

The (missing) link between Type Ia SNe and their SNRs



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RIKEN – iTHEMS
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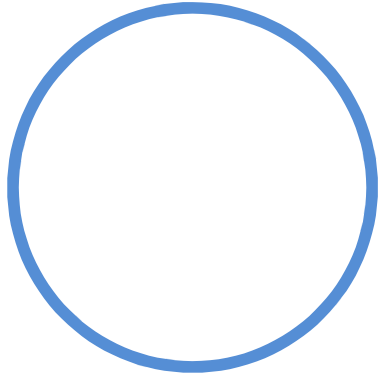
donald.warren@riken.jp

with

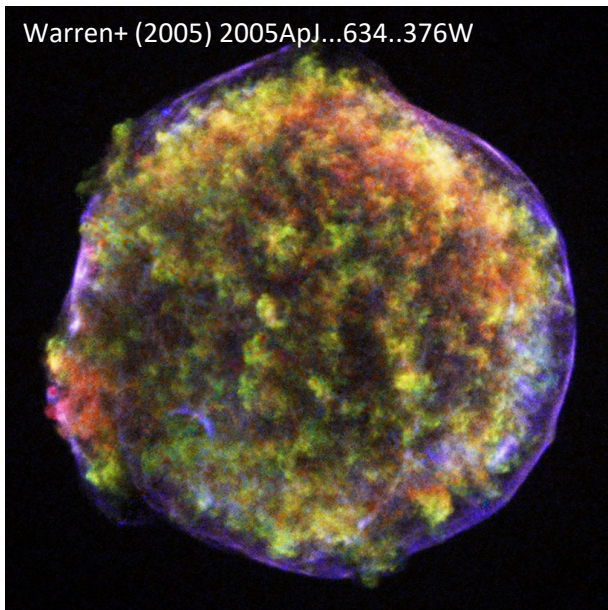
Gilles Ferrand (RIKEN)
Shigehiro Nagataki (RIKEN)
Masaomi Ono (RIKEN)
Fritz Röpke (HITS)
Ivo Seitenzahl (UNSW)

A hopefully non-controversial statement

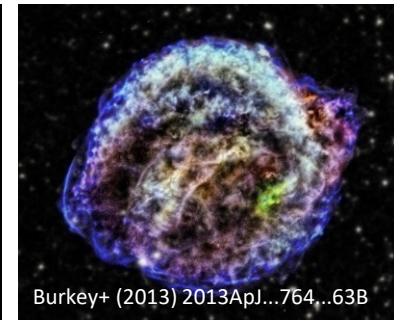
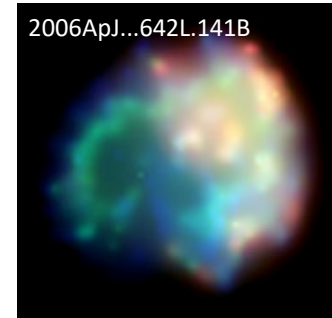
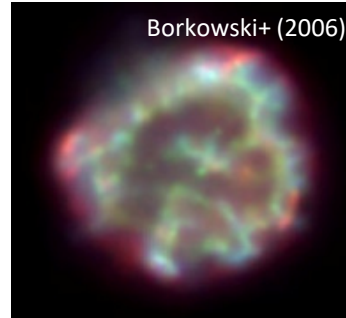
This is a perfect circle



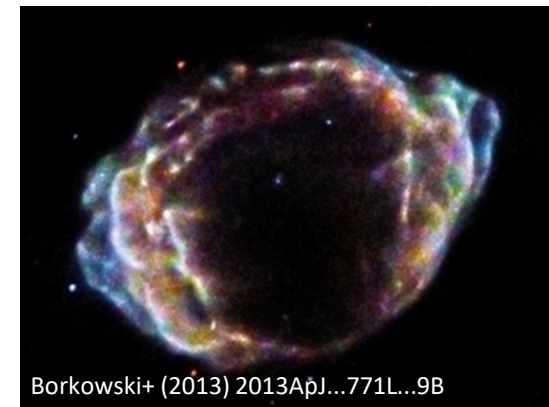
This is not



Neither are these



And especially
not this

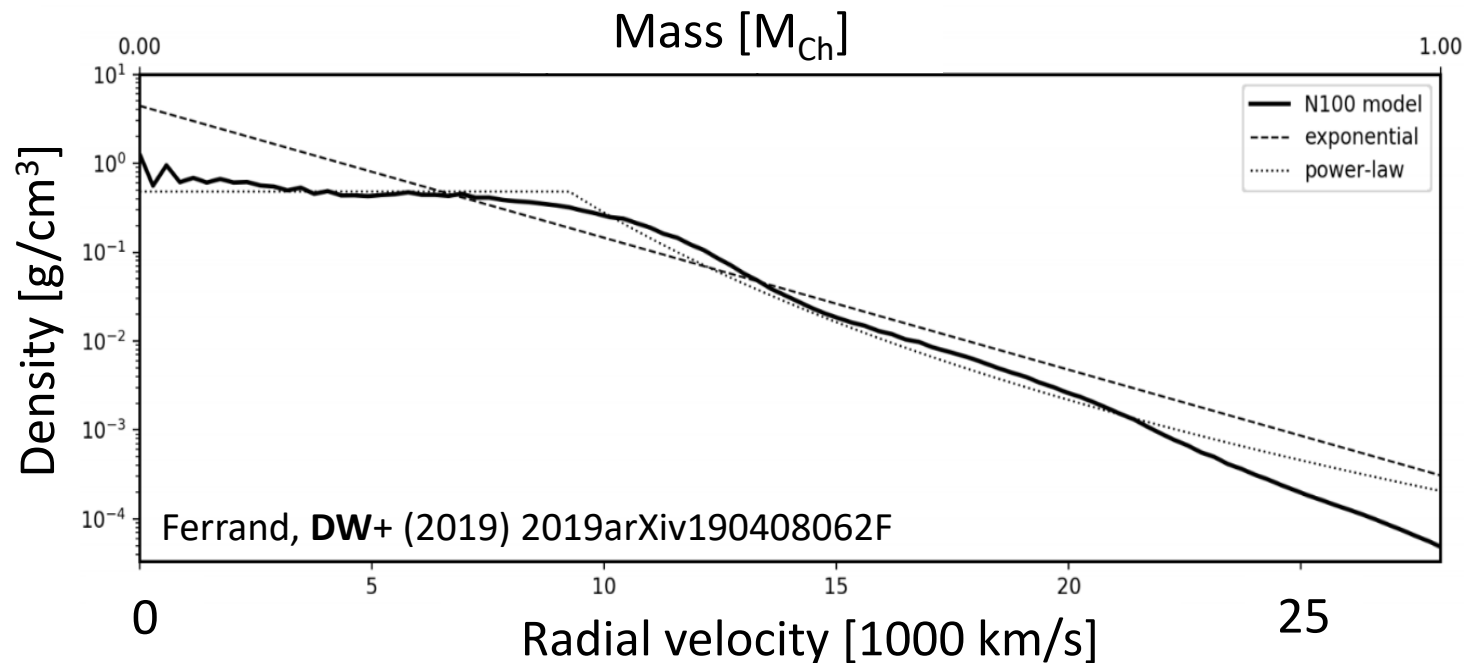


Reasons SNRs might not be spheres

- Inhomogeneous ejecta (Wang & Chevalier 2001; Orlando+ 2012; Warren & Blondin 2013; Williams+ 2017)
- Asymmetric CSM (Lu+ 2011; Williams+ 2011; Liu+ 2012; Fang+ 2018)

Reason ejecta might not be homogeneous

- The supernova
- SNR models all 1-D to start



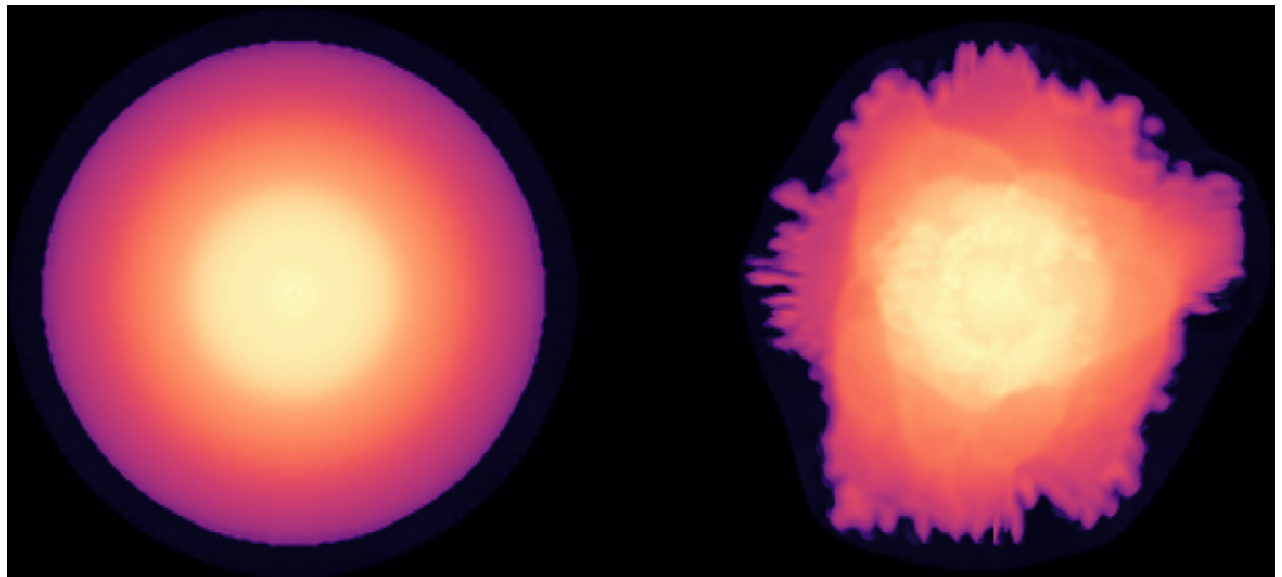
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Reason ejecta might not be homogeneous

- The supernova
- SNR models all 1-D to start
- But supernovae are decidedly not

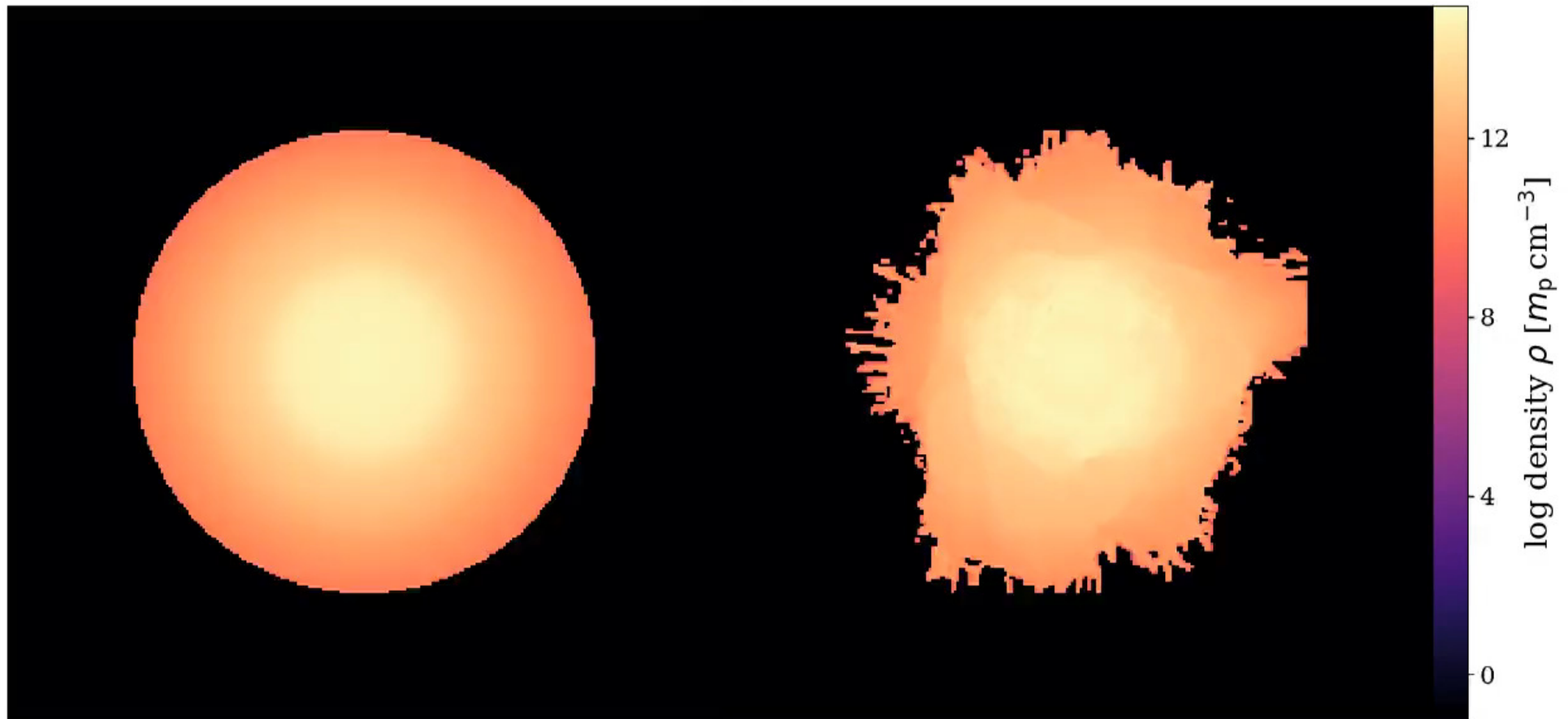
Ferrand, DW+ (2019) 2019arXiv190408062F



Starting an SNR simulation with the SN

- Effects of 1-D initial conditions vs 3-D initial conditions?
- How long does “memory” of SN persist into SNR evolution?

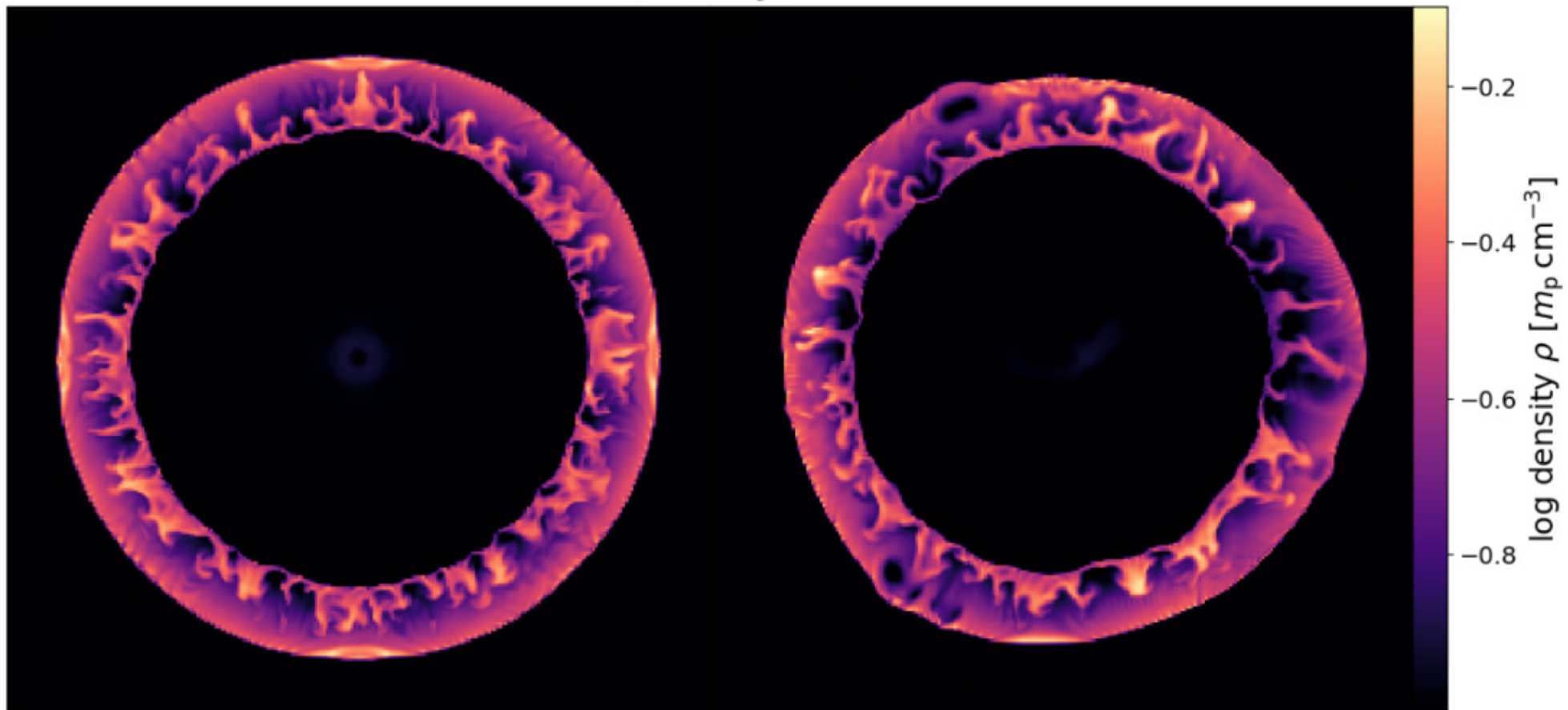
$t = 0 \text{ yr}$



Starting an SNR simulation with the SN

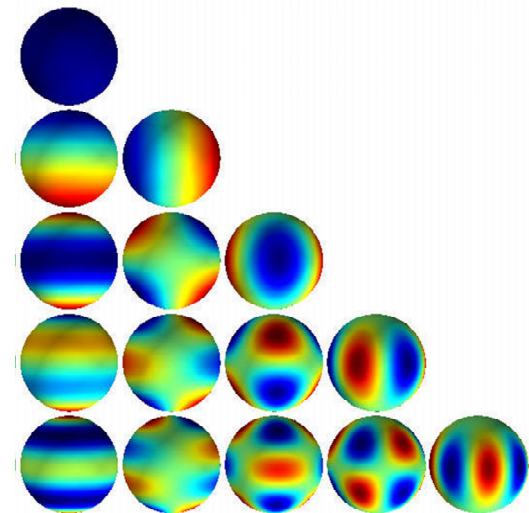
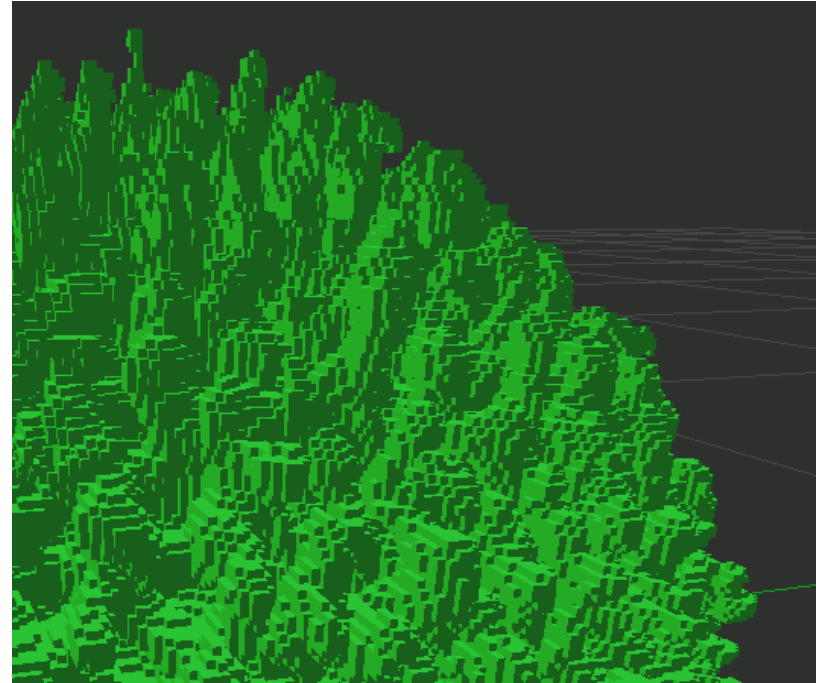
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t = 500 yr

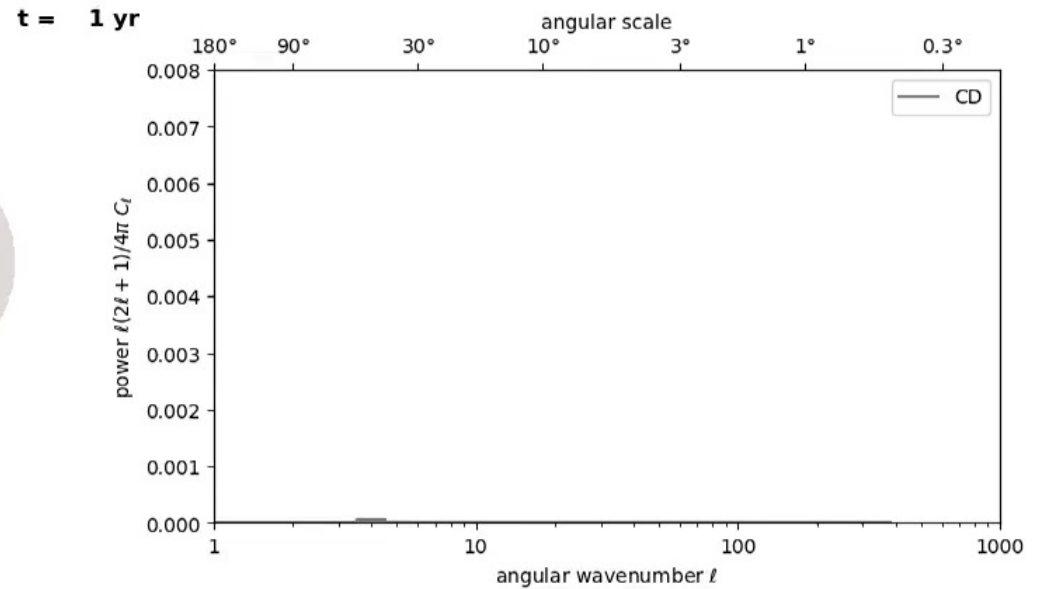
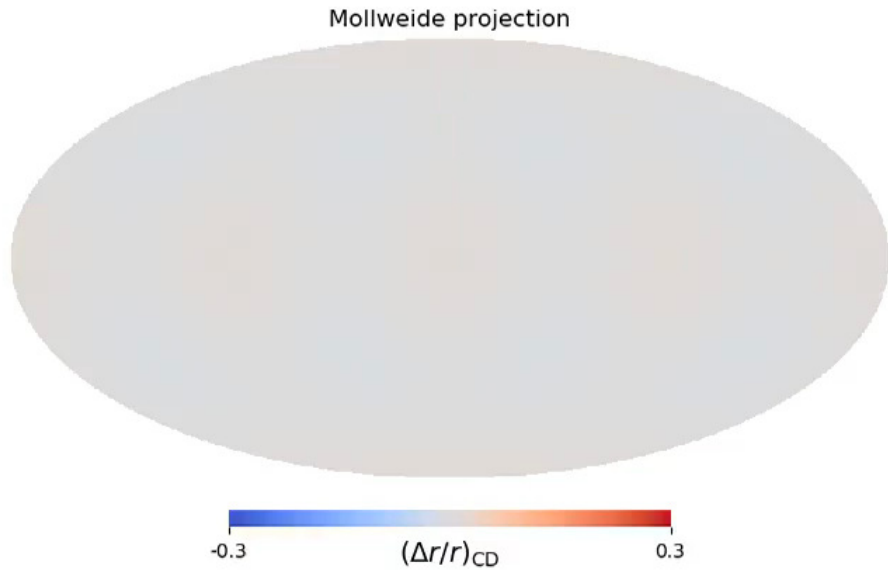
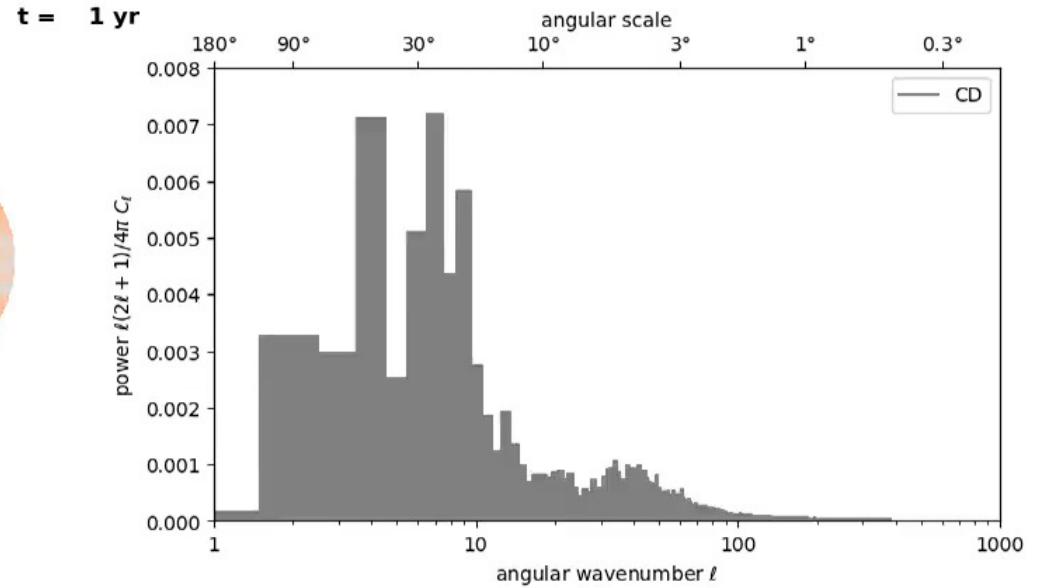
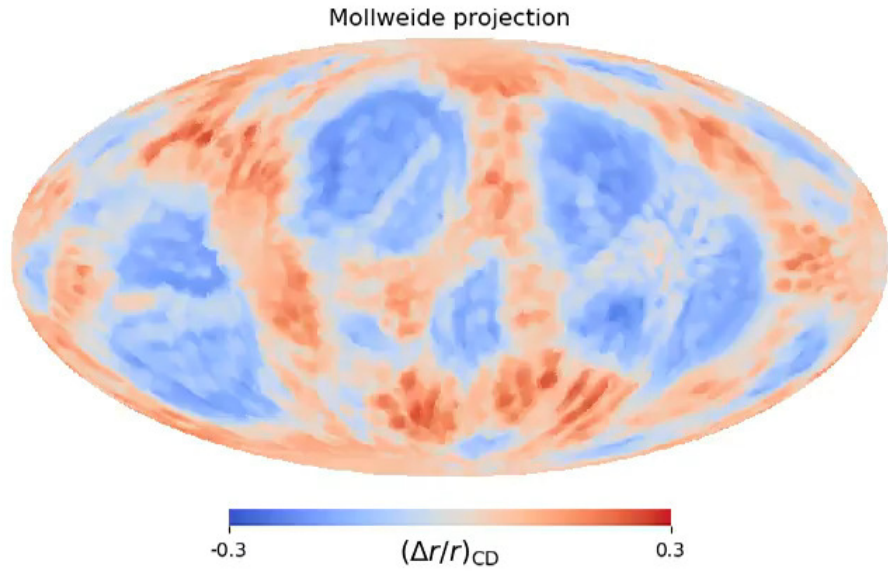


1-D vs 3-D initial conditions

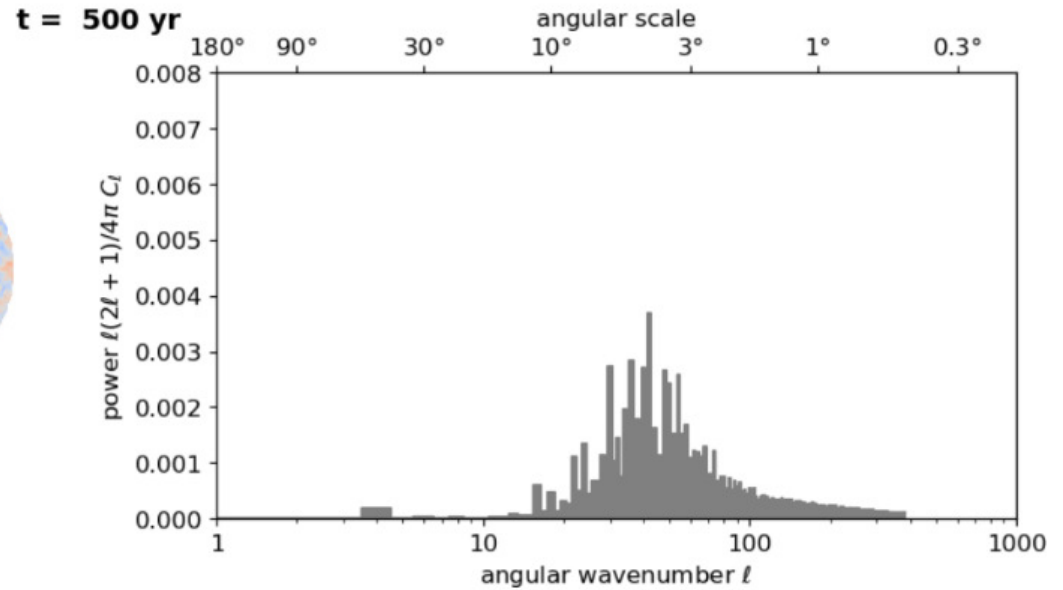
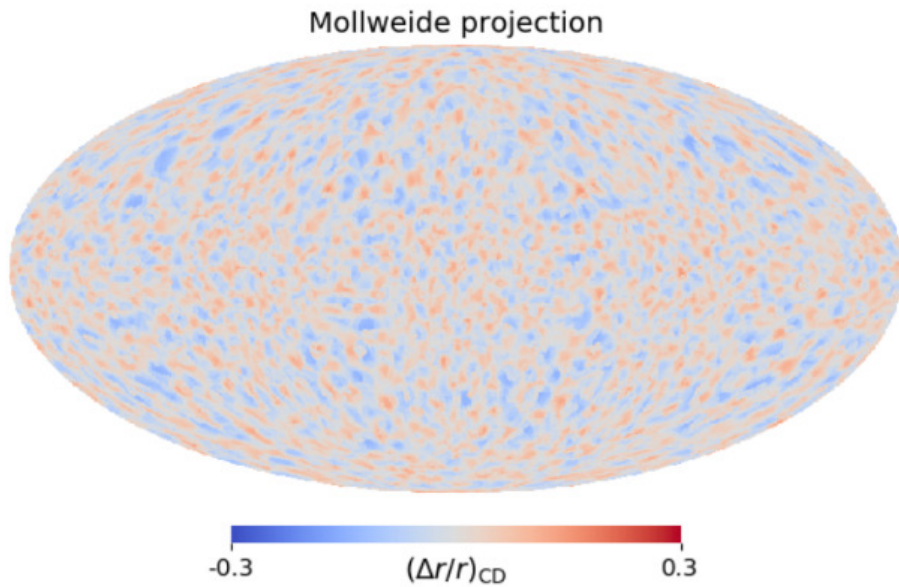
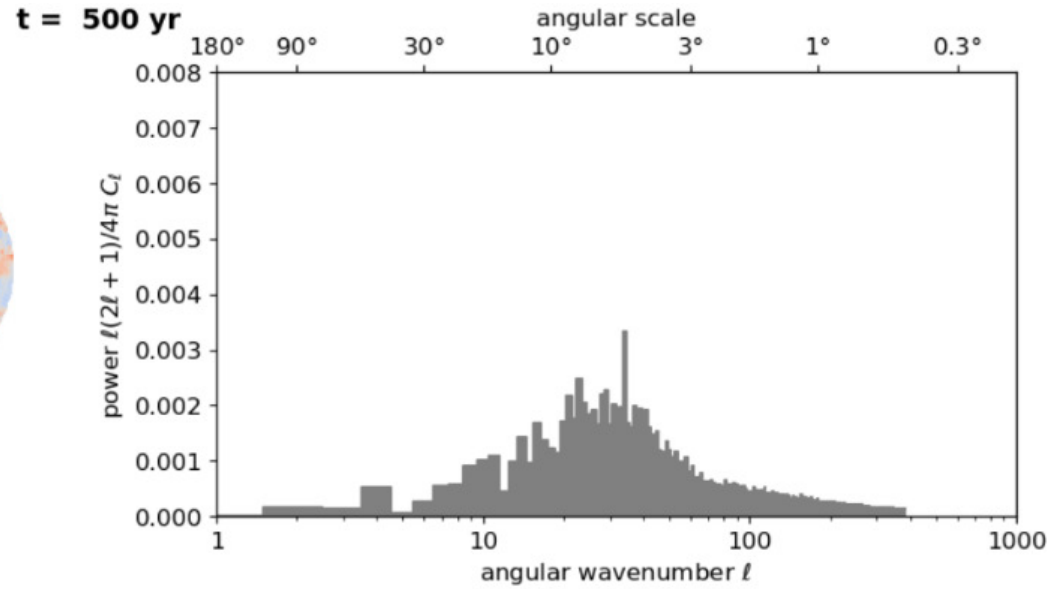
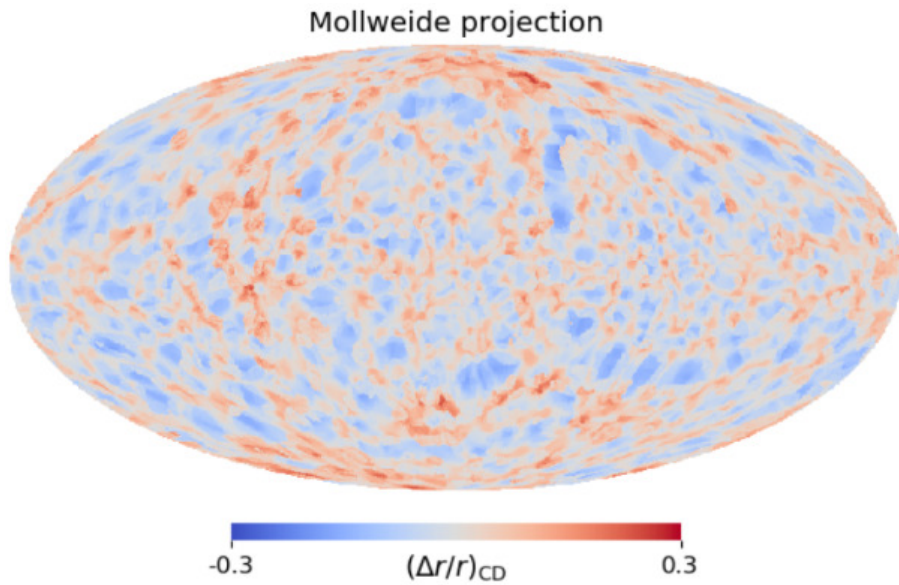
- Find forward shock (FS), contact discontinuity (CD) and reverse shock (RS) as continuous surfaces
- Cast radial rays through each interface to project radius on a sphere (HEALPix for tessellation)
- Plot interfaces in Mollweide projection
- Compute angular power spectrum for interfaces



1-D vs 3-D initial conditions



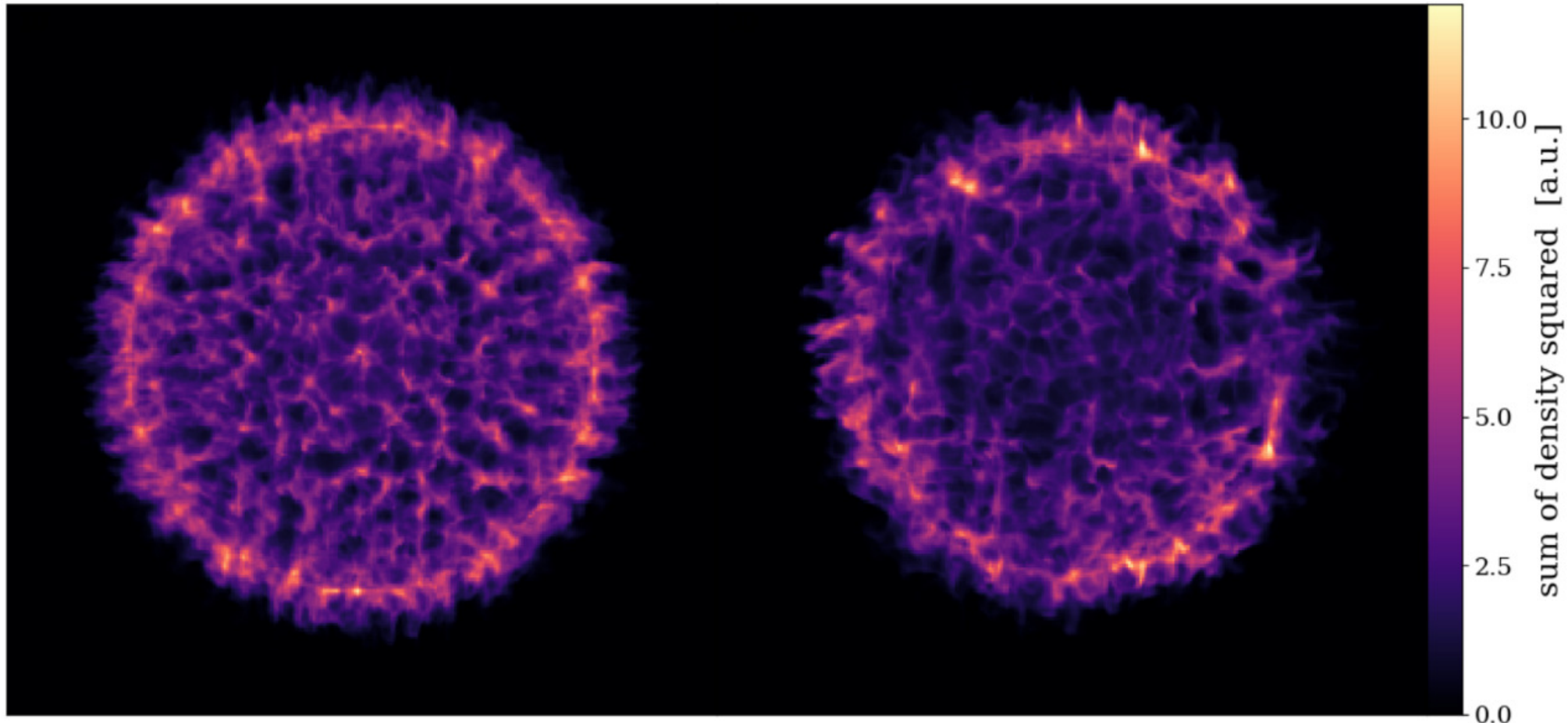
1-D vs 3-D initial conditions



1-D vs 3-D initial conditions

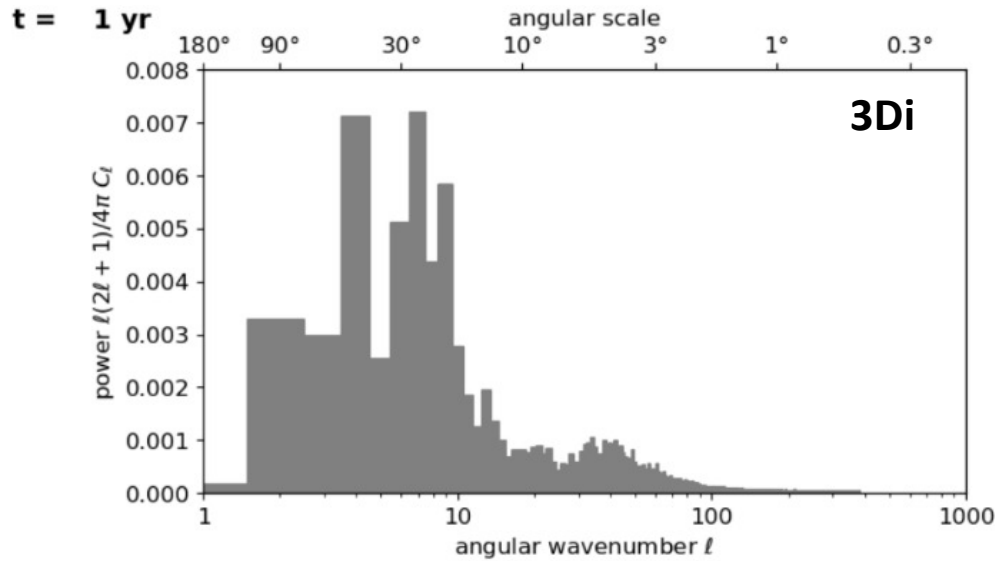
- Power at low wavenumber easily noticeable in projection (c.f. “genus statistic” of Sato+ 2019)

$t = 500 \text{ yr}$

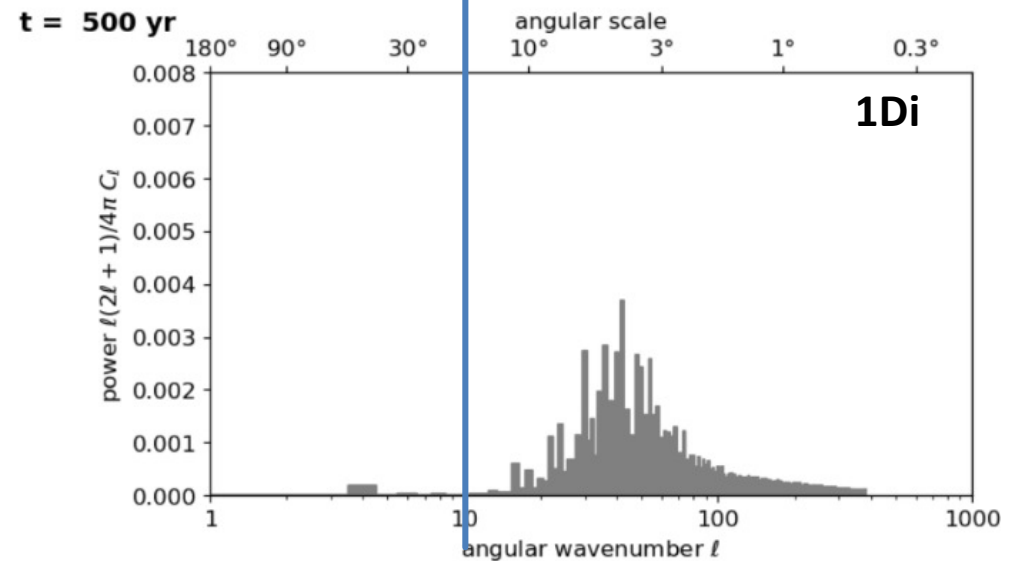
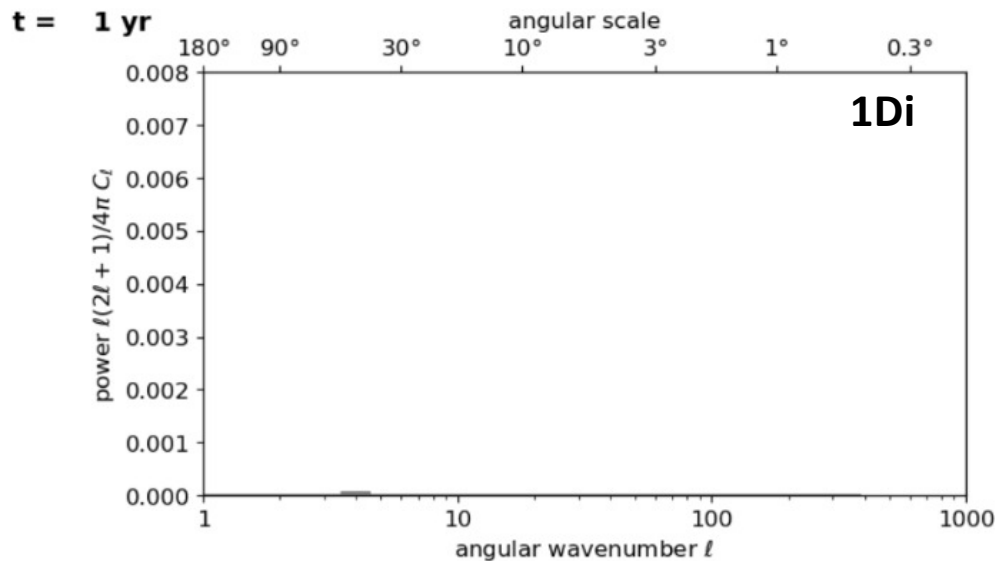
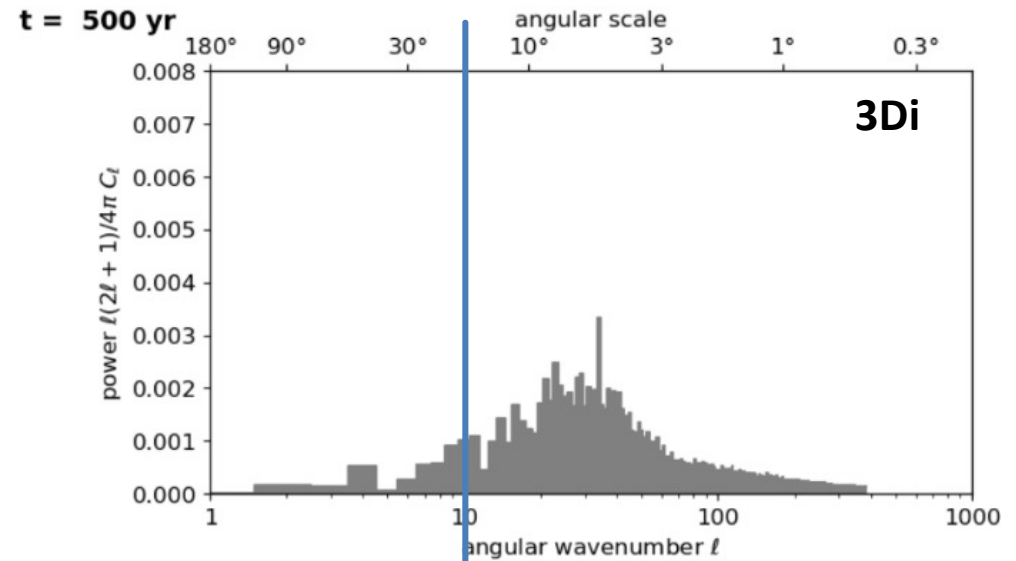


SNR memory of the SN

CD power at 1 yr

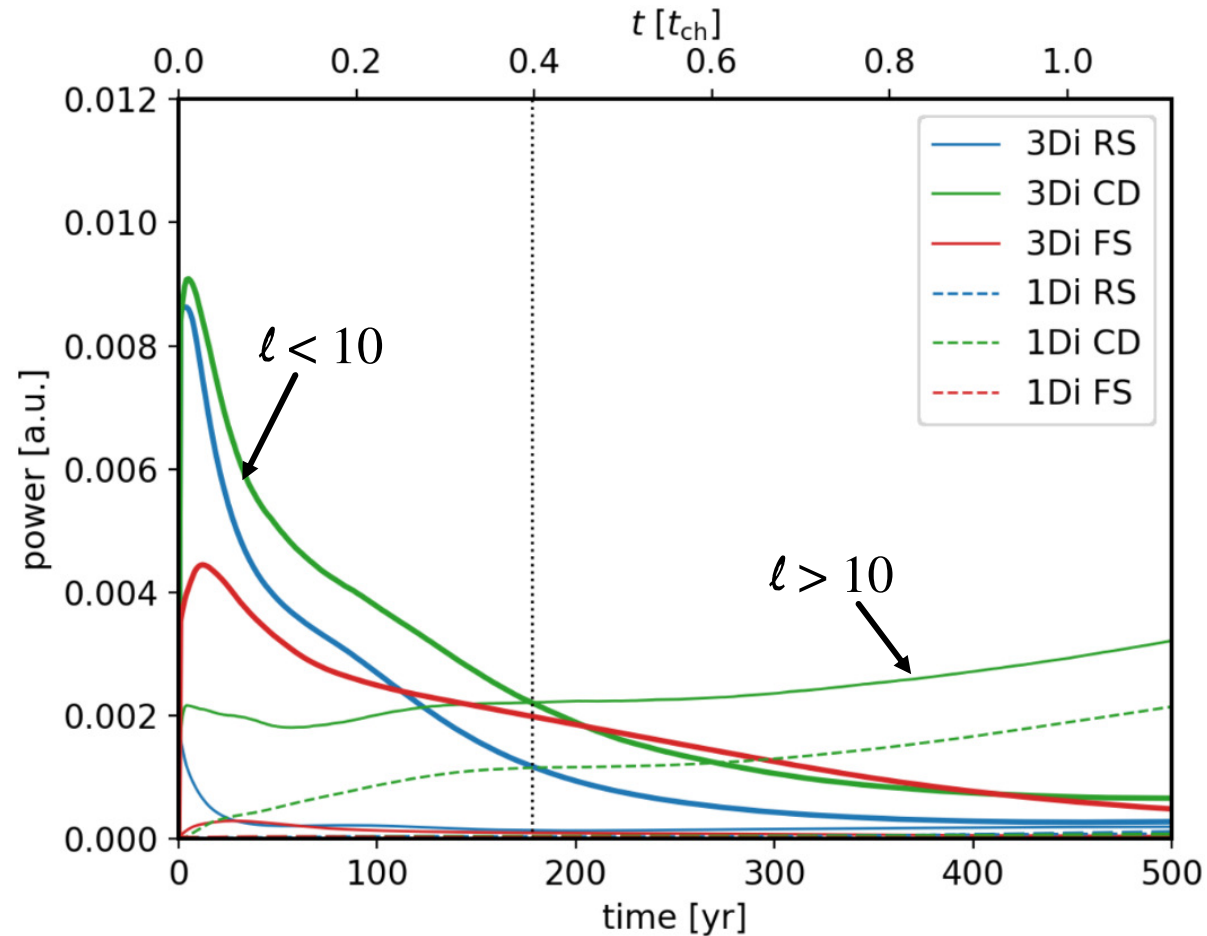
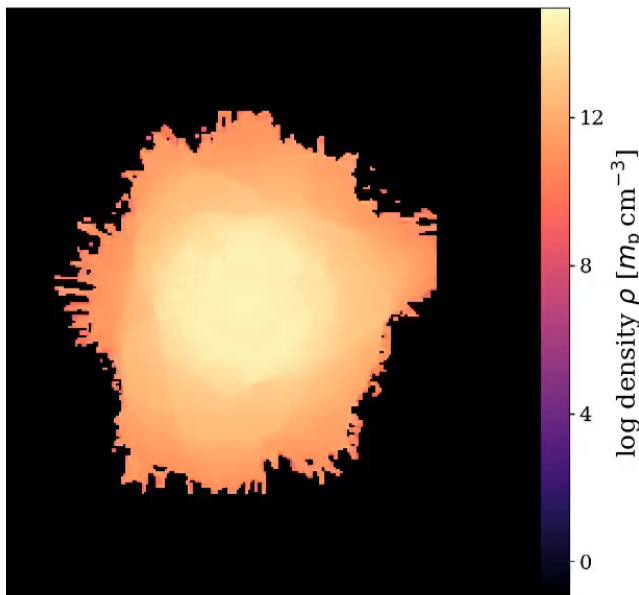


CD power at 500 yrs



SNR memory of the SN

- Remnants circularize & lose power at low ℓ
- Development of RT instability creates power at large ℓ



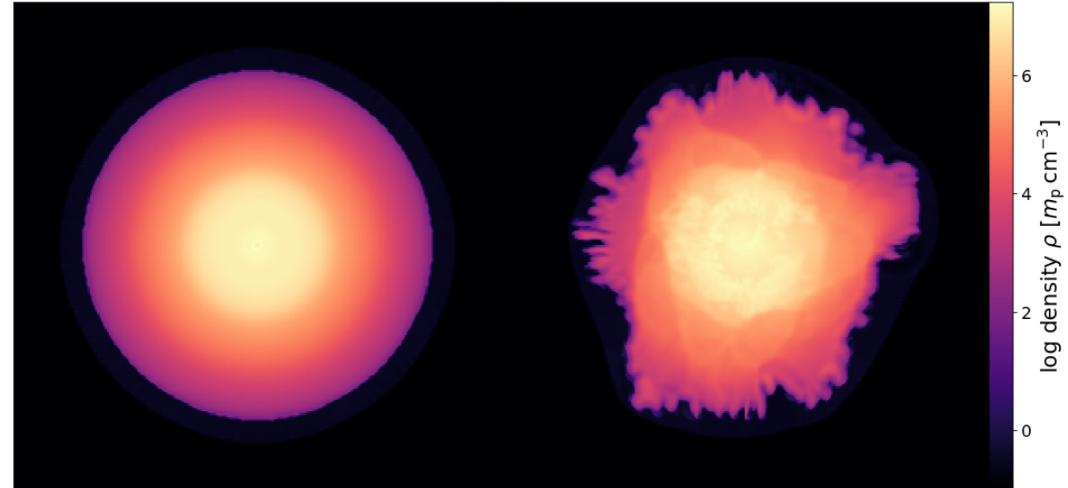
- By ≈ 175 years, SNR power dominates SN power \rightarrow SN transitions to SNR

Summary

Ferrand, DW+ (2019) 2019arXiv190408062F

t = 1 yr

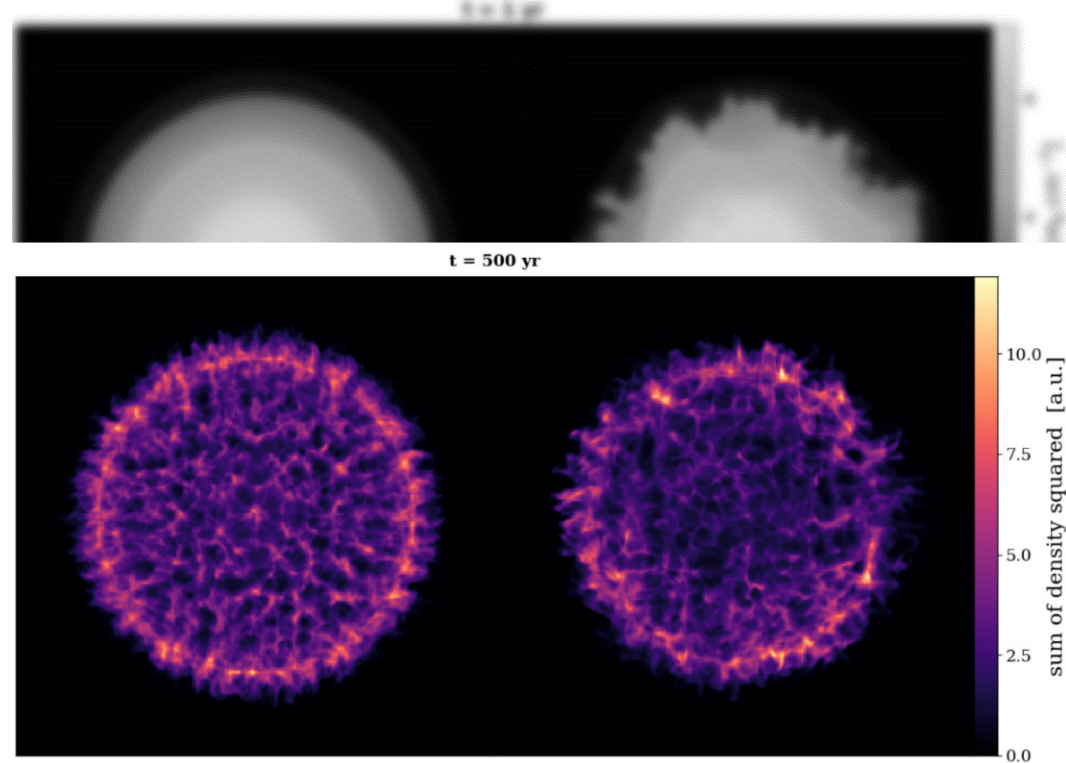
- Simulations of SNe are not spherically symmetric



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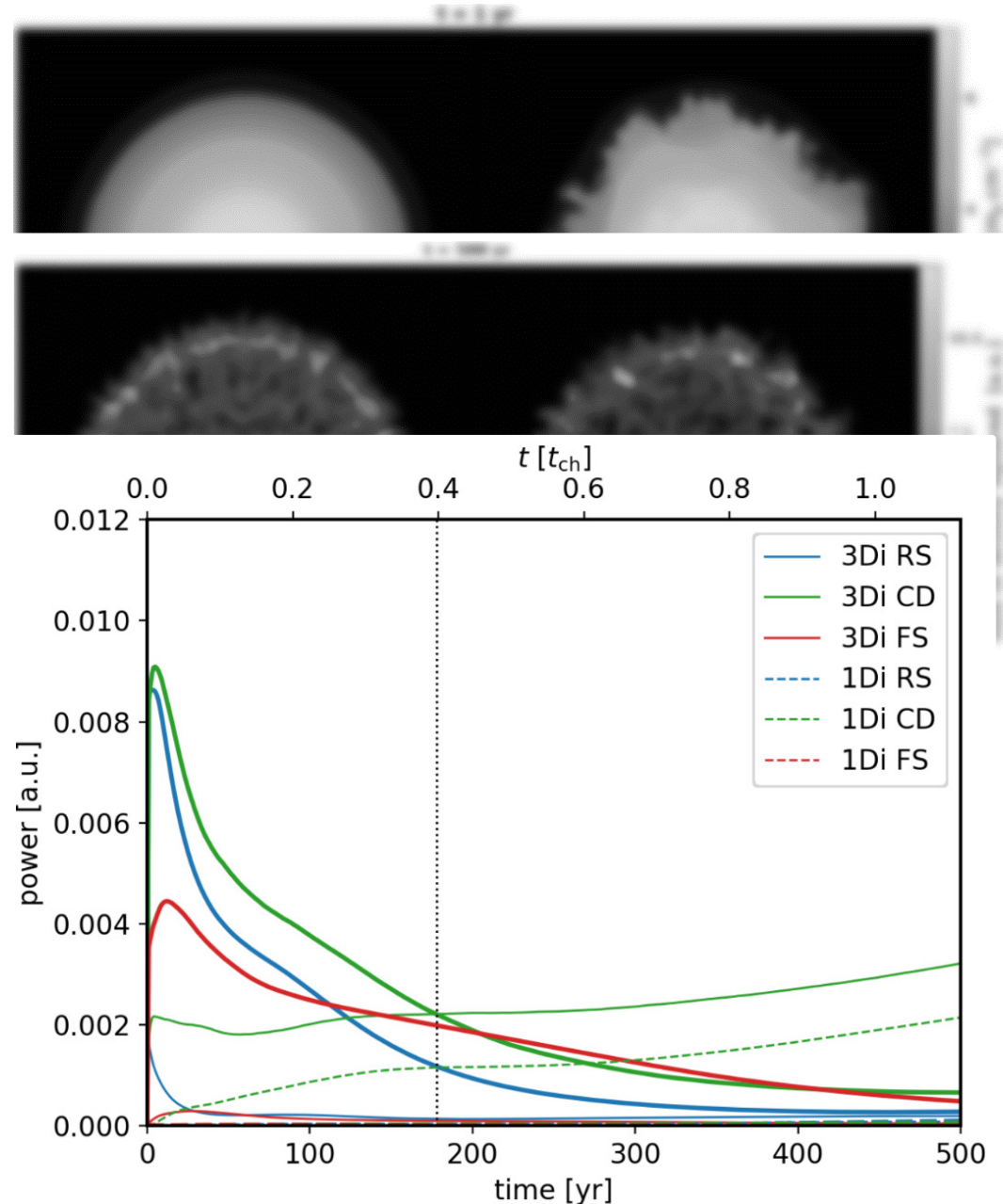
- Simulations of SNe are not spherically symmetric
- Using SN as initial conditions causes detectable differences to SNR → possible to work backwards and infer SN mechanism?



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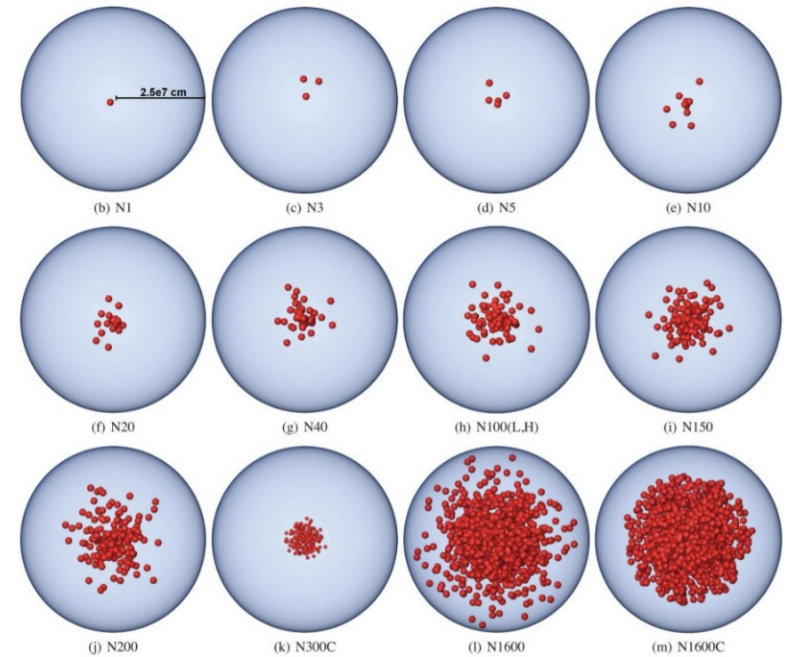
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- Simulations of SNe are not spherically symmetric
- Using SN as initial conditions causes detectable differences to SNR \rightarrow possible to work backwards and infer SN mechanism?
- Imprint of SN in SNR can be detected long after ejecta morphology dominated by interaction with ISM



Perspectives

- Compute thermal emission for accurate predictions of observations
- Analyze our projected images quantitatively (e.g. genus statistic of Sato+ 2019)
- Run more models (single degenerate? double degenerate?)
- Eventually work to diagnose explosion model based on appearance of SNR
- We will evolve your favorite SN model. Just ask!



One last plug for VR

