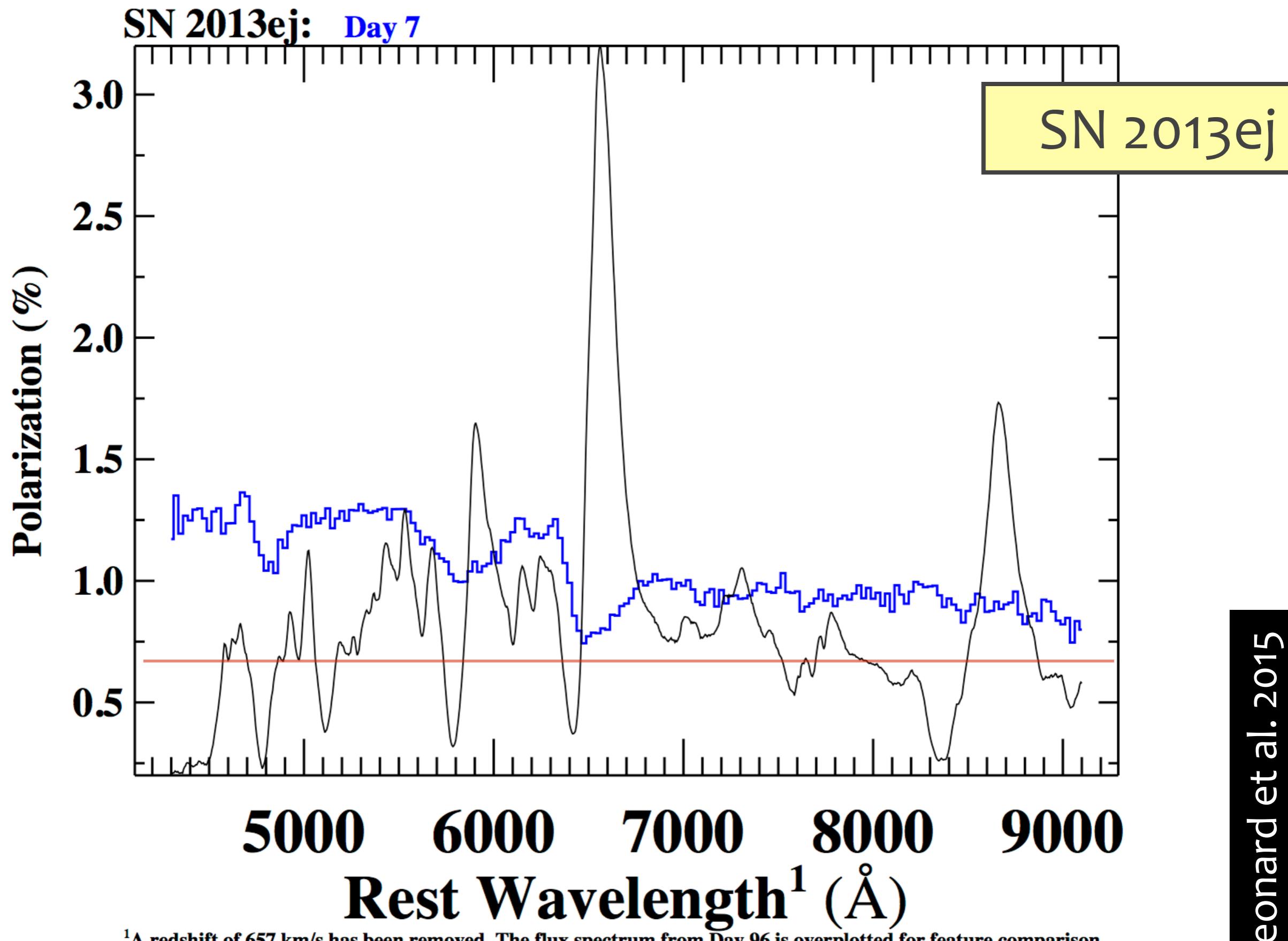
Khandrika et al. 2015

SN2013ej - Richmond (2014) vs MLO

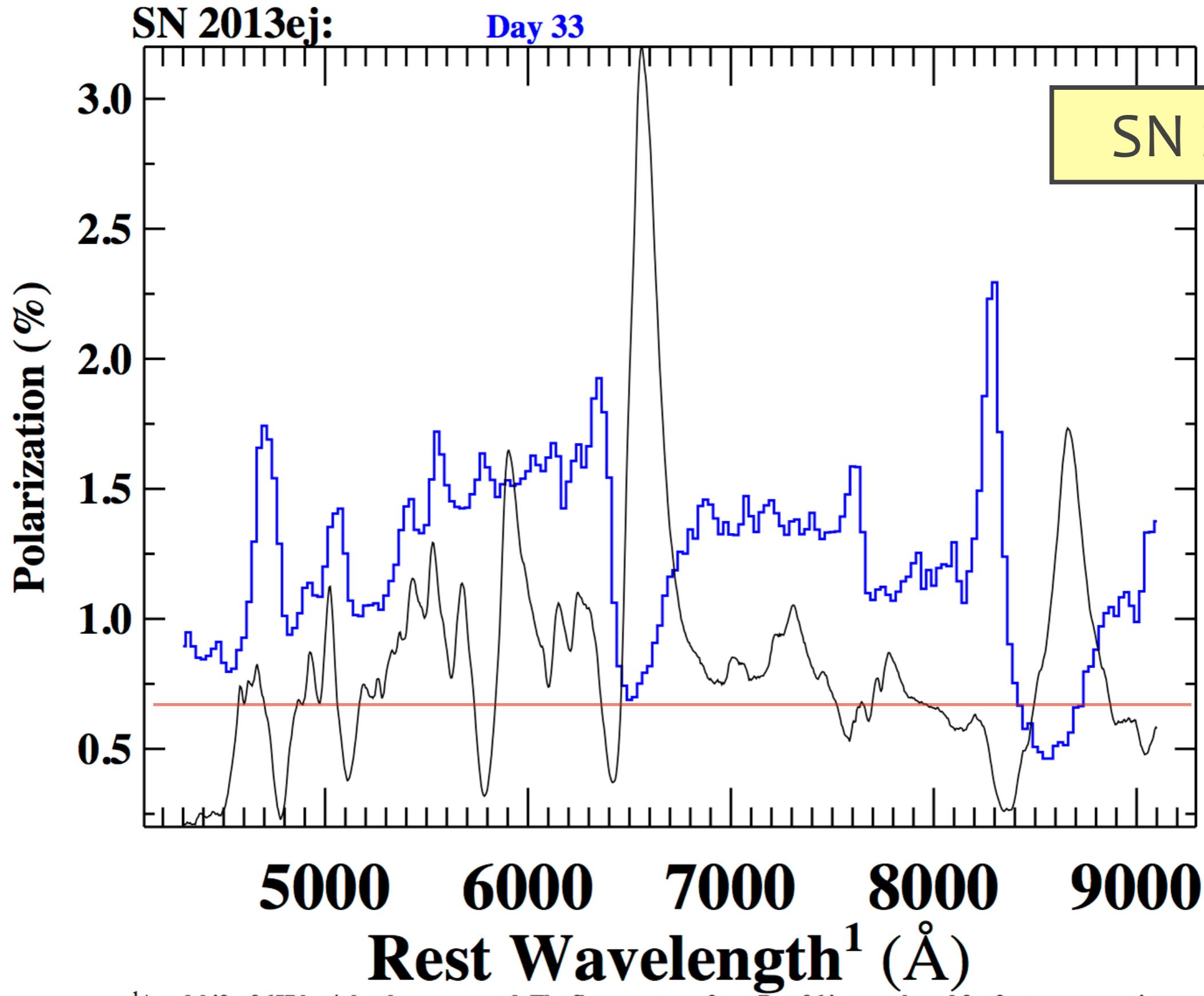


SN 2004dj

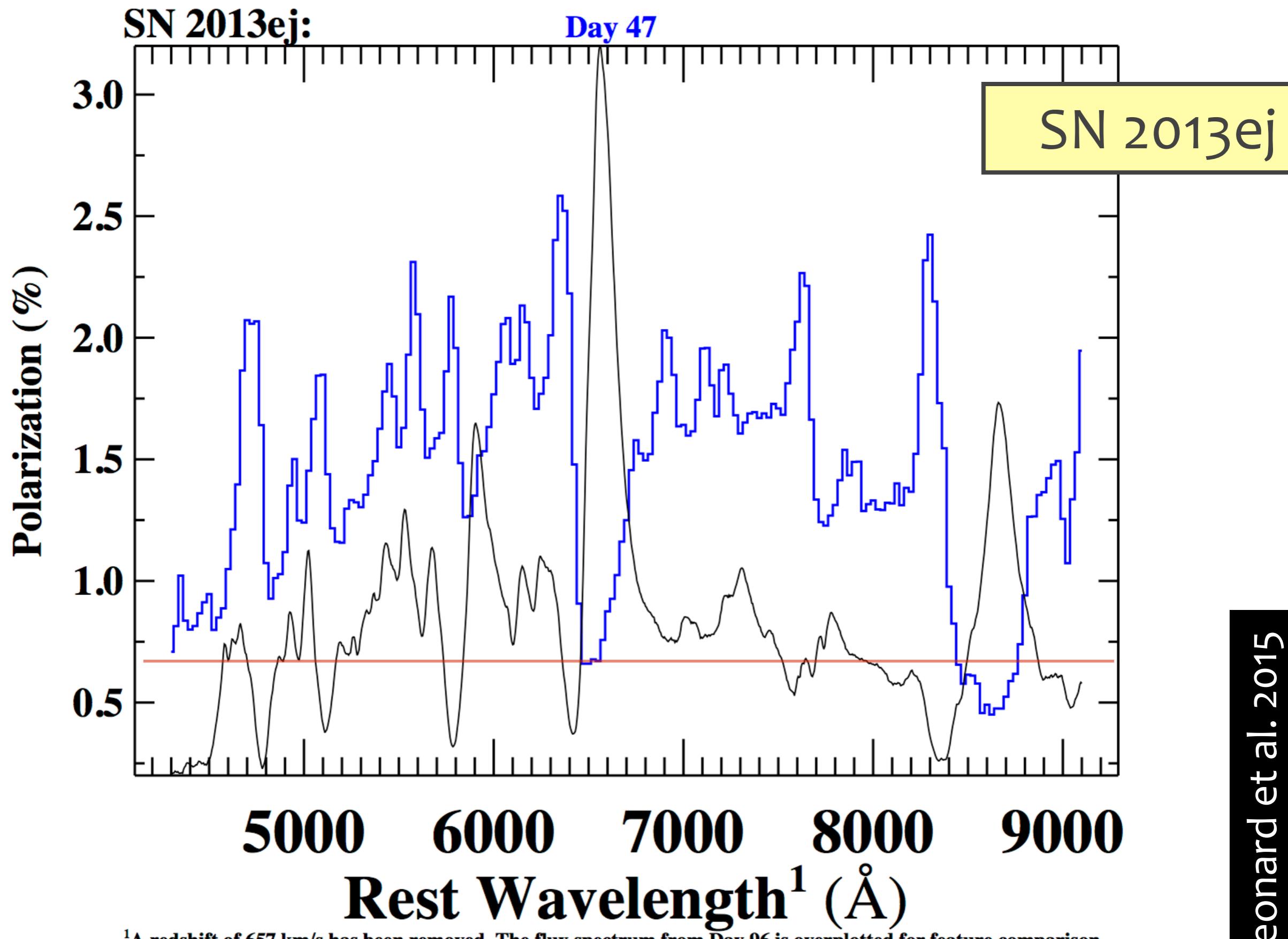




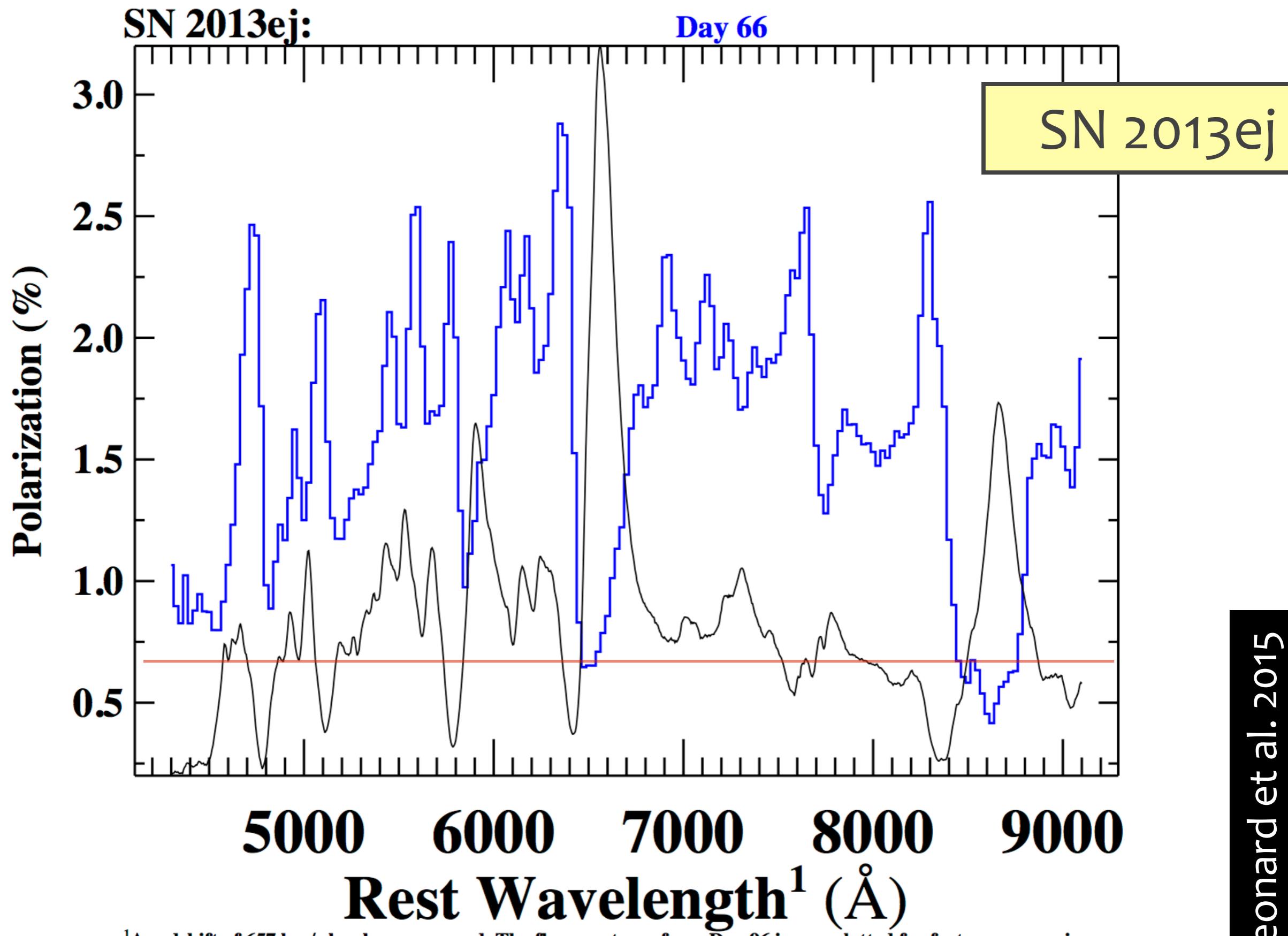
Leonard et al. 2015



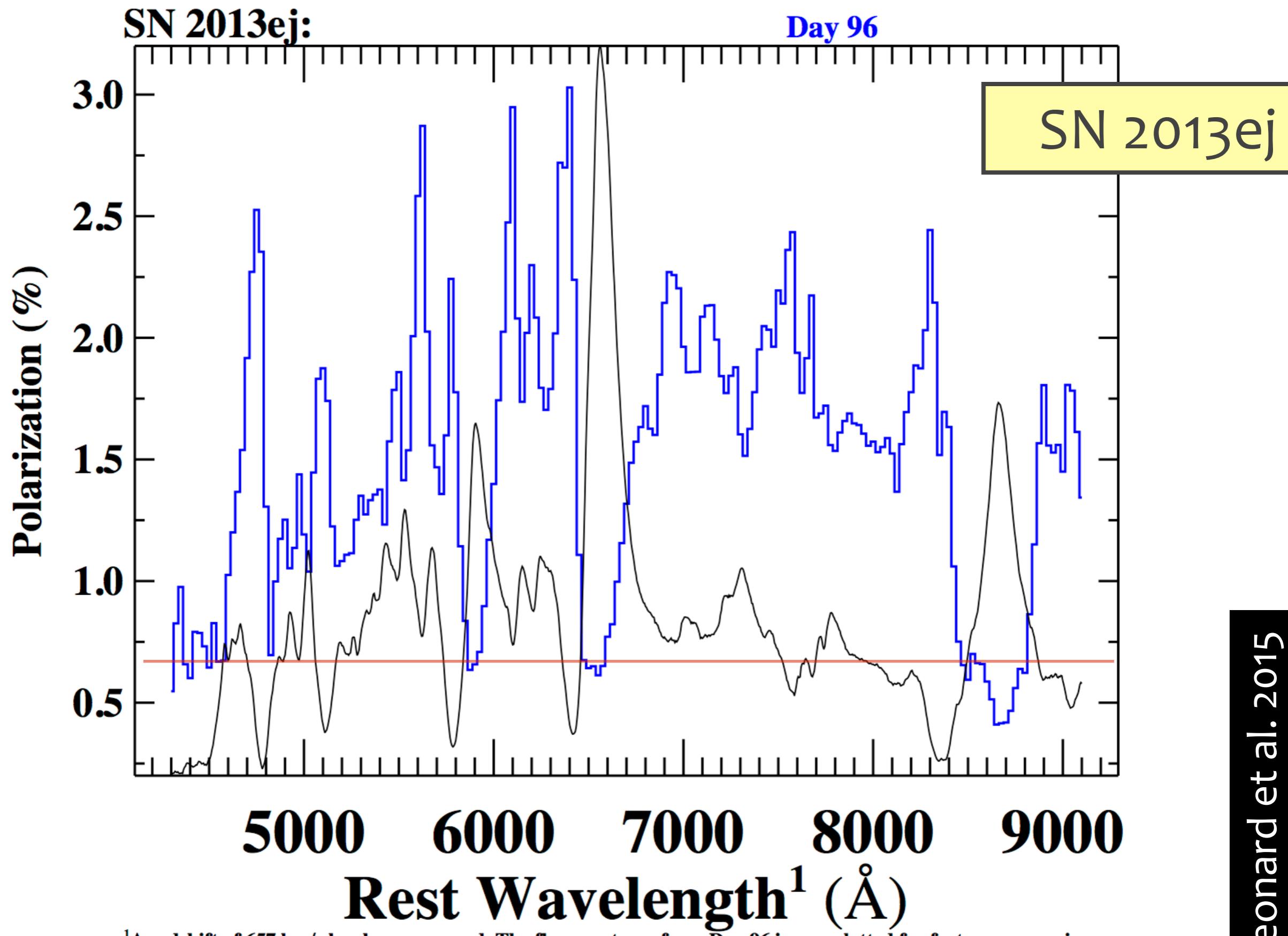
¹A redshift of 657 km/s has been removed. The flux spectrum from Day 96 is overplotted for feature comparison.



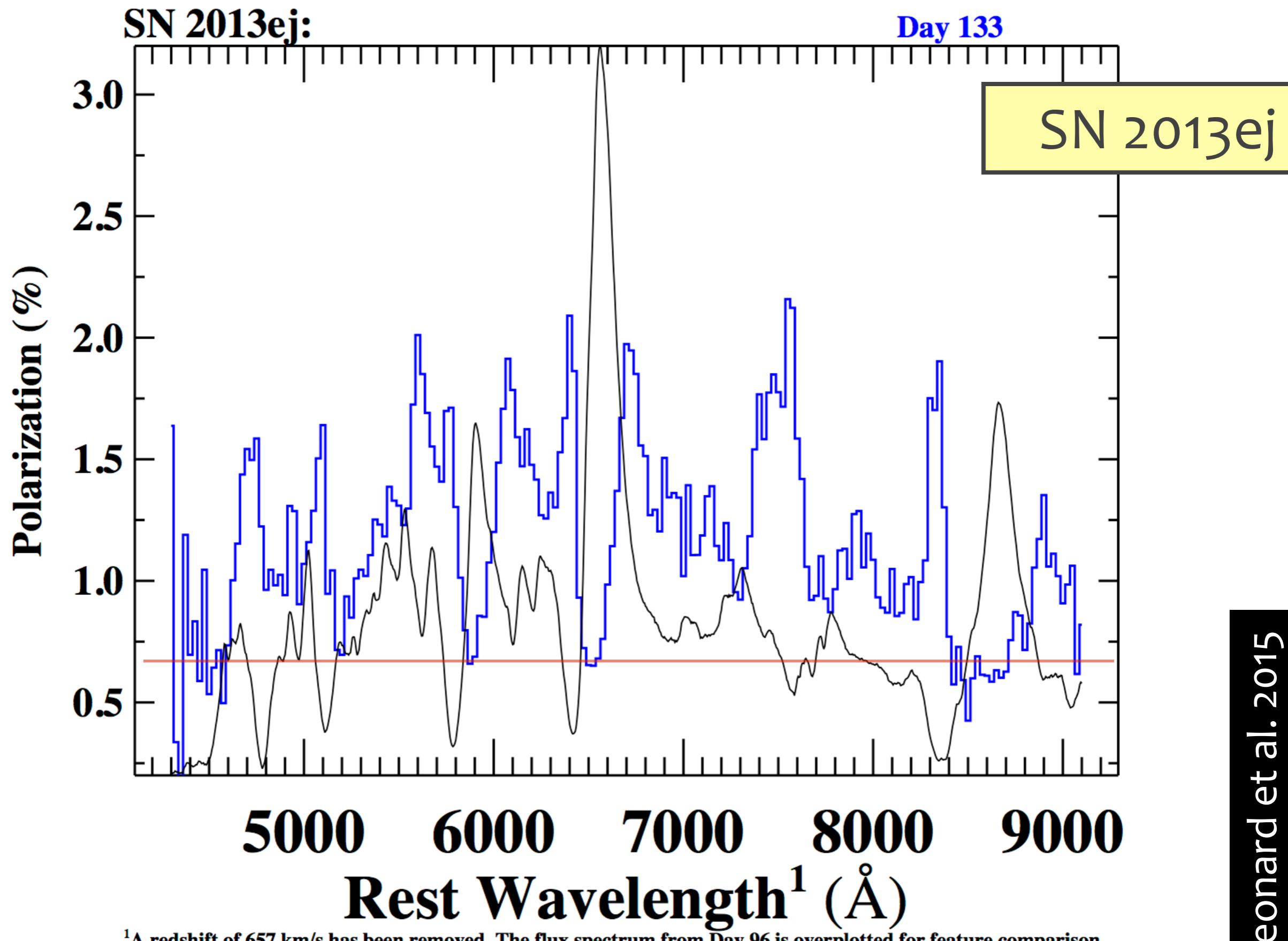
Leonard et al. 2015



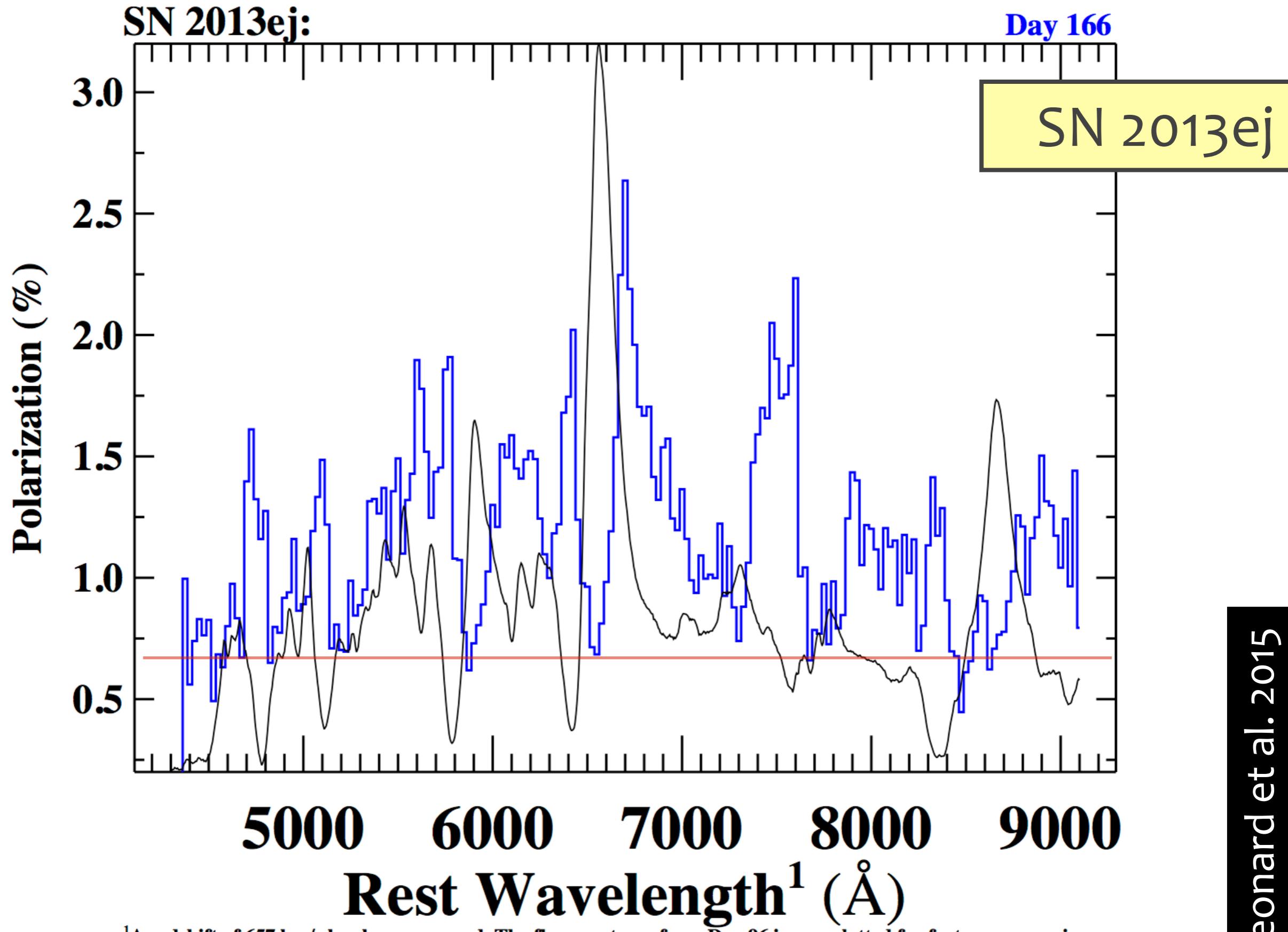
Leonard et al. 2015



¹A redshift of 657 km/s has been removed. The flux spectrum from Day 96 is overplotted for feature comparison.

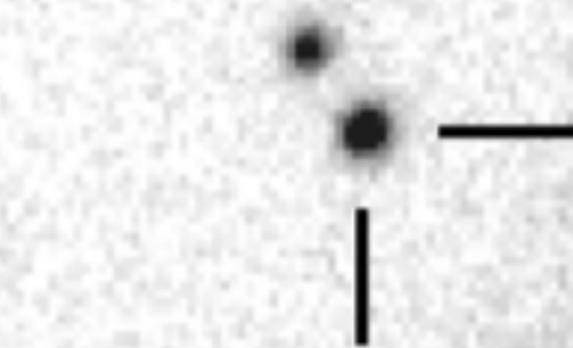


Leonard et al. 2015



Leonard et al. 2015

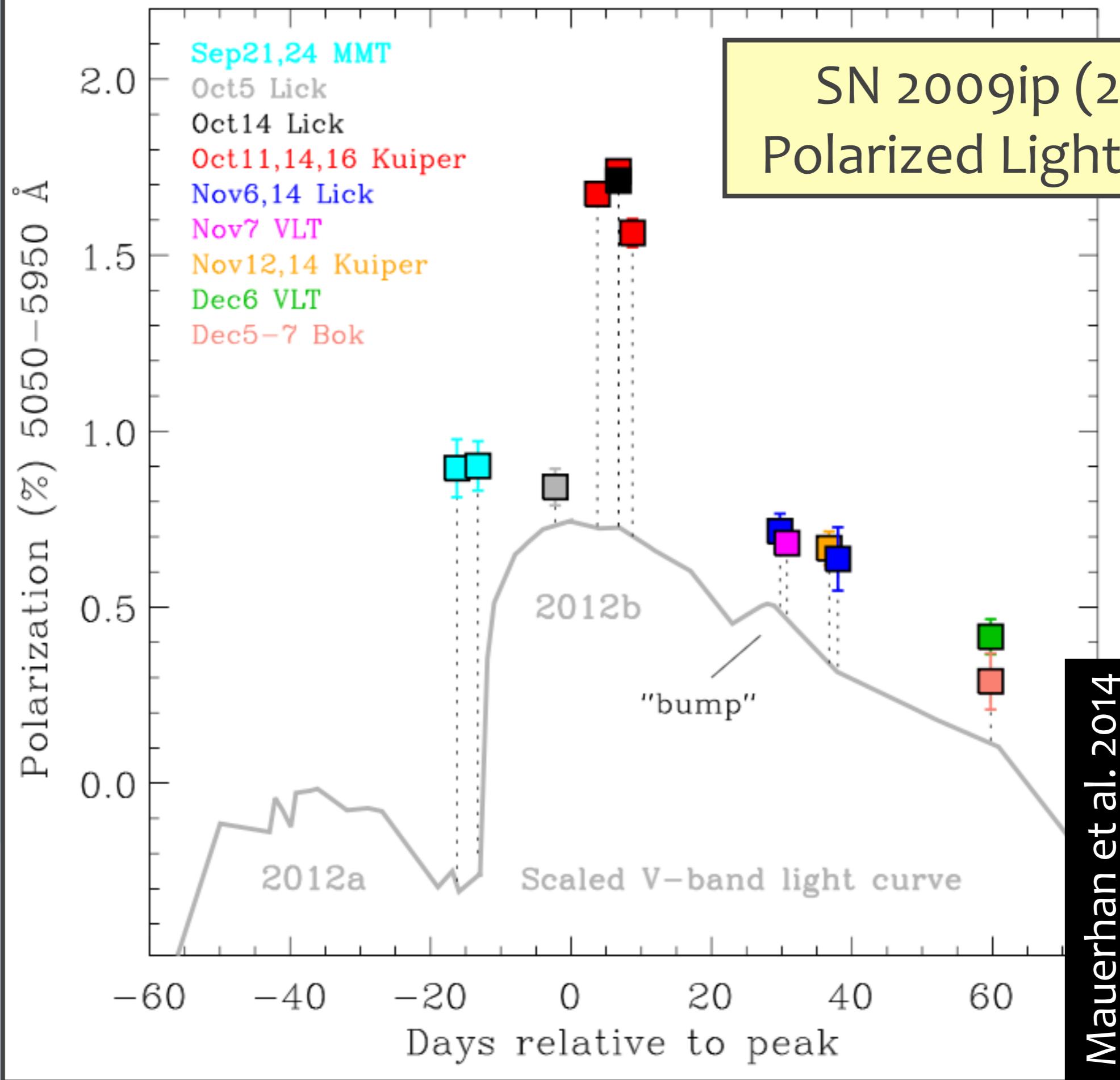
SN 2009ip
Type IIn
4 Epochs + 5 Lick/VLT

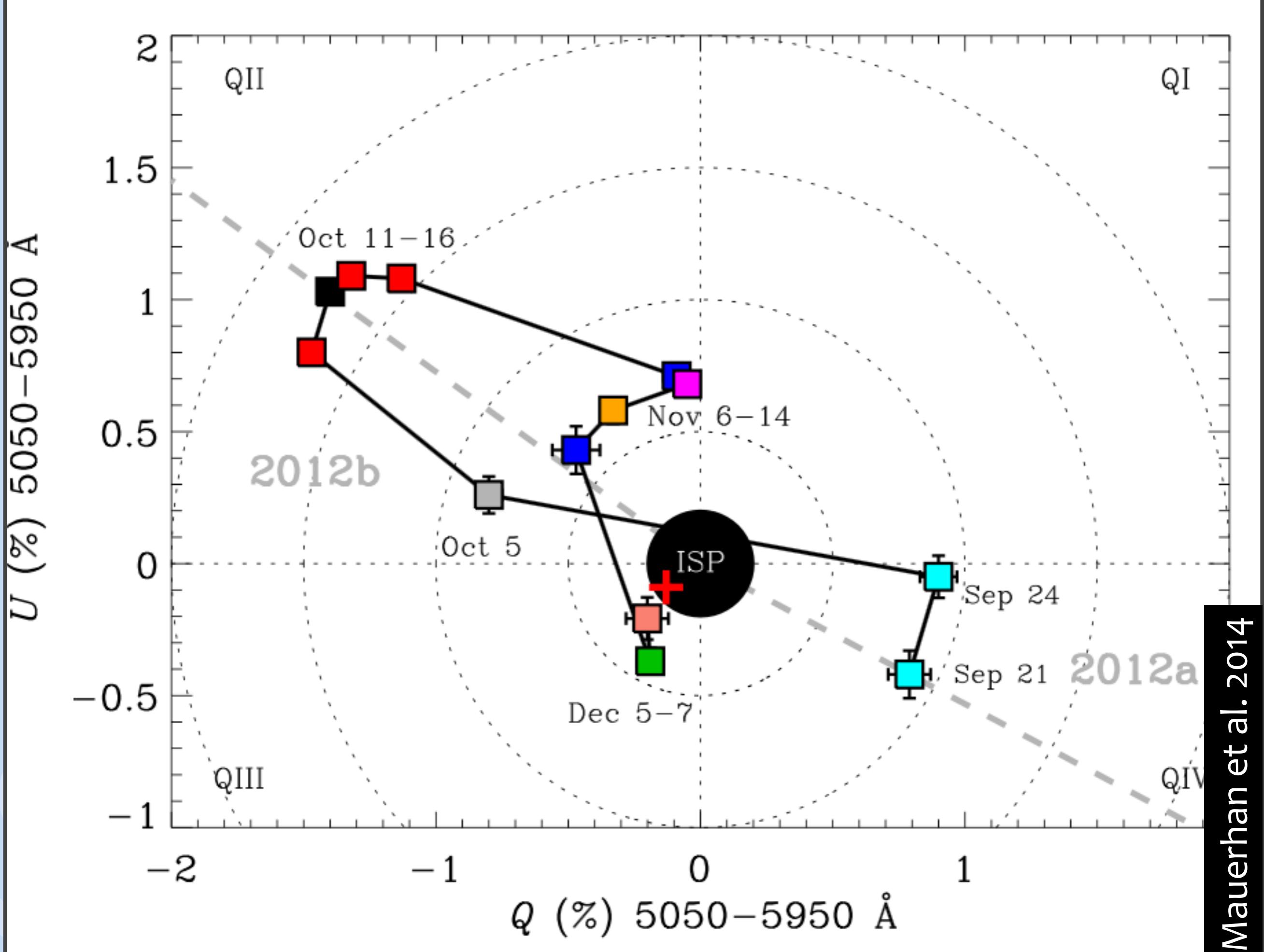


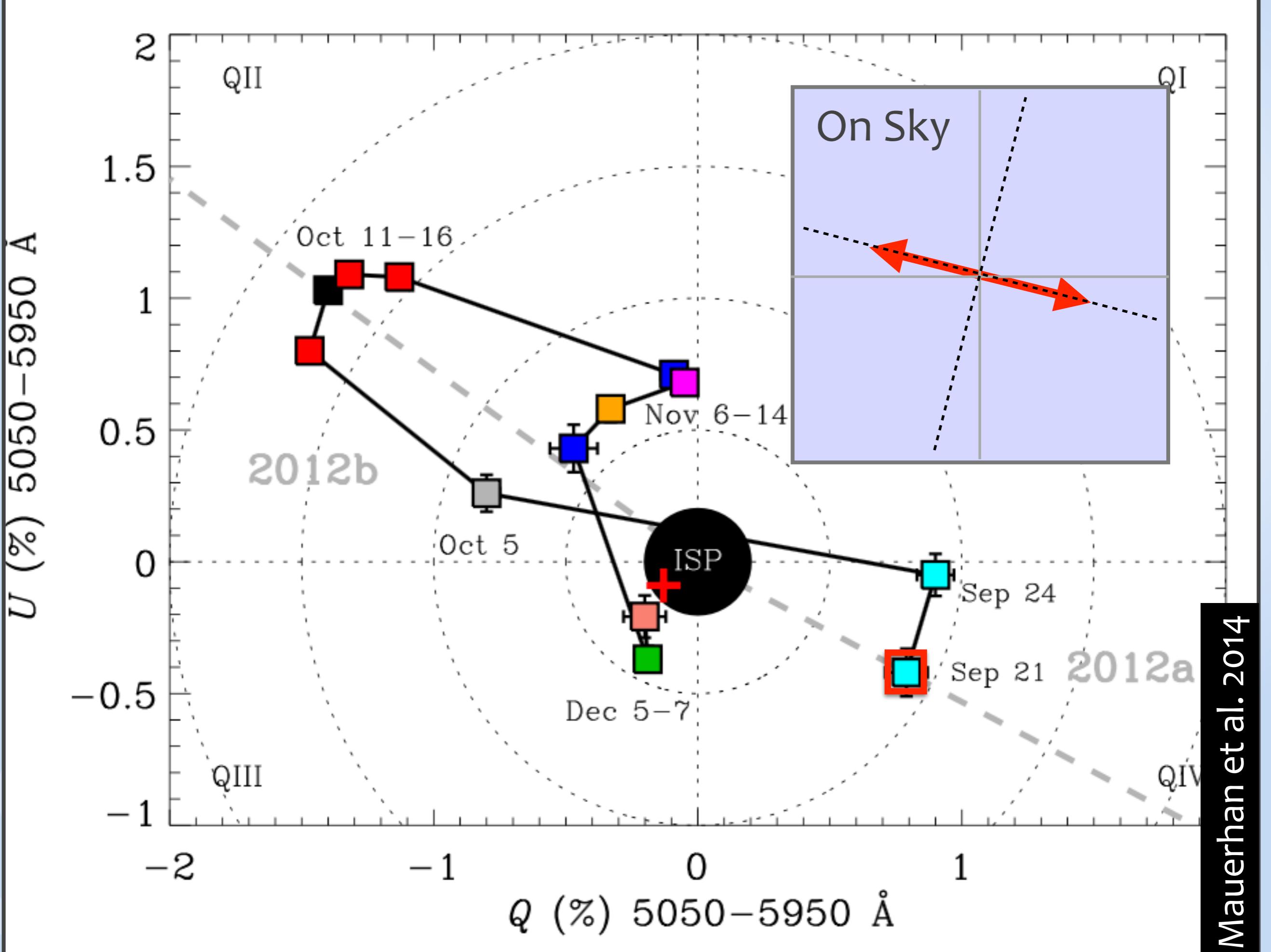
20"
↔

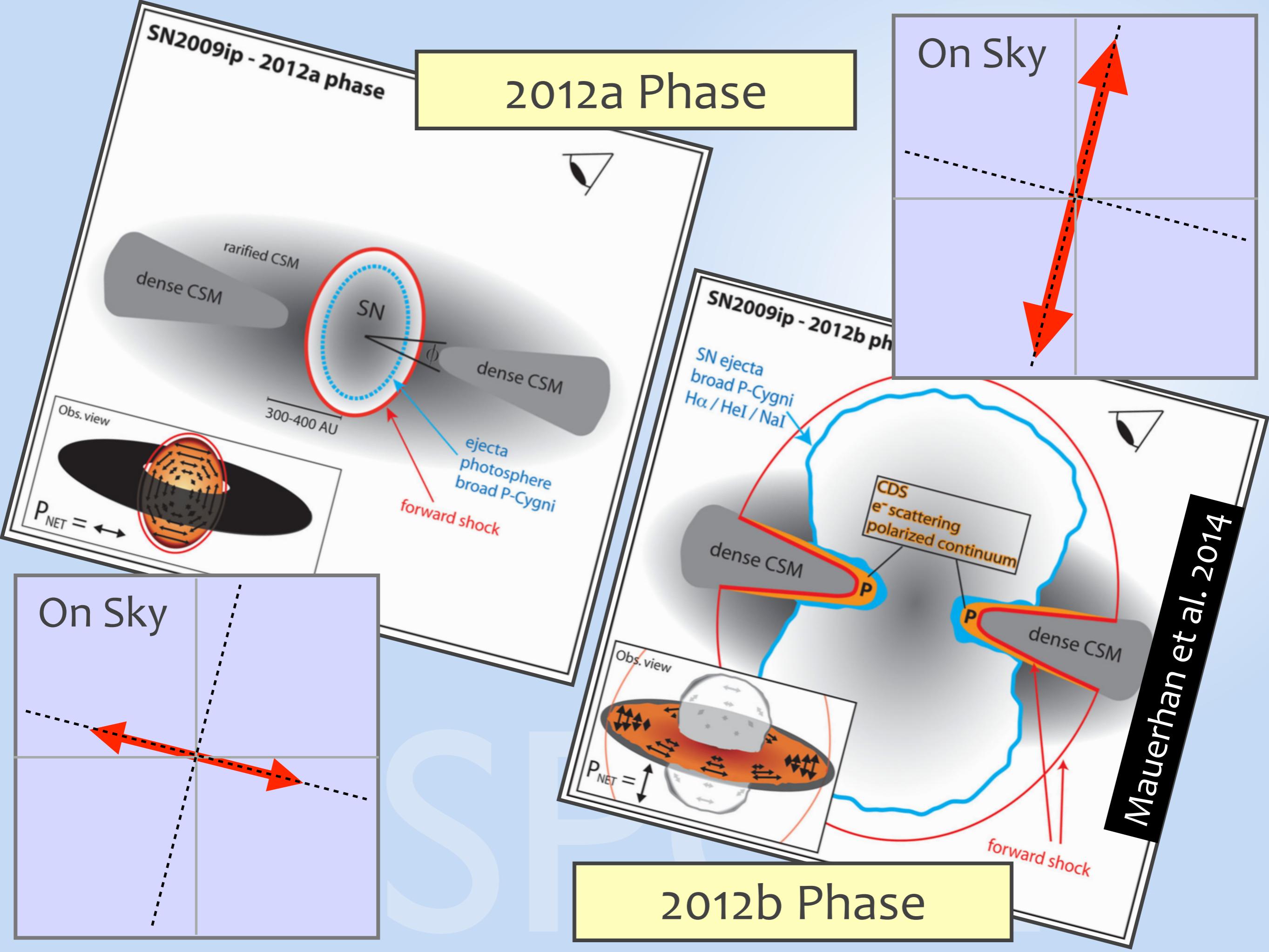
Mauerhan et al. 2014

SN 2009ip (2012) Polarized Light Curve







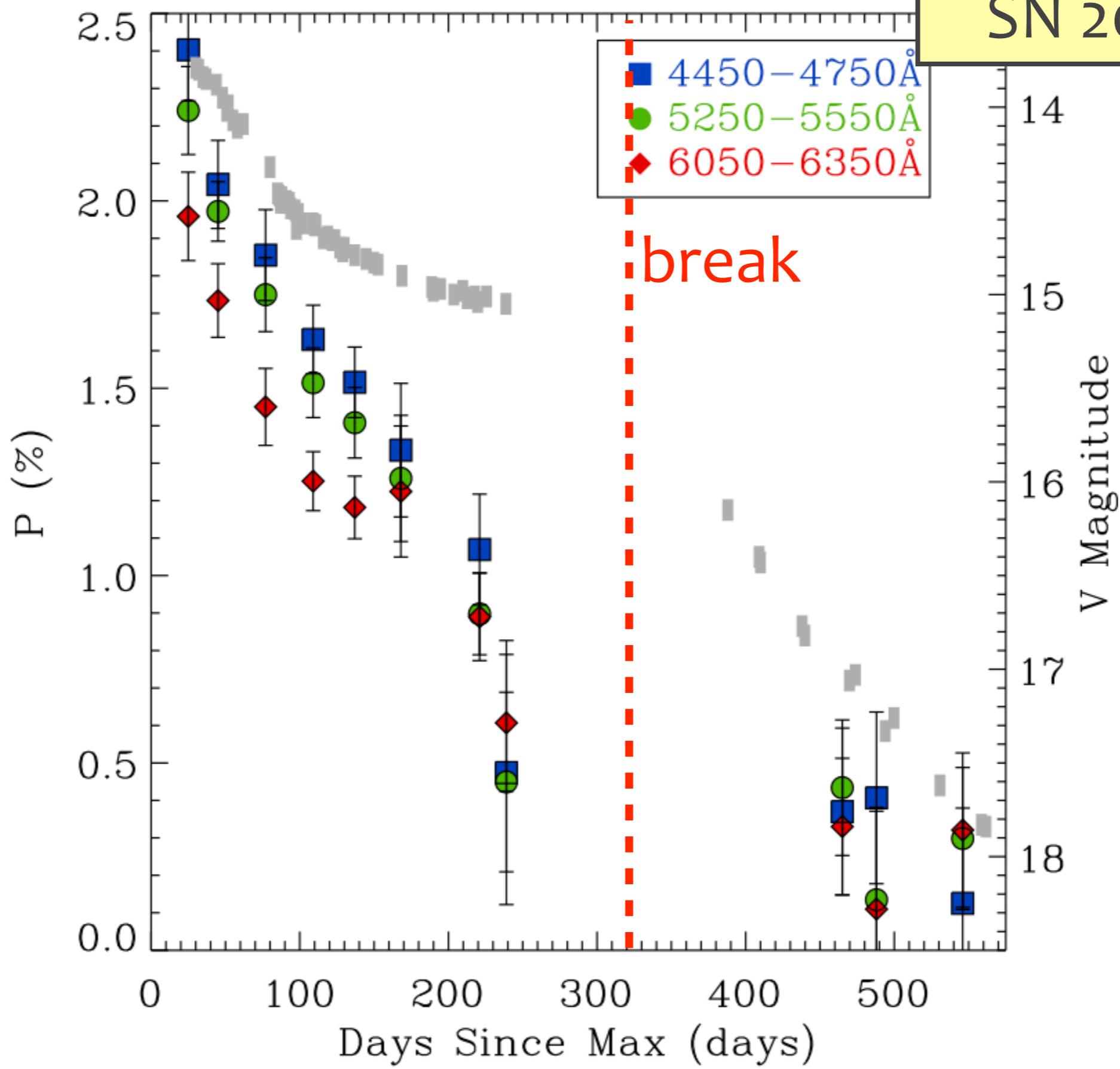




SN 2010jl
Type IIn
11 Epochs
Day 25 - 546

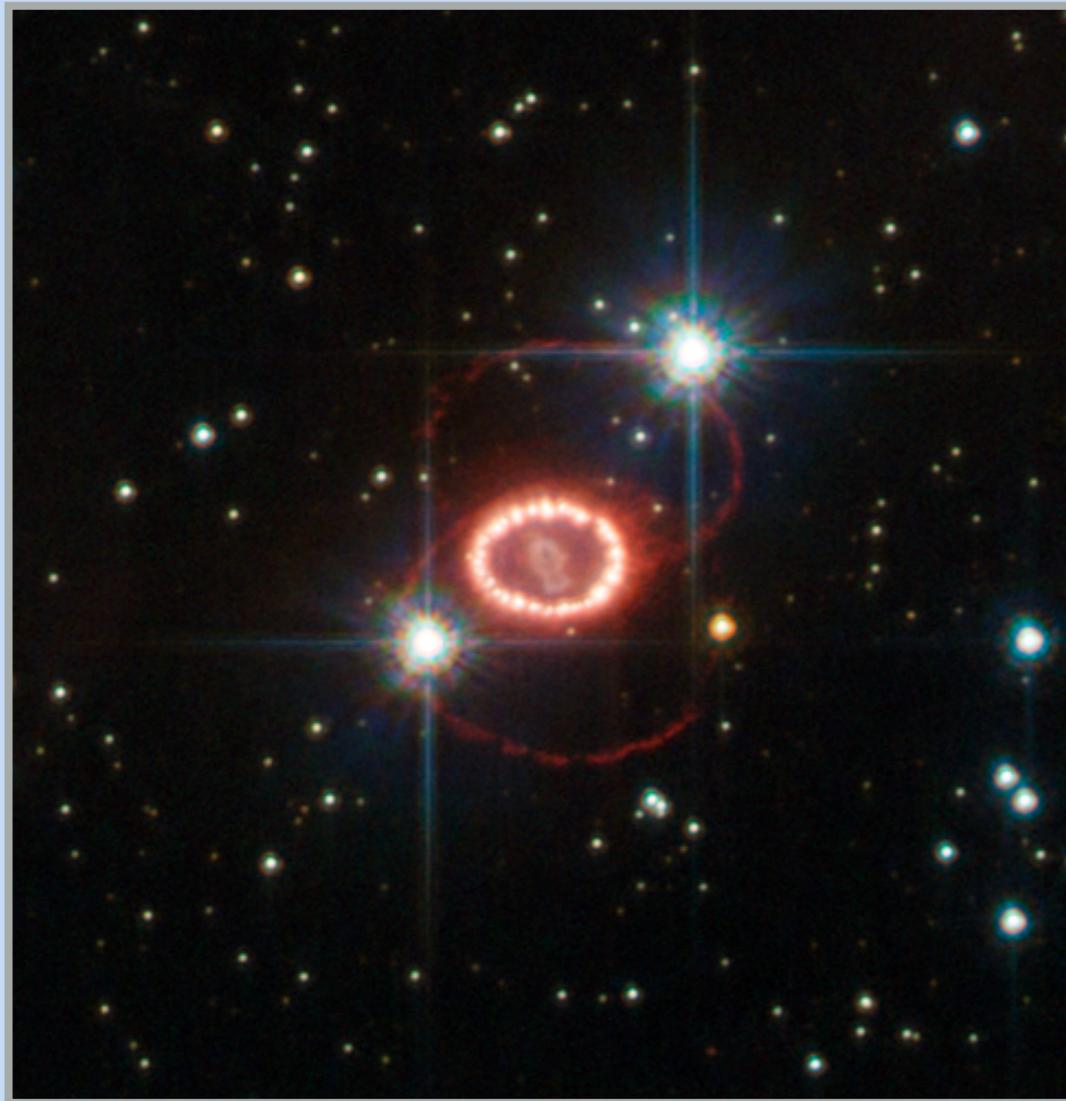
NASA/CXC/Royal Military College of
Canada/P.Chandra et al

SN 2010jl

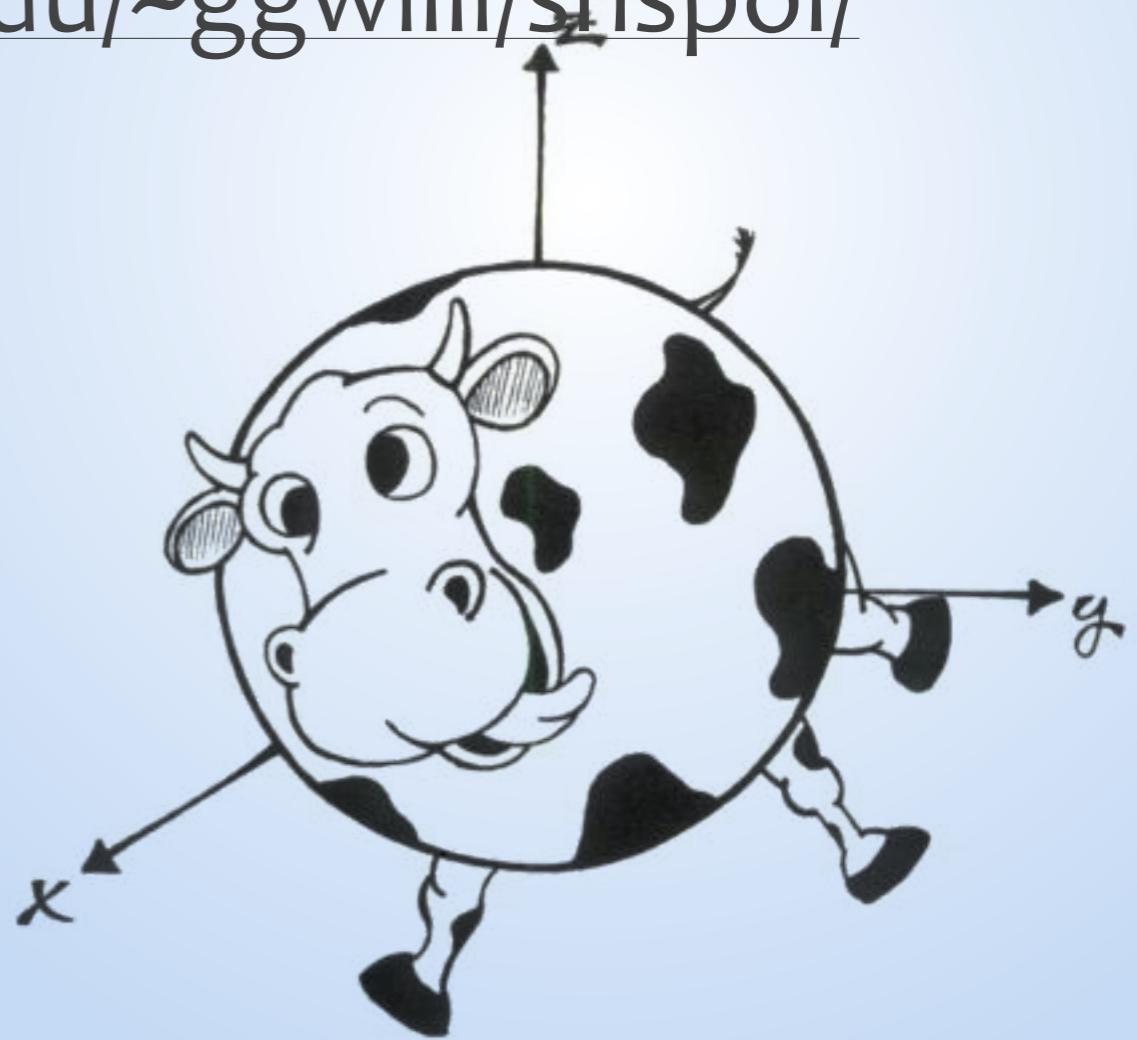


SNSPOL: Multi-epoch spectropolarimetry is a powerful tool for studying the shapes of supernovae.

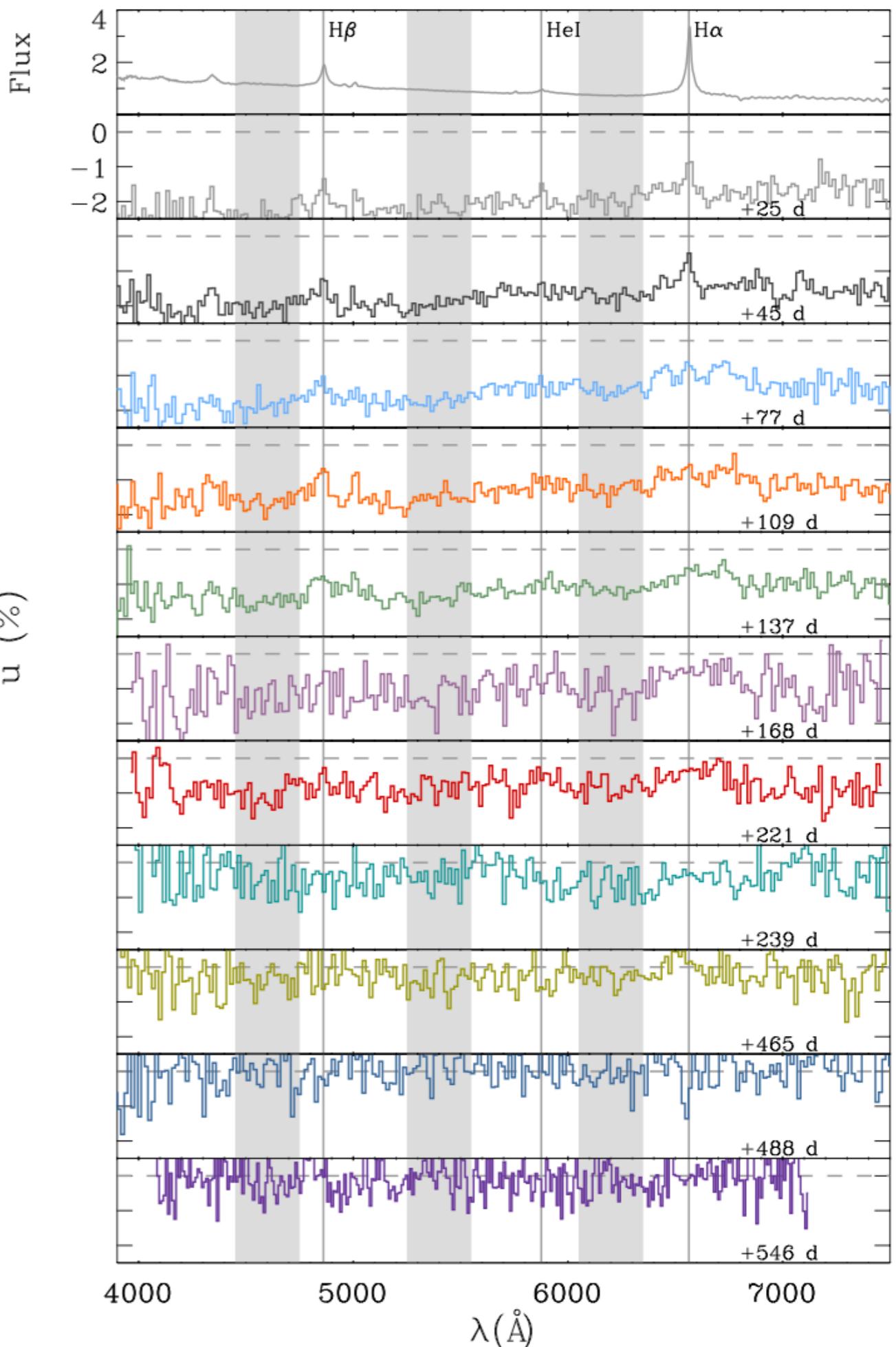
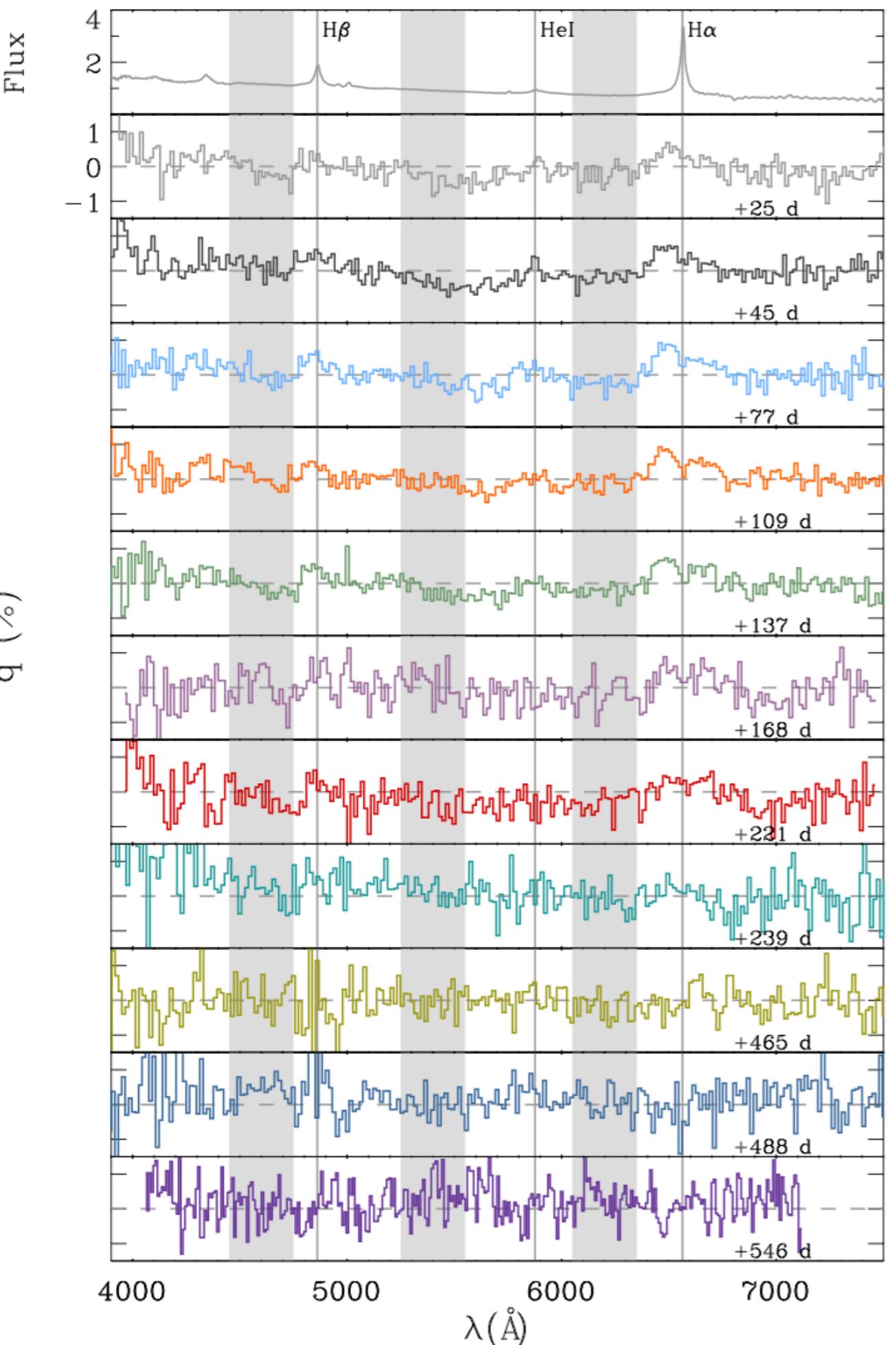
<http://grb.mmto.arizona.edu/~ggwilli/snspol/>



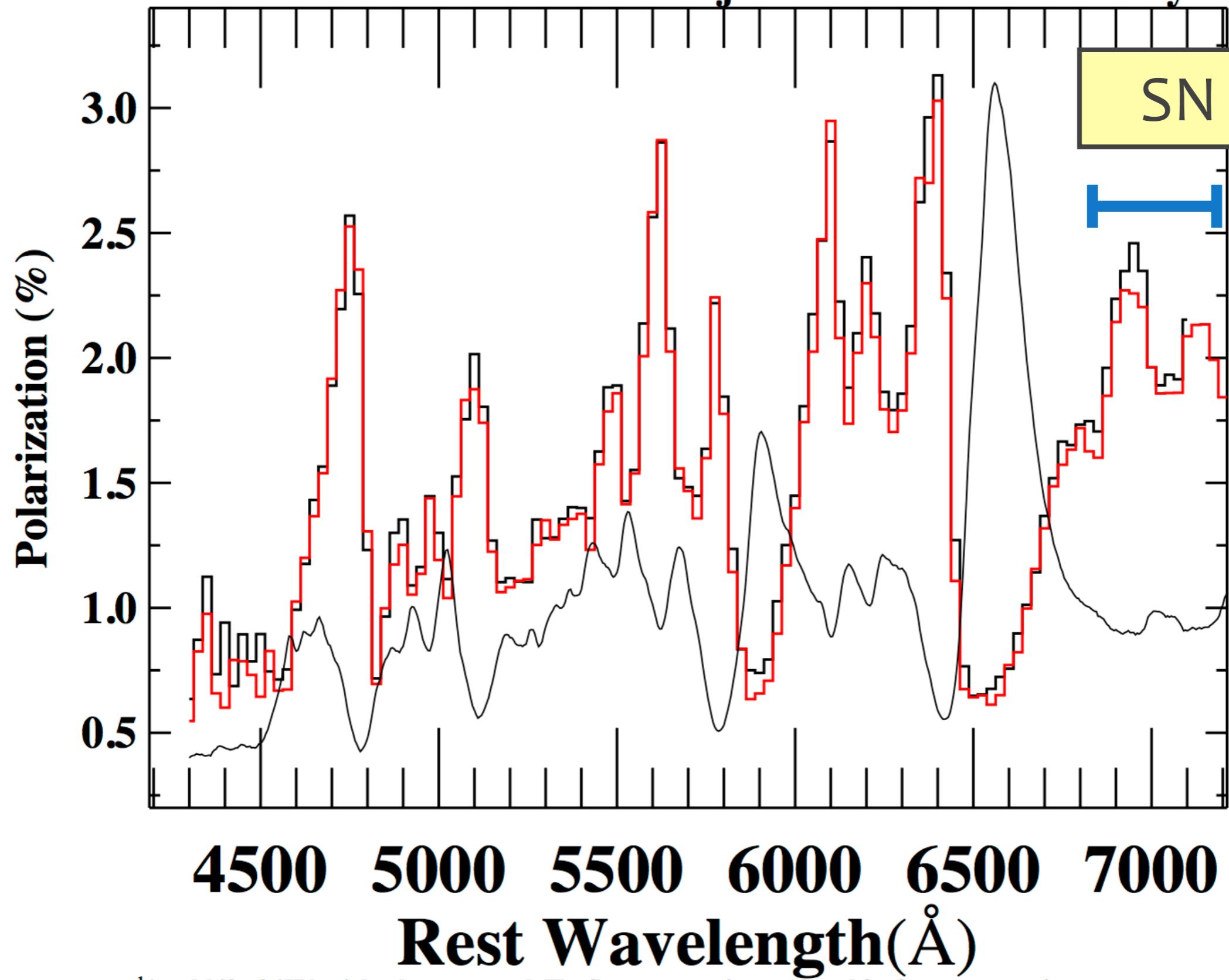
\neq



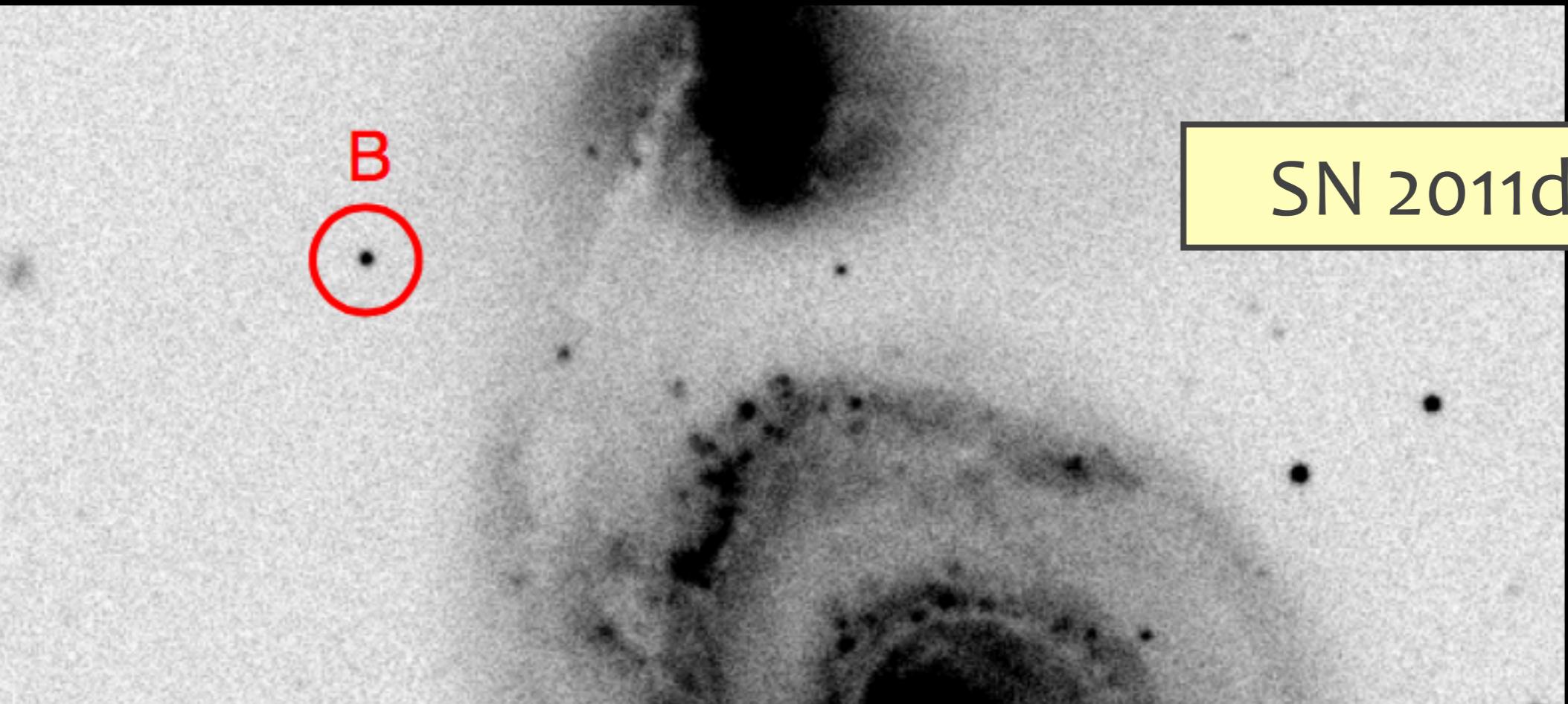
Thank You!



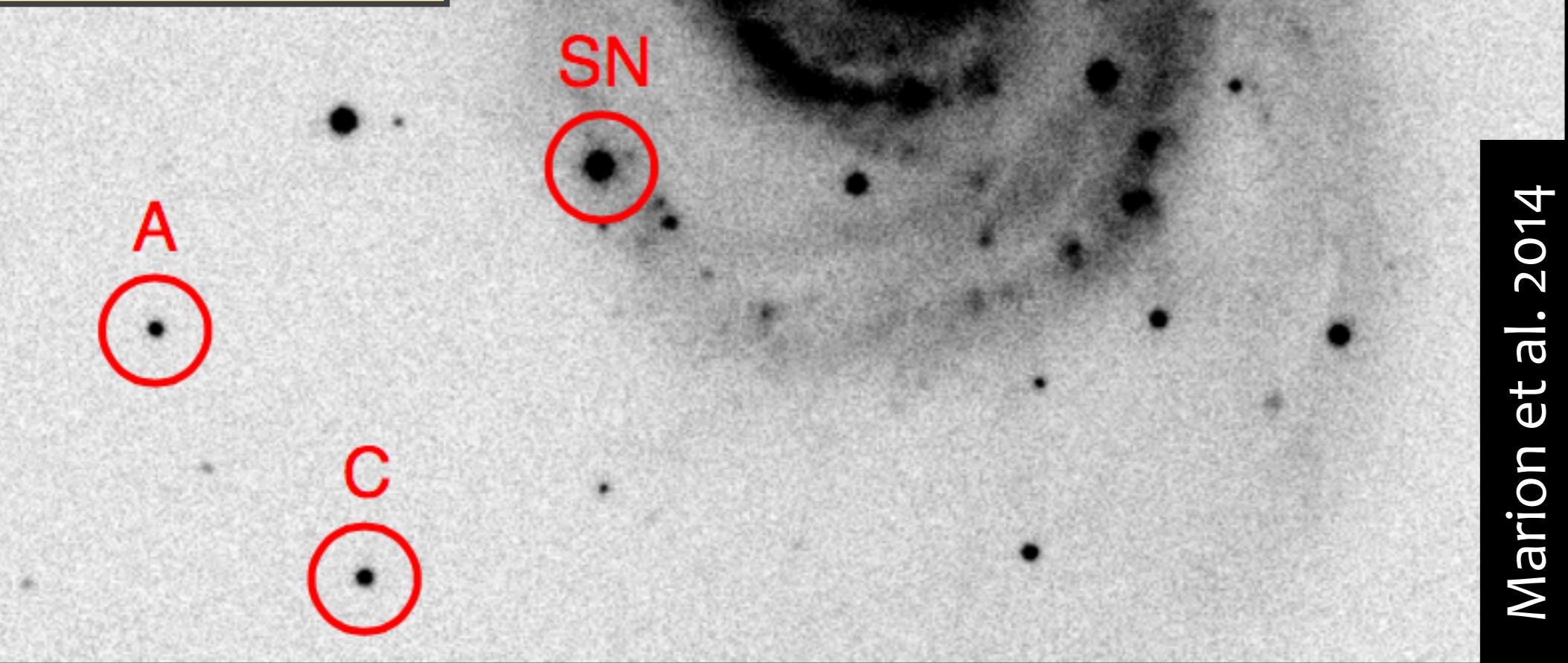
The Polarization of SN 2013ej: VLT vs. MMT on day 96



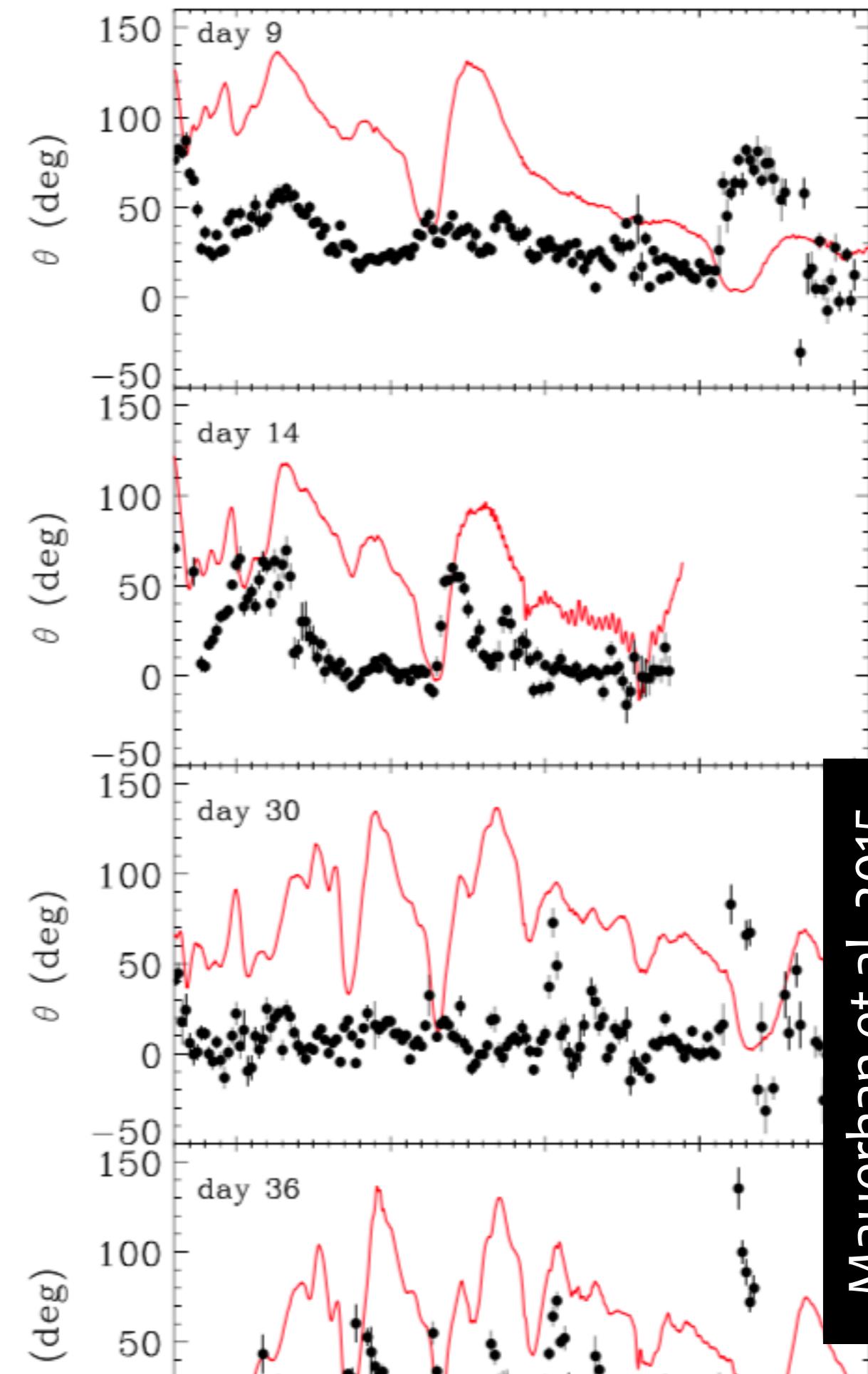
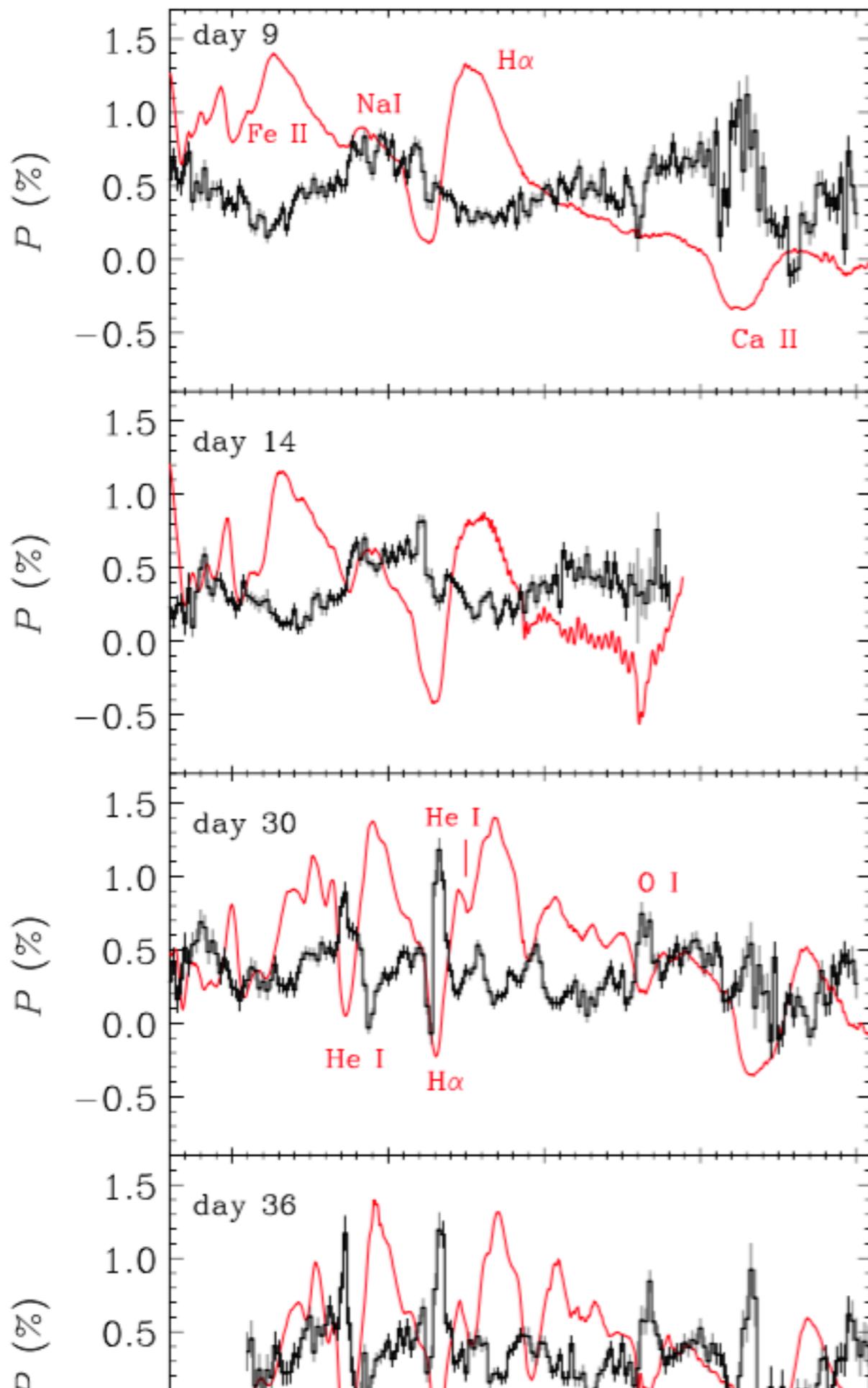
¹A redshift of 657 km/s has been removed. The flux spectrum is overplotted for feature comparison.

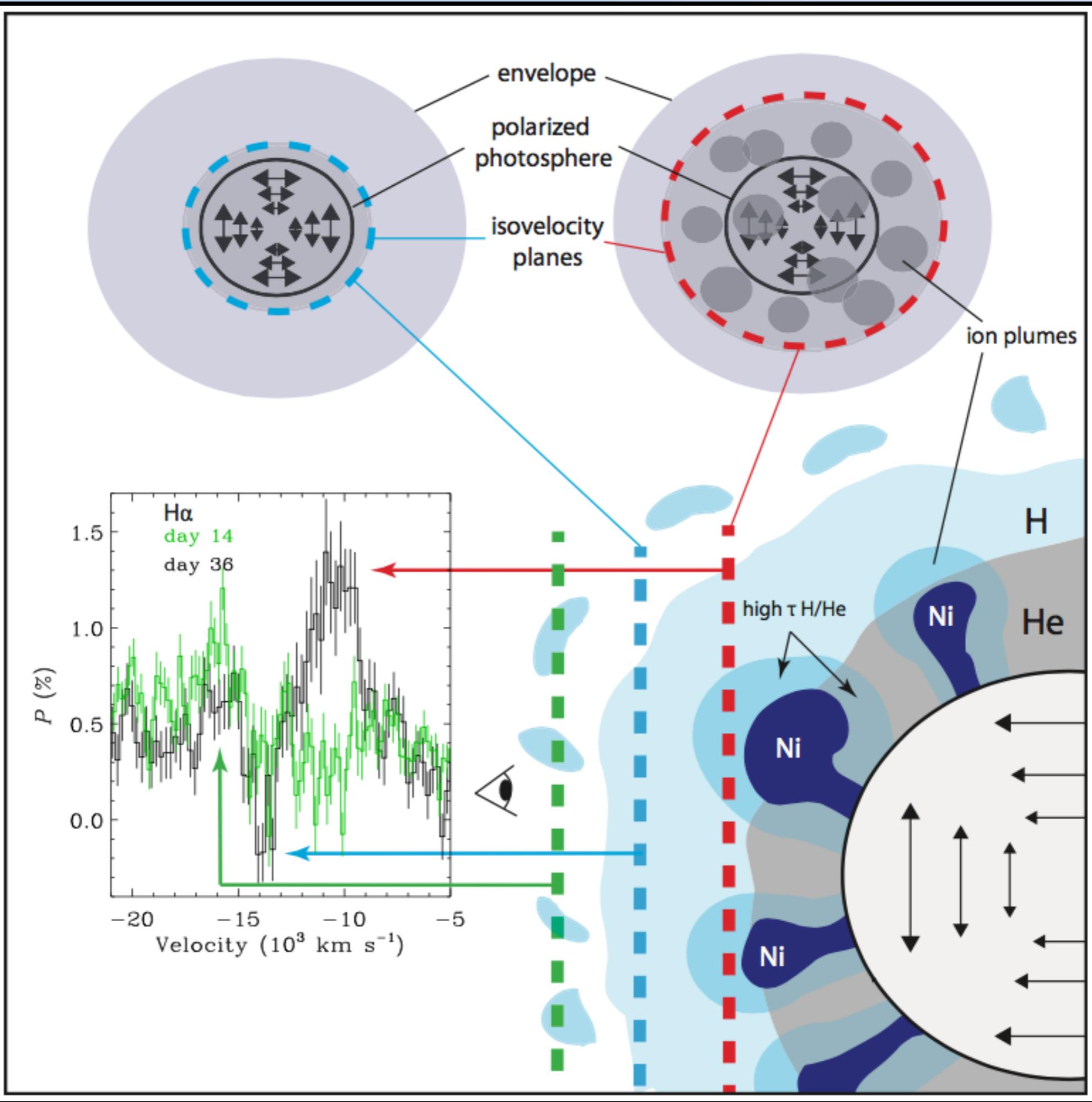


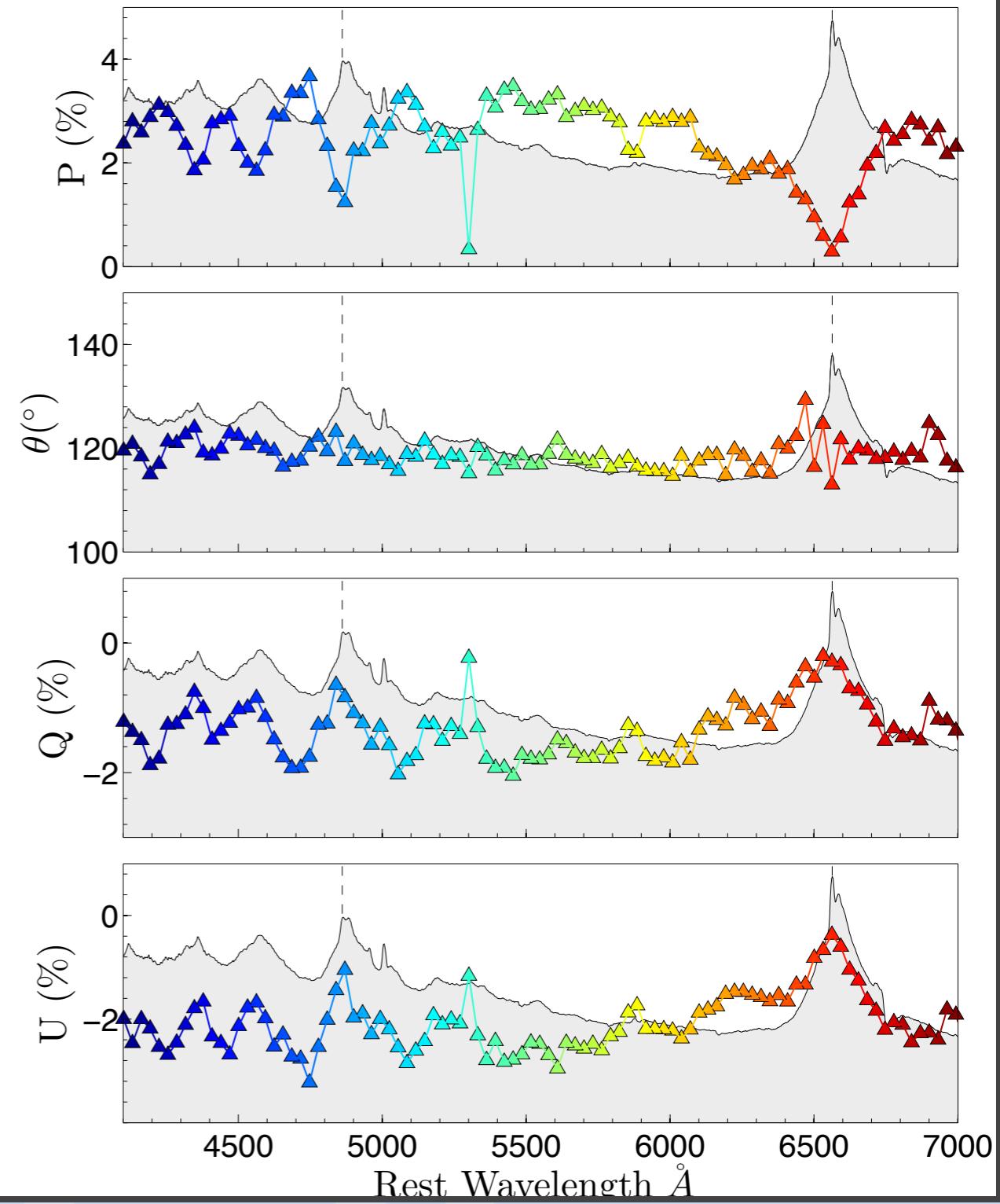
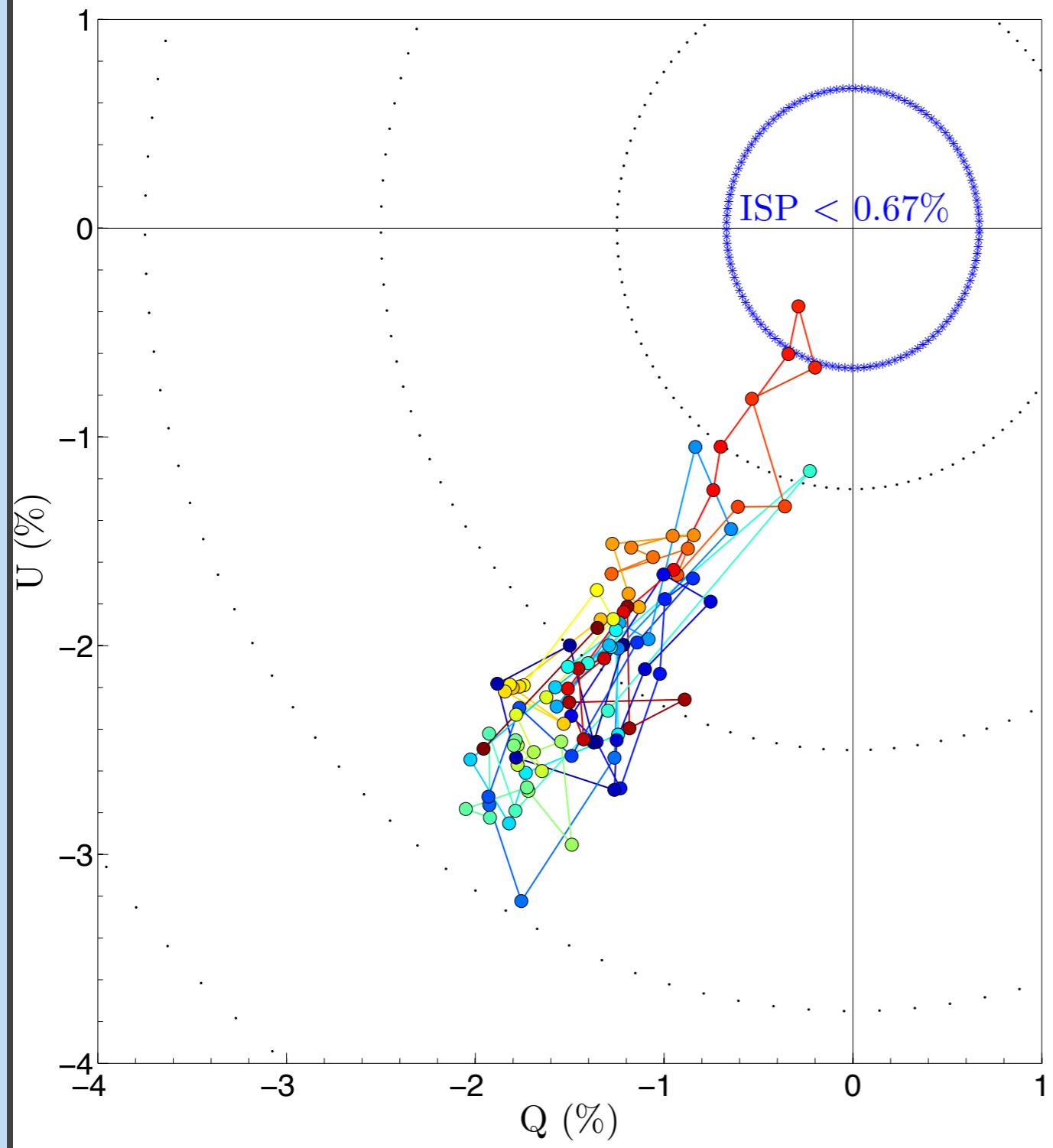
FLWO/KeplerCam

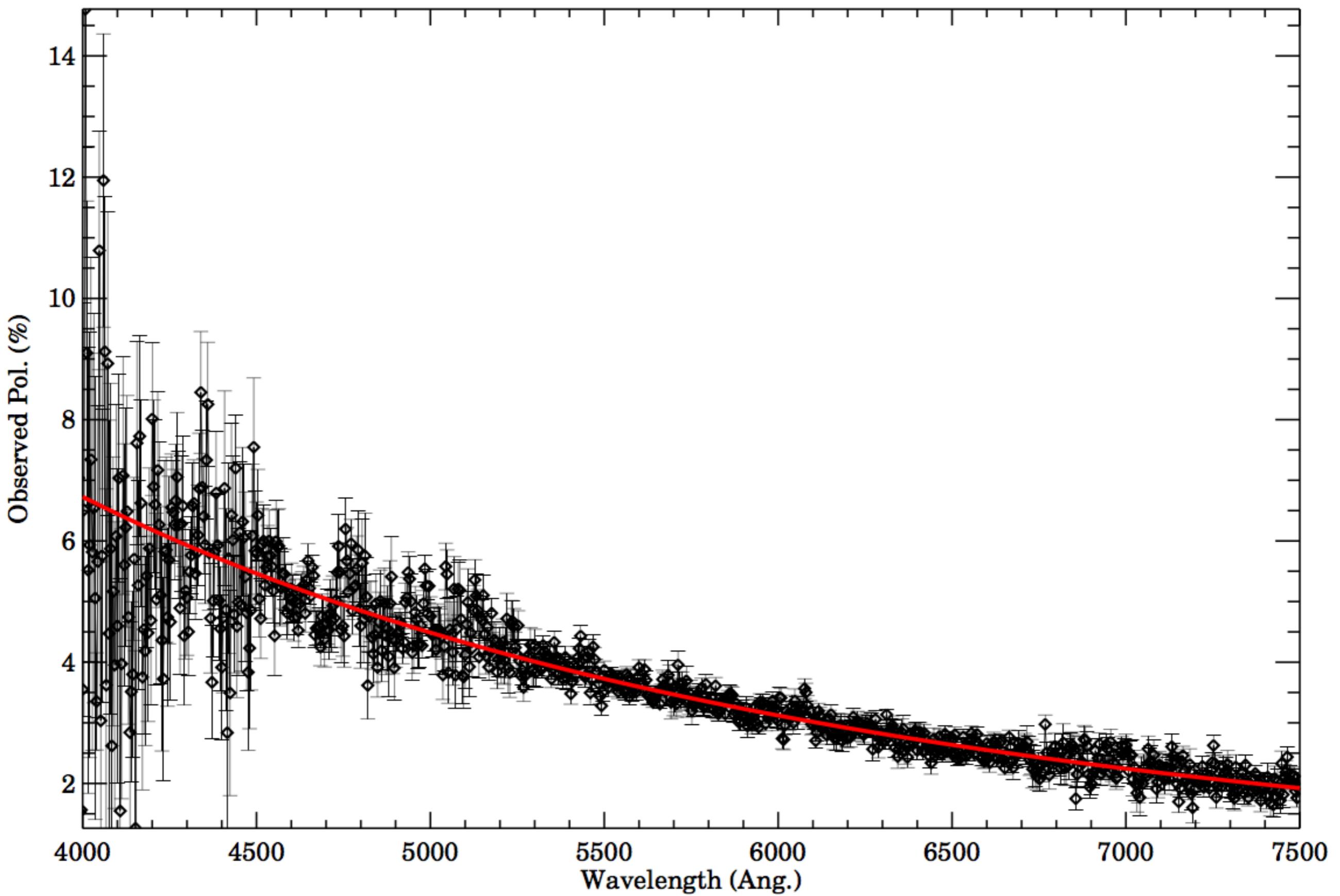


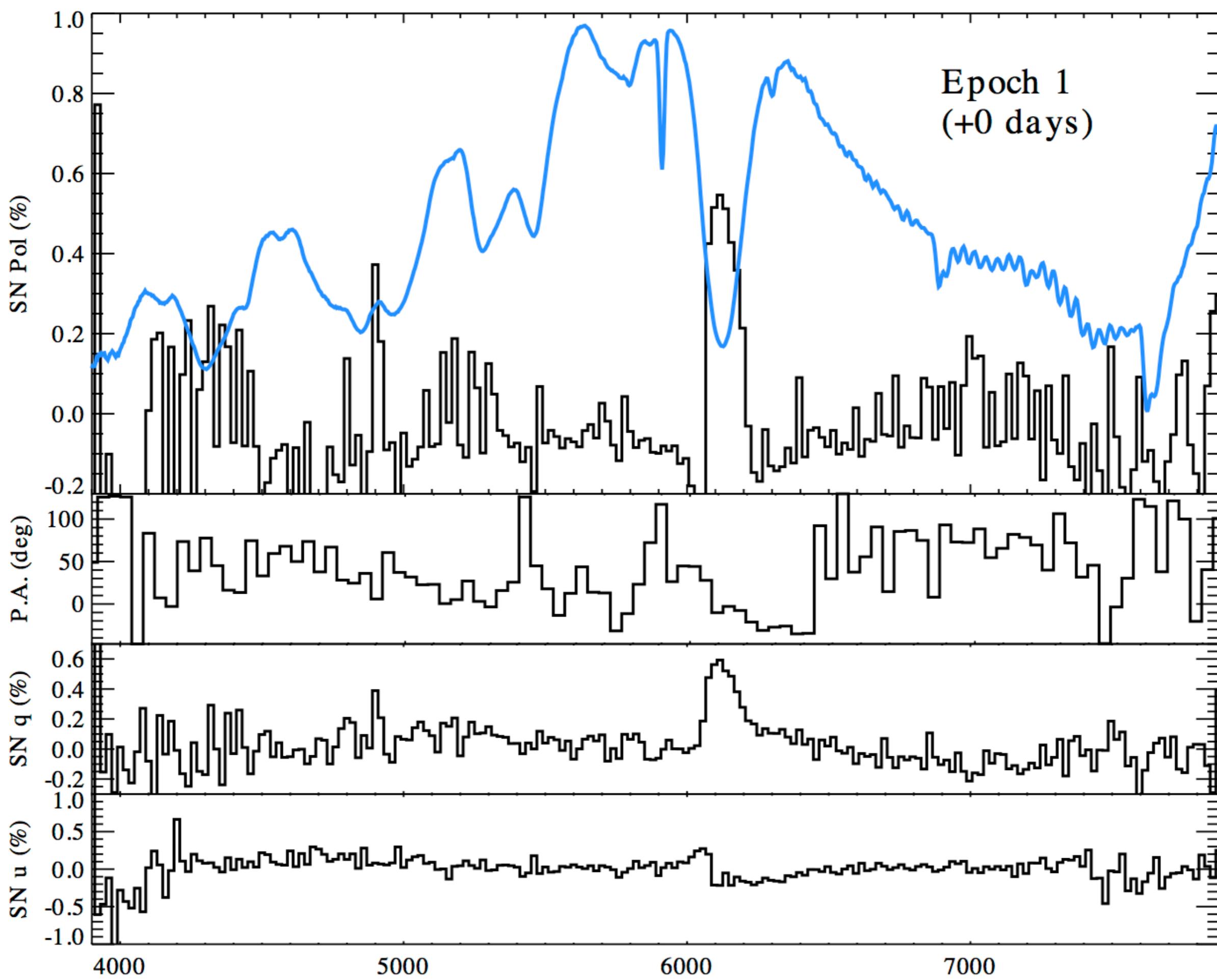
ISS

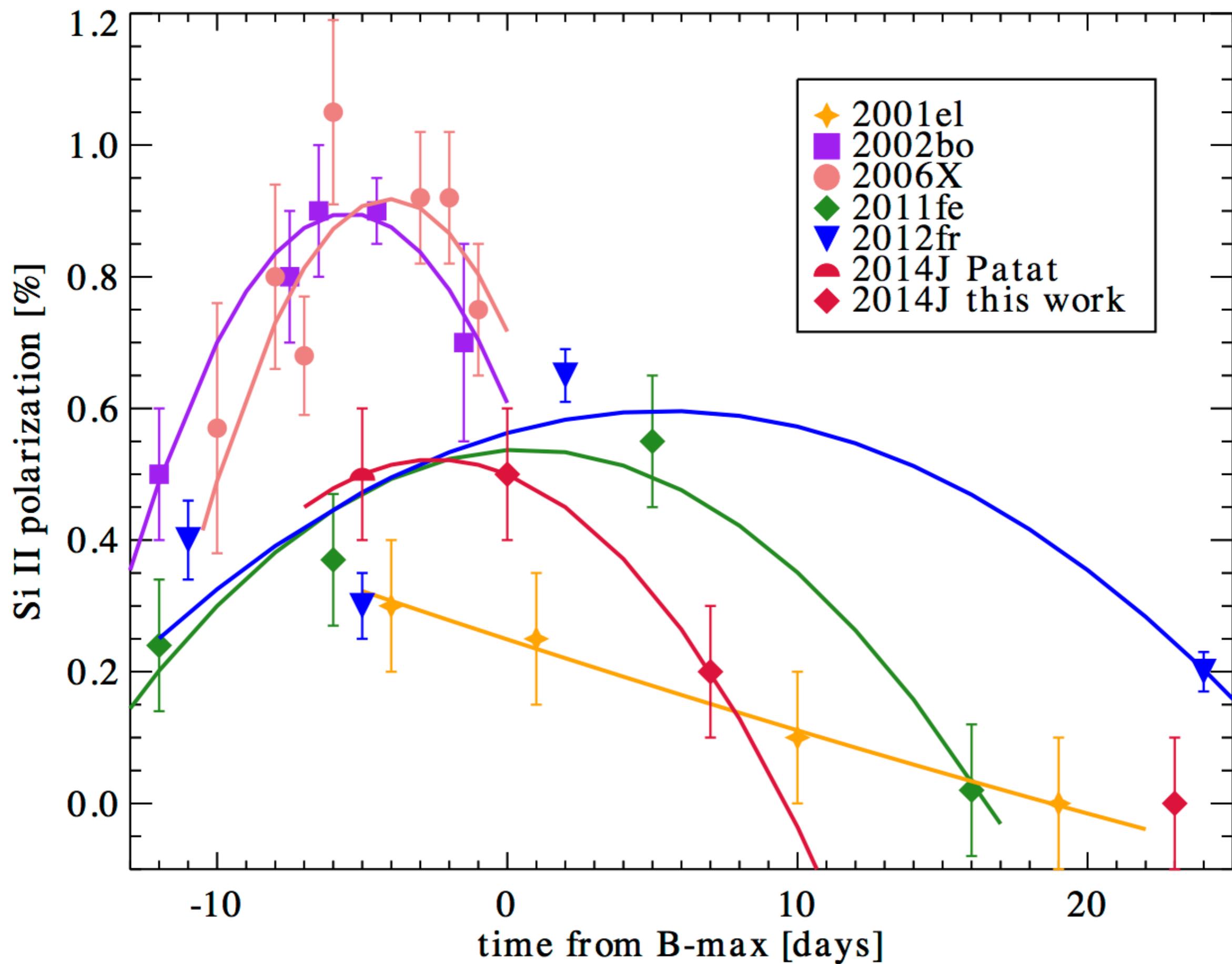


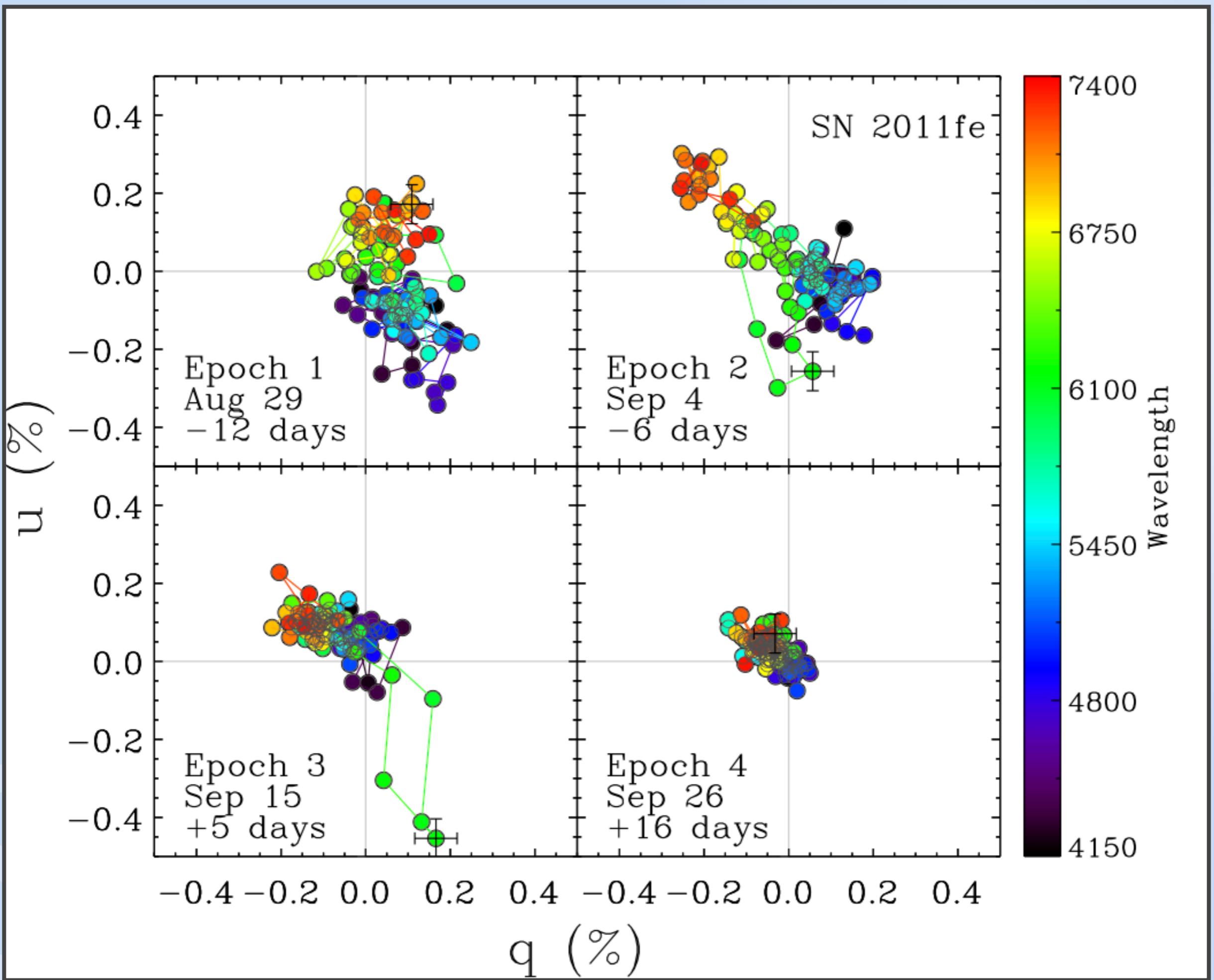


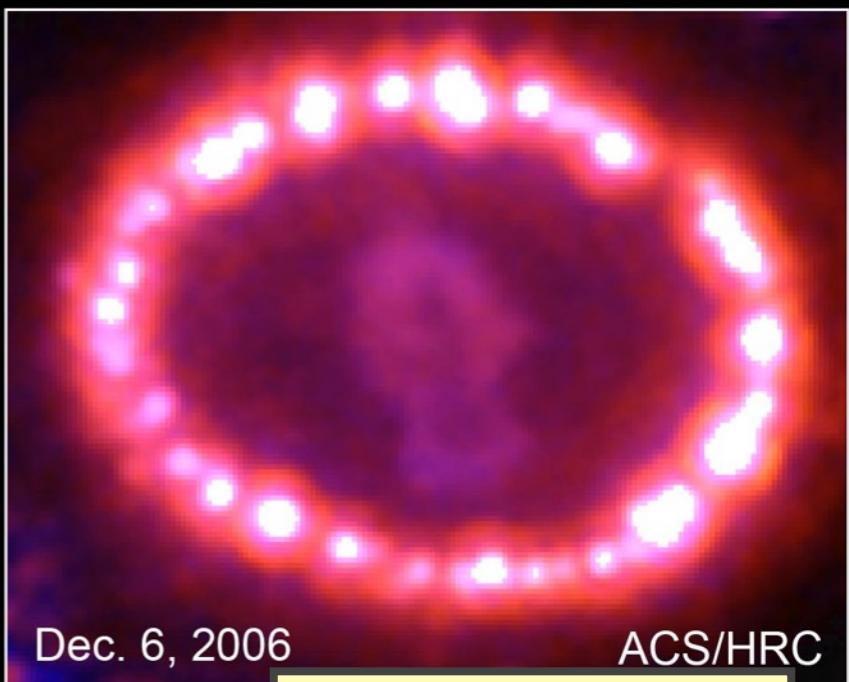
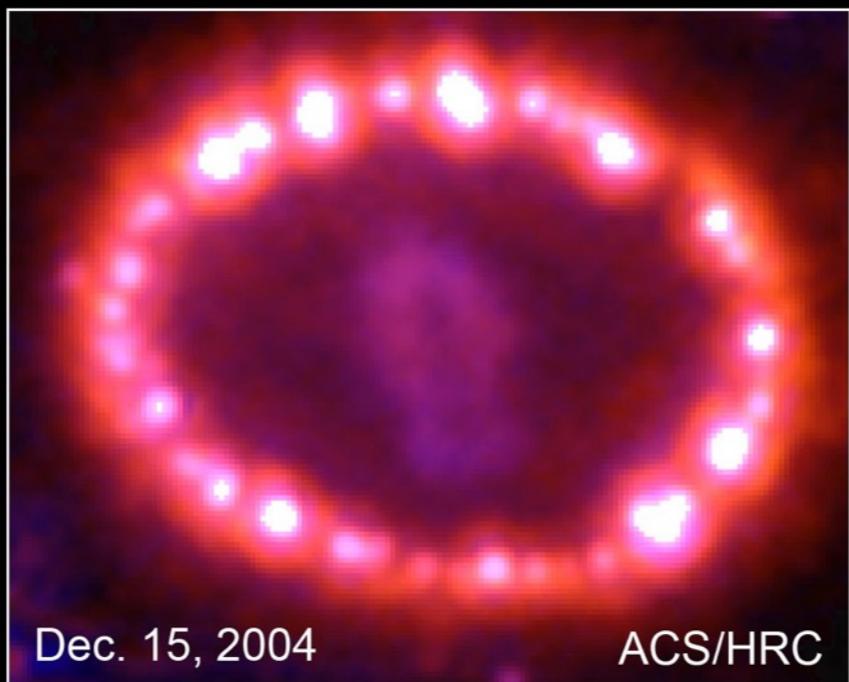
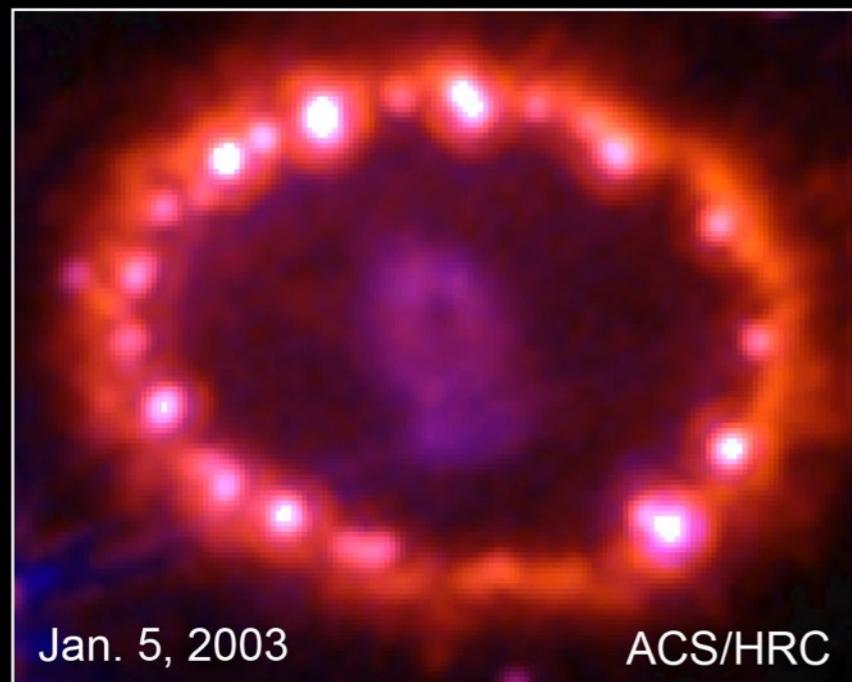
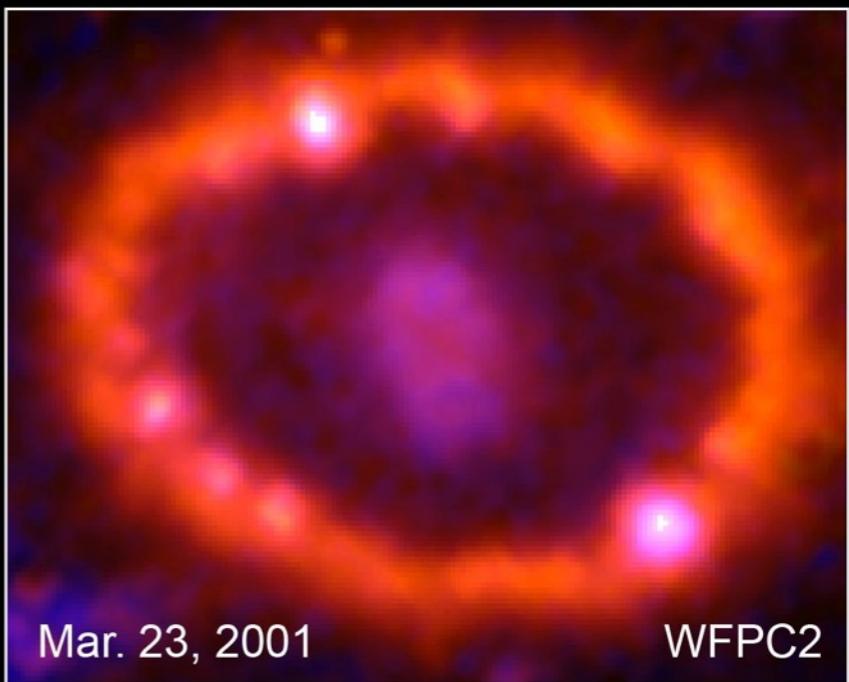
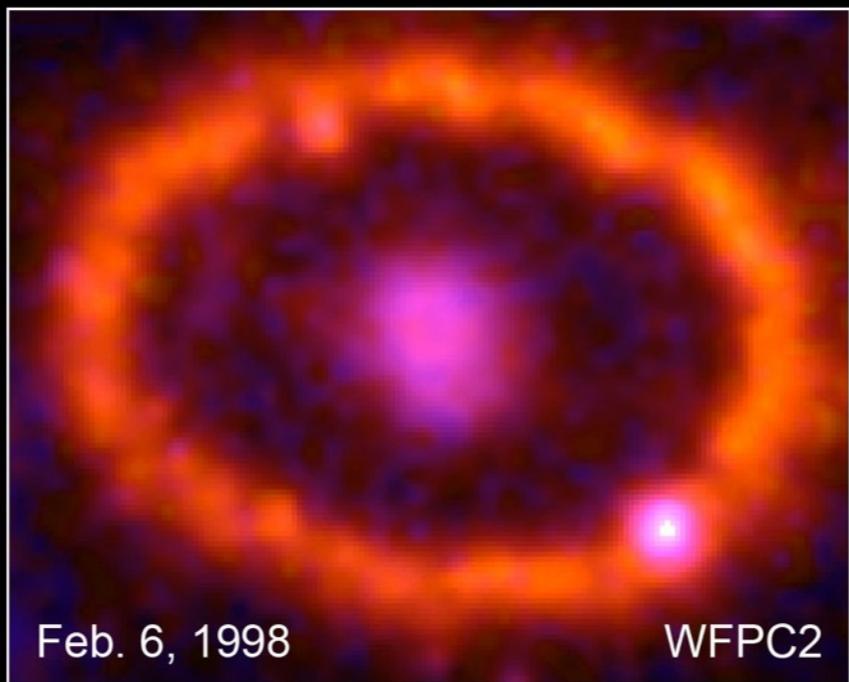
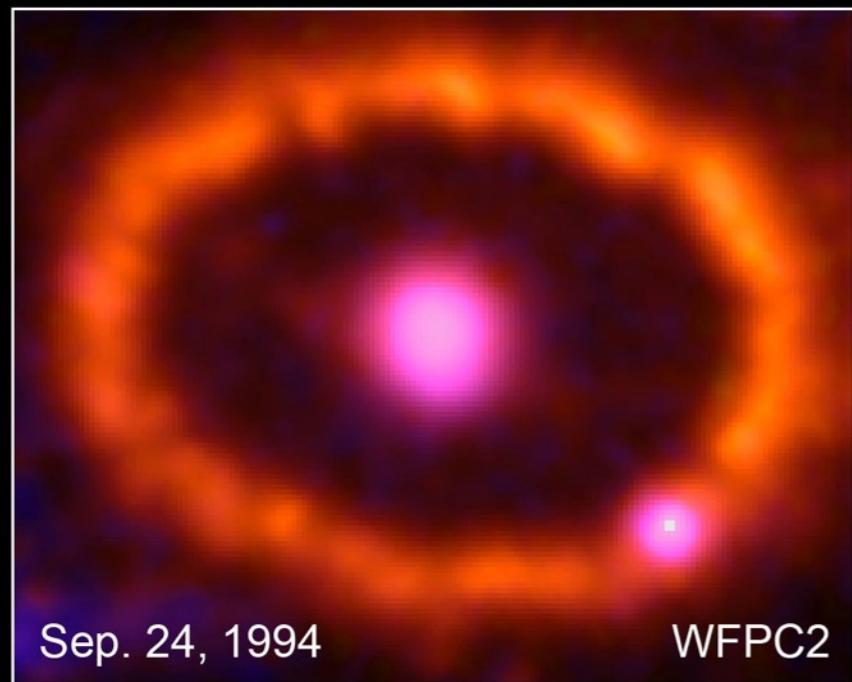












~20 Years After
Explosion

Supernova 1987A • 1994-2006

Hubble Space Telescope • WFPC2 • ACS