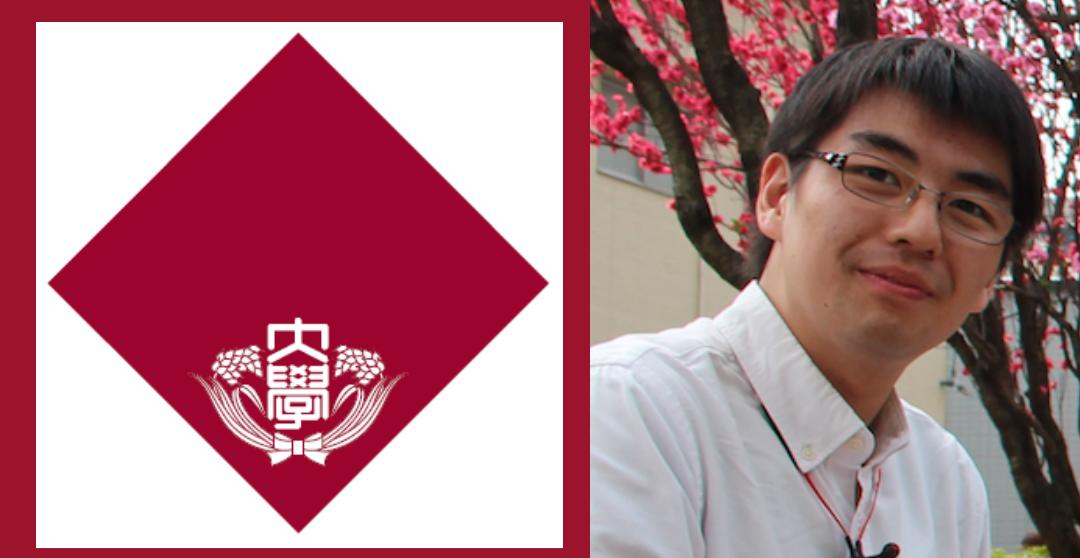


# Linear Analysis of Shock Dynamics in CCSNe

## ~Effect of Acoustic Injection and LESA~



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### Introduction and Main Purpose

Successful supernova explosions are supported by multidimensional effects such as

- Standing Accretion Shock Instability (SASI)
- Acoustic mechanism
- Lepton-number Emission Self-sustained Asymmetry (LESA).

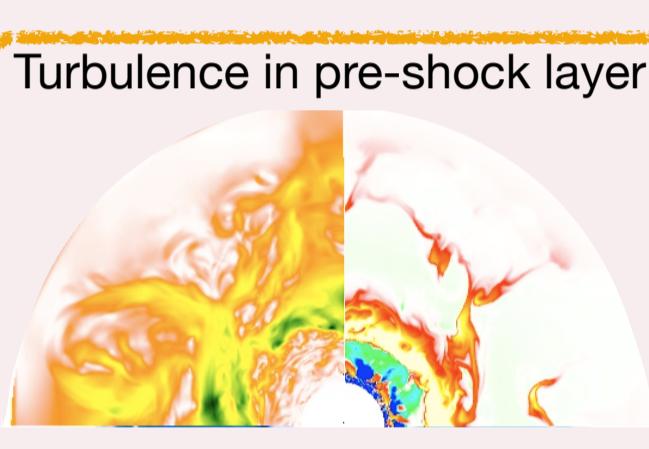
Main purpose of this research is investigating the importance of these multidimensional effects qualitatively by using the linear instability of standing shock.

### Result 1: Effects of acoustic injection

#### Boundary conditions

**Outer b.c. ( $r = r_{sh}$ ):** Linearized Rankine-Hugoniot condition

$$y^*(r_{sh}, s) = (sc + d) \frac{\delta r_{sh}^*(s)}{r_{sh}} + Rz^*(r_{sh}, s)$$



**Inner b.c. ( $r = r_\nu$ ):**

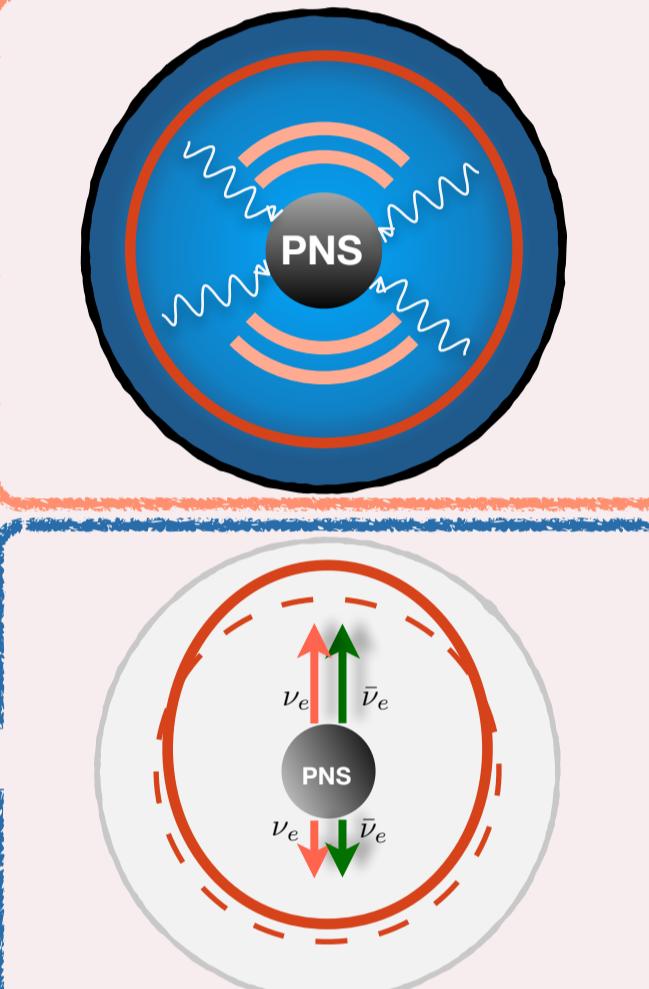
**Acoustic injection**

$$\frac{\delta p}{v_{r0} c_s \rho_0} + \frac{\delta v_r}{v_{r0}} = \sin(\omega_{PNS} t)$$

Outgoing acoustic mode

**Fluctuations of neutrino temperature**

$$\left( \frac{\partial P}{\partial Y_e} \right)_{\rho, T} \delta Y_e(r_\nu, t) + \left( \frac{\partial P}{\partial T} \right)_{\rho, Y_e} \delta T_\nu(t) = 0$$

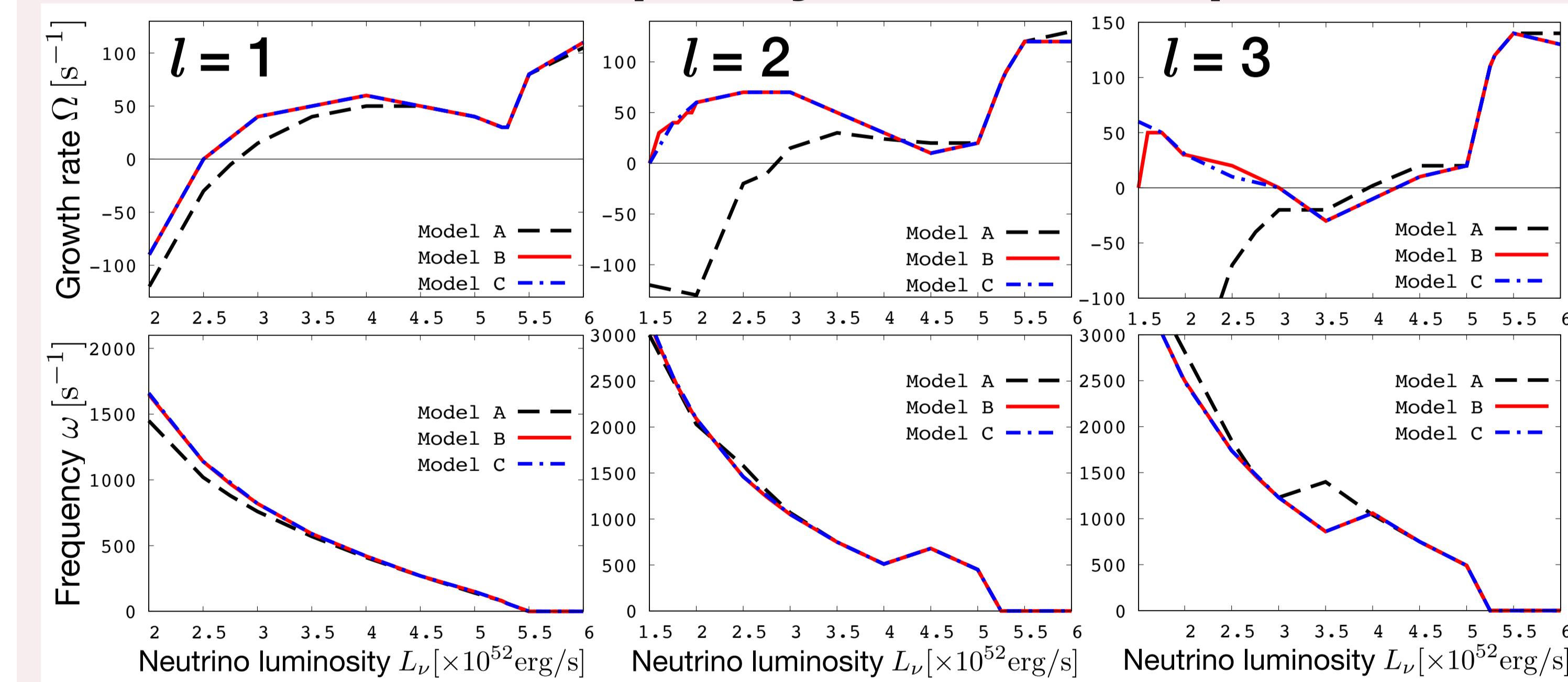


#### Eigenmodes expansion

$$\frac{\delta r_{sh}(t)}{r_{sh}} = \sum_i a_i e^{\Omega_i t} e^{i\omega_i t + \phi_i} \xrightarrow{\text{Laplace trans.}} \frac{\delta r_{sh}^*(s)}{r_{sh}} = \sum_i a_i \frac{e^{\phi_i}}{(s - \Omega_i) - i\omega_i}$$

$\Omega_i$ : growth rate,  $\omega_i$ : frequency

#### Growth rate and frequency of dominant pole

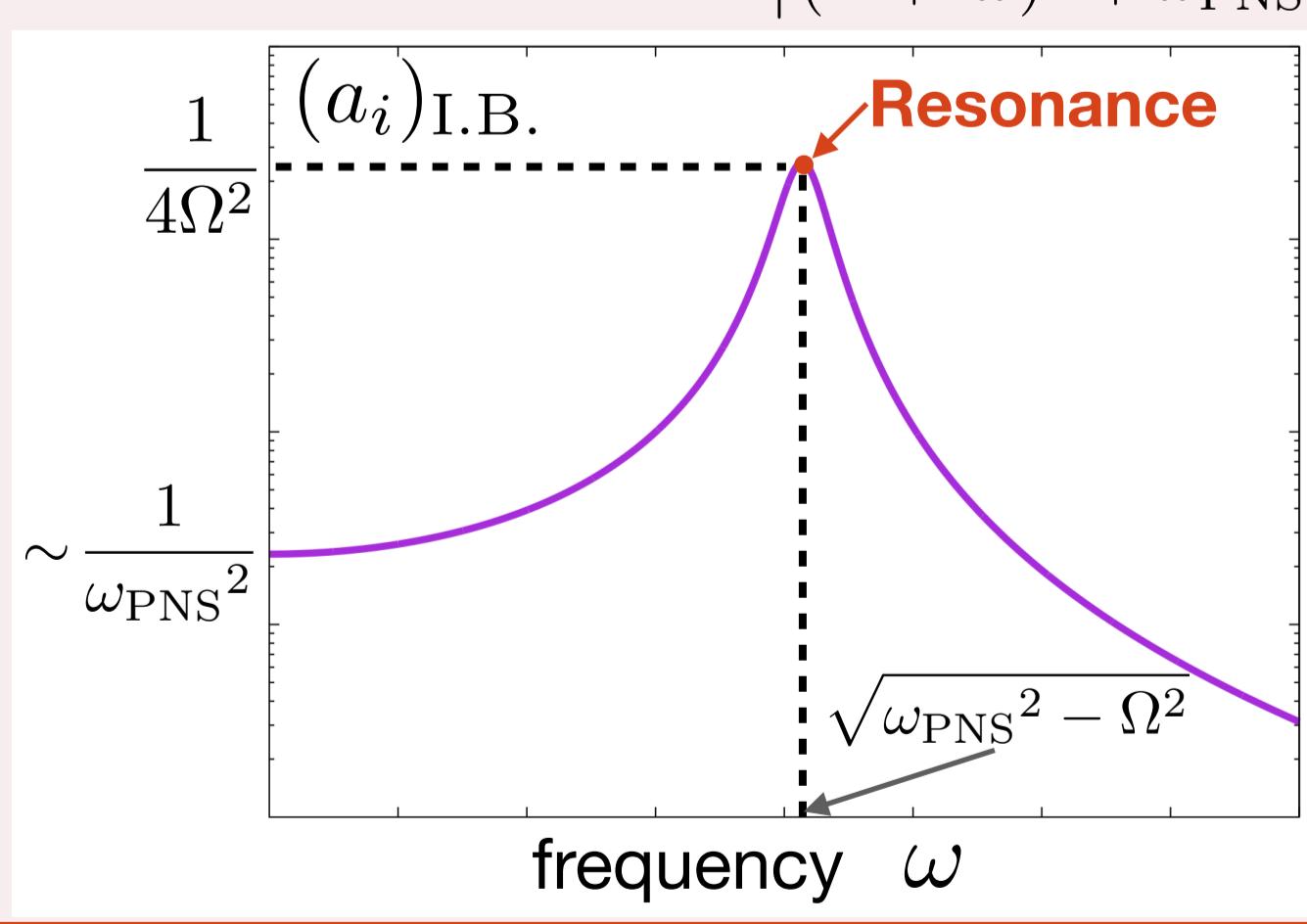
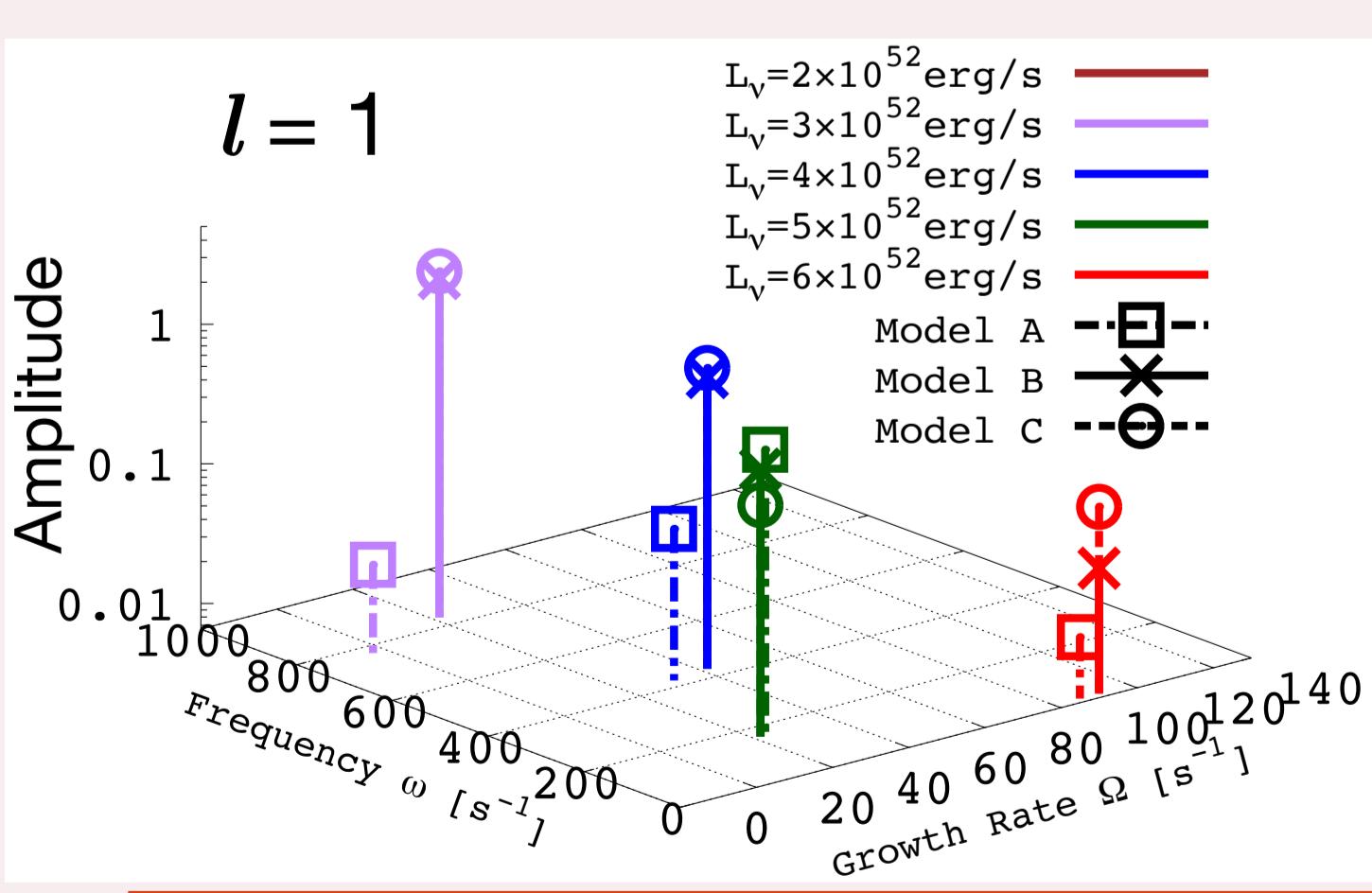


Growth rates of shock deformation are enhanced by acoustic injection.

#### Amplitude of dominant pole and resonance

$$a_i = (a_i)_{\text{upstream}} + (a_i)_{\text{I.B.}}$$

$$(a_i)_{\text{I.B.}} = |(\sin \omega_{PNS} t)^*| = \sqrt{\frac{\omega_{PNS}}{(\Omega + i\omega)^2 + \omega_{PNS}^2}}$$



→ Resonance of SASI freq. and g-mode freq. enhance the amplitude

### 5. Summary

Acoustic injection enhances instability of the standing shock especially for lower luminosity.

For lower luminosity, effect of acoustic injection is enhanced because of resonance of SASI and g-mode oscillation of PNS.

Fluctuation of neutrino luminosity does not influence linear instability of standing shock. Furthermore, LESA itself is not closely related with shock deformation.

### Basic Equations

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \mathbf{v}) = 0, \quad \frac{\partial}{\partial t}(\rho \mathbf{v}) + \nabla \cdot (\rho \mathbf{v} \mathbf{v} + p \mathbf{I}) = -\rho \frac{GM}{r^2} \frac{\mathbf{r}}{r}$$

$$\frac{d\varepsilon}{dt} + p \frac{d}{dt} \left( \frac{1}{\rho} \right) = q, \quad \frac{\partial}{\partial t} (nY_e) + \nabla \cdot (nY_e \mathbf{v}) = \lambda$$

Linearization around spherical steady b.g. with spherical harmonics  $Y_{(l,m)}(\theta, \phi)$

+

Laplace transformation  $f^*(s) := \int_0^\infty f(t) e^{-st} dt$  ( $s \in \mathbb{C}$ )

$$M \frac{dy^{*(l,m)}}{dr}(r, s) = (sA^{(l)} + B) y^{*(l,m)}(r, s) - A^{(l)} y_0^{(l,m)}(r) + \mathbf{v} \delta T_\nu^{(l,m)}$$

$$\mathbf{y}^*(r, s) = \left( \frac{\delta \rho^*}{\delta \rho_0}, \frac{\delta v_r^*}{\delta v_{r0}}, \frac{\delta v_\perp^*}{\delta v_{r0}}, \frac{\delta \varepsilon^*}{\delta \varepsilon_0}, \frac{\delta Y_e^*}{\delta Y_{e0}}, \frac{\delta v_{rot}^*}{\delta v_{r0}} \right)^T, \quad \mathbf{v} = \left( 0, 0, 0, \frac{1}{v_{r0}} \frac{\partial q}{\partial T_\nu}, \frac{m_b}{\rho_0 Y_{e0}} \frac{\partial \lambda}{\partial T_\nu}, 0 \right)^T$$

### Result 2: Steady solution of perturbed eq.

**LESA** is known as systematic deformation of standing shock by asymmetric neutrino emission. But it is unclear which of asymmetry in lepton number flux  $F_{\nu_e}^n - F_{\bar{\nu}}^n$  or deformation of the sum  $F_{\nu_e} + F_{\bar{\nu}}$  is closely related to shock deformation.

#### Boundary conditions

**Outer b.c. ( $r = r_{sh}$ ):** Linearized Rankine-Hugoniot cond.

$$y^*(r_{sh}, s) = (sc + d) \frac{\delta r_{sh}^*(s)}{r_{sh}}$$

No turbulence in pre-shock layer

**Inner b.c. ( $r = r_\nu$ ):**

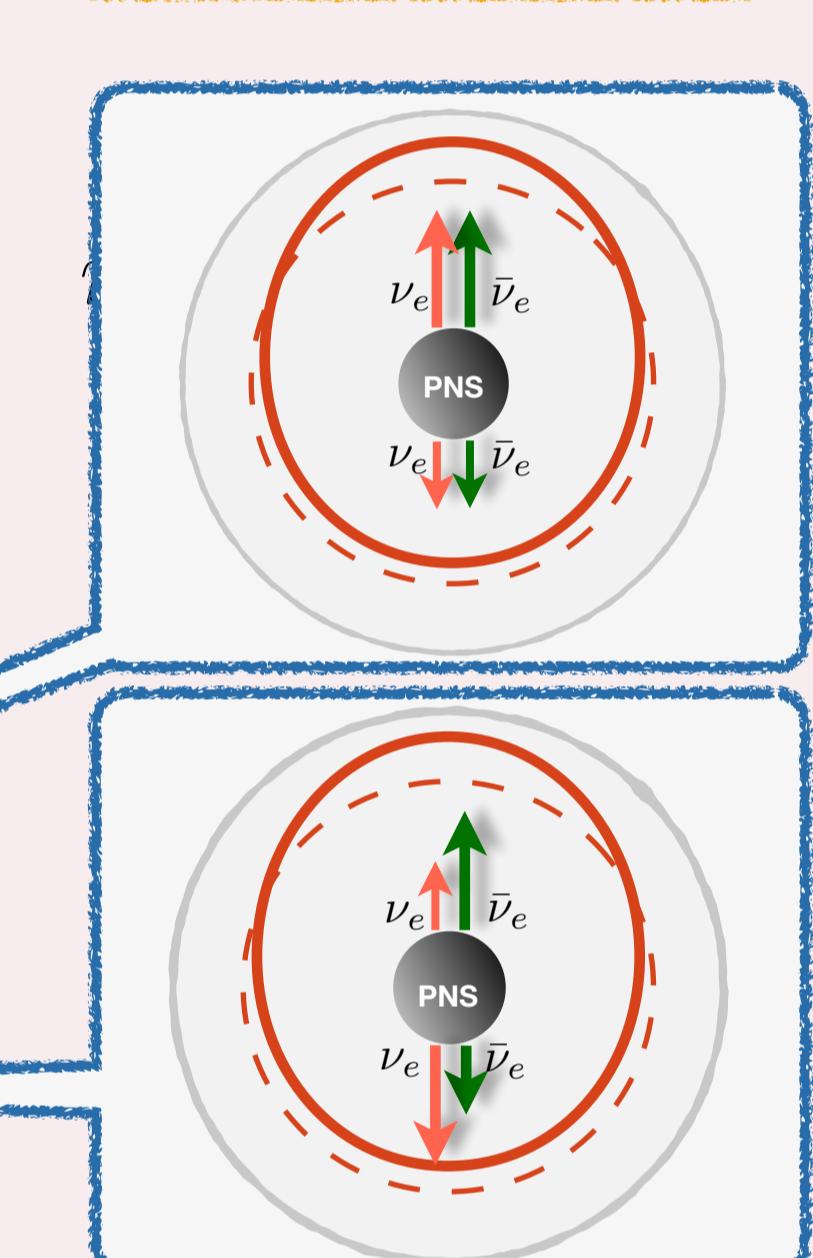
**Fluctuations of neutrino luminosity**

$$\frac{\delta L_{\nu_e}}{L_0} = 4 \frac{\delta T_{\nu_e}}{T_{\nu_e 0}} + c_{Y_e} \frac{\delta Y_e}{Y_{e0}}$$

$$\frac{\delta L_{\bar{\nu}_e}}{L_0} = 4 \frac{\delta T_{\bar{\nu}_e}}{T_{\bar{\nu}_e 0}} - c_{Y_e} \frac{\delta Y_e}{Y_{e0}}$$

We consider 2 cases.

$$\begin{cases} c_{Y_e} = 0 & (\Leftrightarrow \delta L_{\nu_e} - \delta L_{\bar{\nu}_e} = 0) \\ c_{Y_e} = 3.5 & (\Leftrightarrow \delta L_{\nu_e} - \delta L_{\bar{\nu}_e} \neq 0) \end{cases}$$

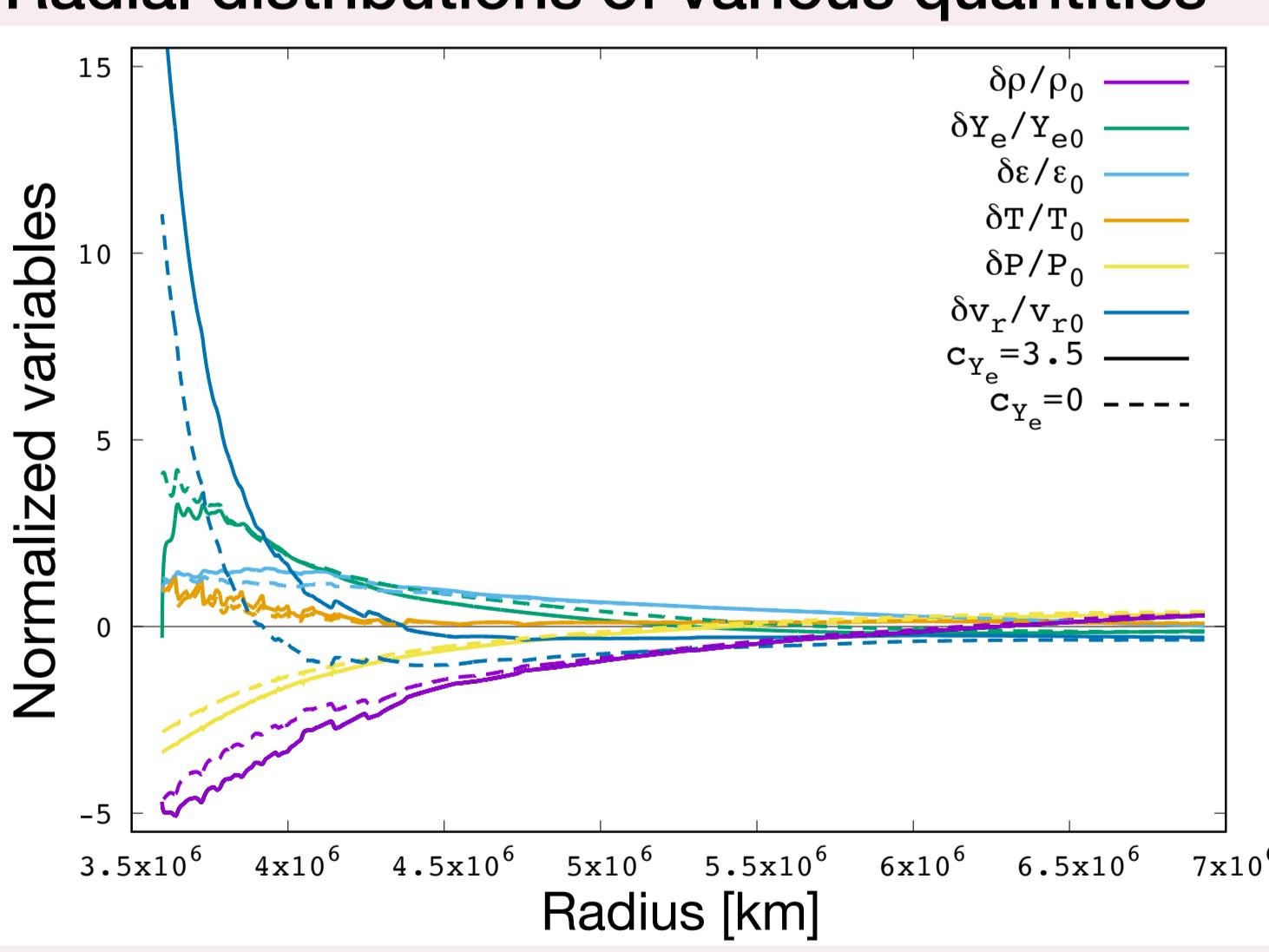


#### Steady solution

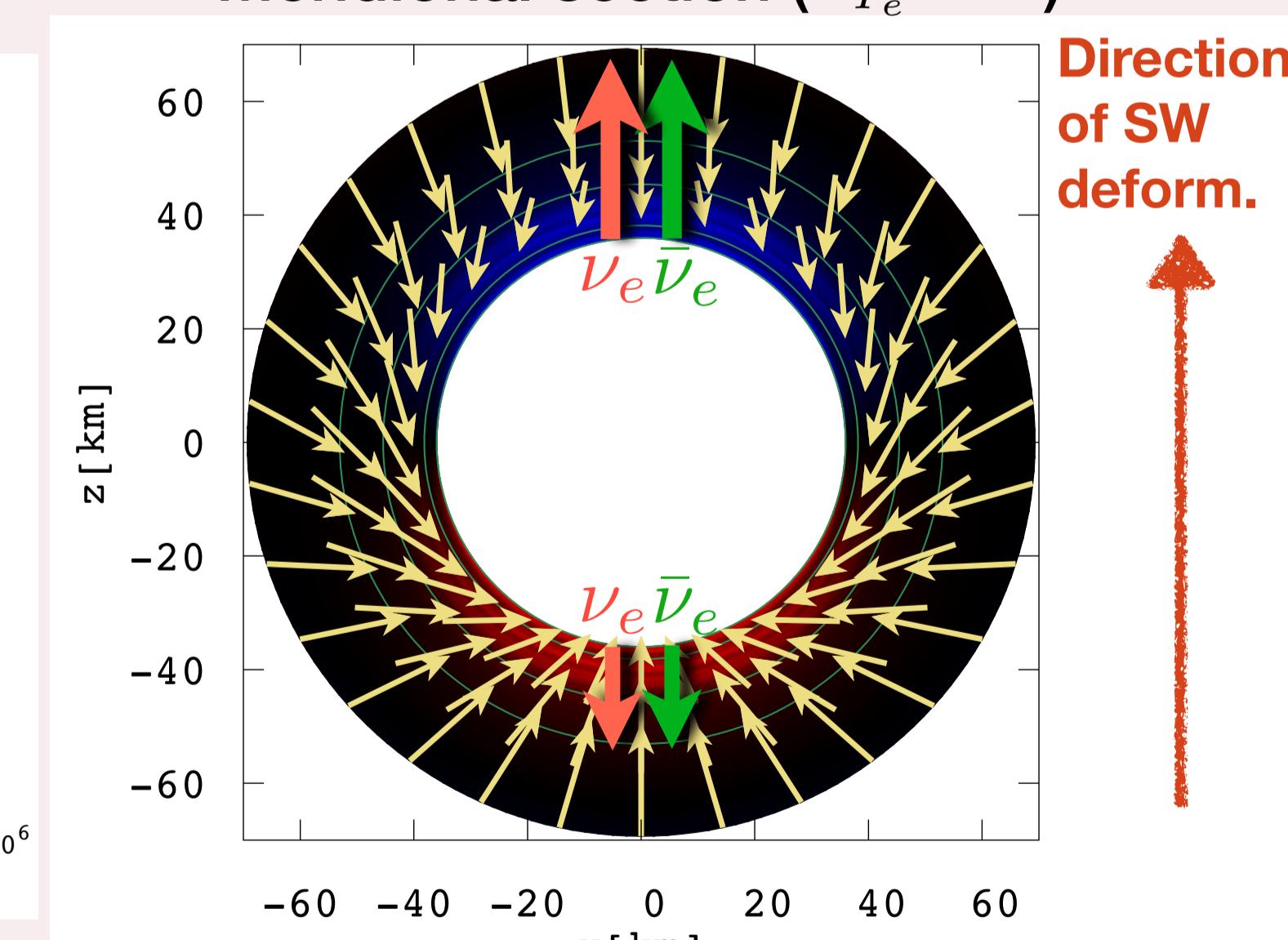
We found self-sustained structure in both cases.

Neutrino luminosity (erg s <sup>-1</sup> )	$c_{Y_e} = 3.5$			$c_{Y_e} = 0$		
	$\delta r_{sh}/r_{sh0}$	$\delta T_\nu/T_{\nu 0}$	$\delta Y_e/Y_{e0}$	$\delta r_{sh}/r_{sh0}$	$\delta T_\nu/T_{\nu 0}$	$\delta Y_e/Y_{e0}$
$2 \times 10^{52}$	$7.66 \times 10^{-3}$	1	-0.650	$2.04 \times 10^{-2}$	1	0
$3 \times 10^{52}$	$1.51 \times 10^{-1}$	1	-0.533	$1.86 \times 10^{-1}$	1	0
$4 \times 10^{52}$	$4.43 \times 10^{-1}$	1	-0.421	$4.81 \times 10^{-1}$	1	0
$5 \times 10^{52}$	1.60	1	-0.219	1.63	1	0

#### Radial distributions of various quantities



#### Meridional section ( $c_{Y_e} = 0$ )



• Steady structure and sign of shock deformation doesn't change in both cases.  
• If LESA exist, shock deformation is suppressed because of stronger cooling.

→ It is natural to consider the sum of neutrino luminosity fluctuation is closely related with shock deformation.

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